

Polarization of HL Tau

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supervised by Aso Yusuke

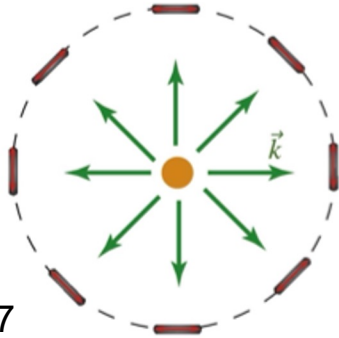
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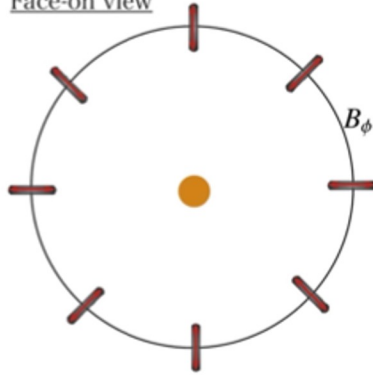
Introduction

Polarization Mechanisms in Young Stellar Objects

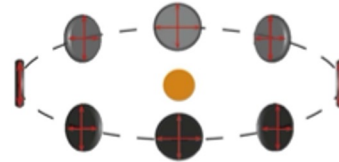
Face-on view



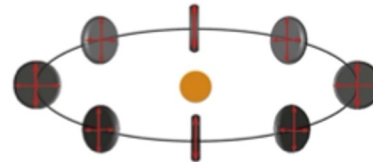
Face-on view



Inclined view

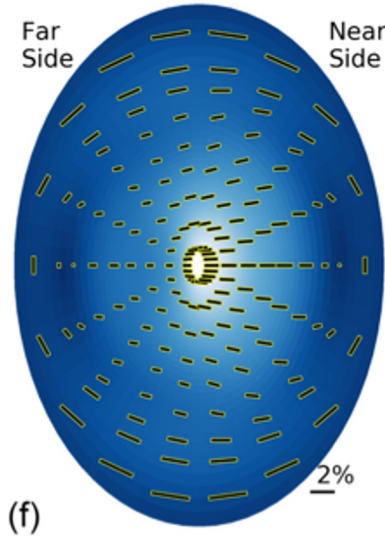


Inclined view



Grain alignment
along radiation

Grain alignment
along
magnetic fields



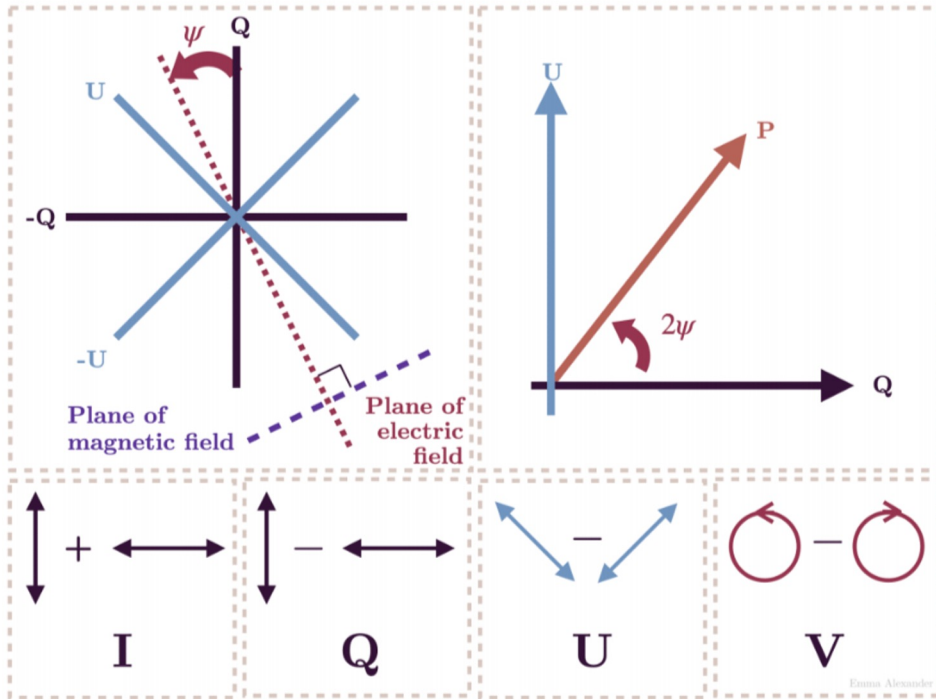
(f)

Yang+'17

Self-scattering
(90°-scattering is strongest)

Tazaki+'17

Observable polarization signals



Spectral index α , β

Spectral index is related to dust properties, especially **size of grain**.
Larger grain size gives smaller spectral index. (Draine, 2006)

$$F_\nu \approx F_{\nu_0} \left(\frac{\nu}{\nu_0} \right)^\alpha$$

$$F_\nu \approx \kappa_\nu B_\nu(T_d) \frac{M_T}{D^2} \quad (\text{optically thin})$$

$$\approx \kappa_{\nu_0} \left(\frac{\nu}{\nu_0} \right)^\beta \frac{2kT_d}{c^2} \nu^2 \frac{M_T}{D^2}$$

therefore $\alpha \approx \beta + 2$.

Kwon+'09

Large grains: Flux follows Rayleigh-Jeans law ($\propto \nu^2$), $\alpha \sim 2$, $\beta \sim 0$

Small grains: Flux follows Rayleigh scattering ($\propto \nu^4$), $\alpha \sim 4$, $\beta \sim 2$

Target HL Tau

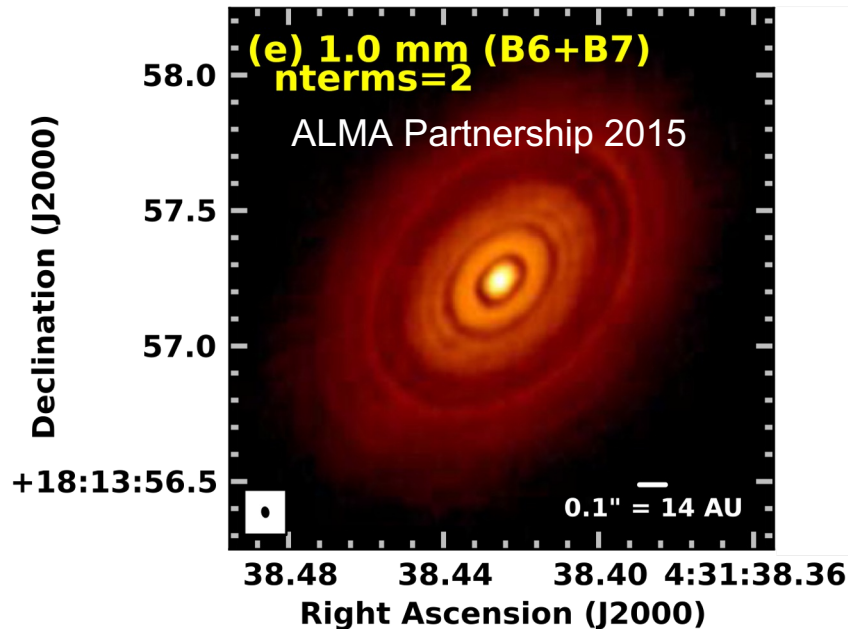
HL Tauri is a Class I/II young stellar object in the constellation Taurus (황소자리).

Distance: 140pc

RA: 04h 31m 38.43s

Dec: 18d 13m 57.12s

Very well studied in multiple-wavelength polarized continuum emission using ALMA.



Data Reduction

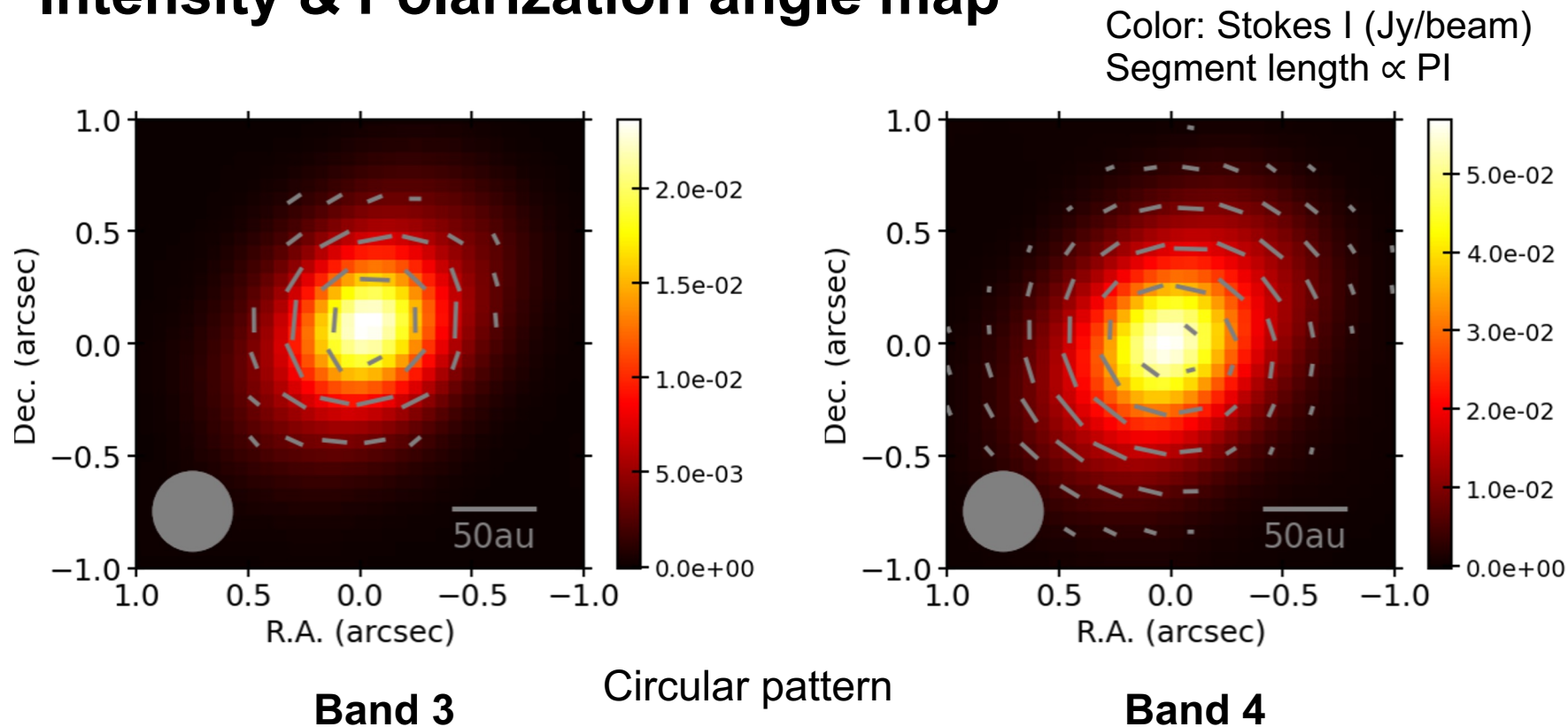
Six measurement sets

	Frequency (GHz)	Robust	Beam Size† (" × ")	UV distance (kλ)	3σ noise level (mJy/beam)
Band 3	97.5	-1.0	0.356×0.200	3.6 - 780	0.10
Band 4	145.0	0.5	0.312×0.260	15.0 - 1,400	0.50
Band 5	203.0	1.5	0.320×0.279	8.5 - 1,920	0.75
Band 6	233.0	0.5	0.321×0.232	10.0 - 1,800	3.5
Band 7a	343.5	-1.0	0.343×0.259	14.2 - 770	2.3
Band 7b	343.5	0.5	0.328×0.262	14.0 - 2,100	3.0

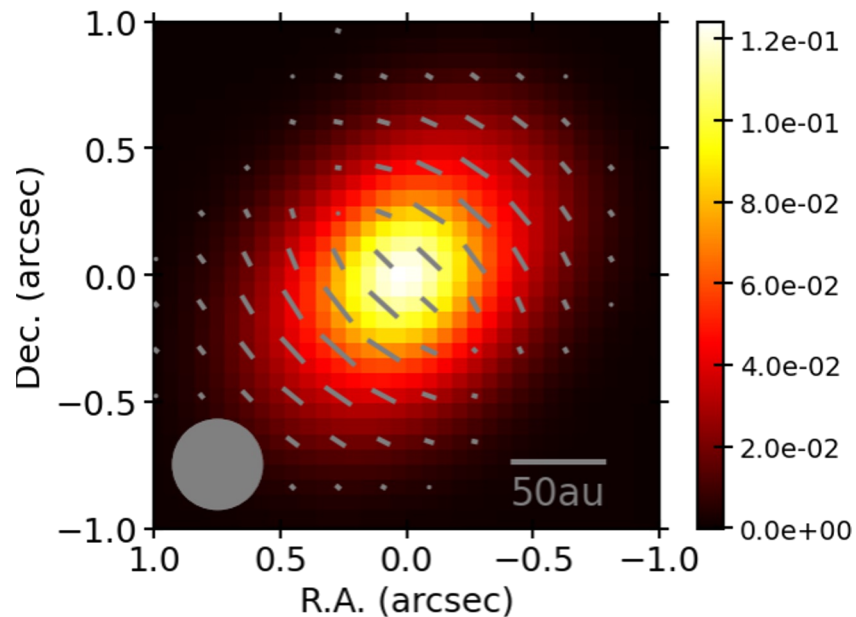
†This beam size is obtained with the common uv-distance range (**15-770 kλ**). We smoothed the images to change all the beam sizes to **0.36"x0.36"**.

Results

Intensity & Polarization angle map



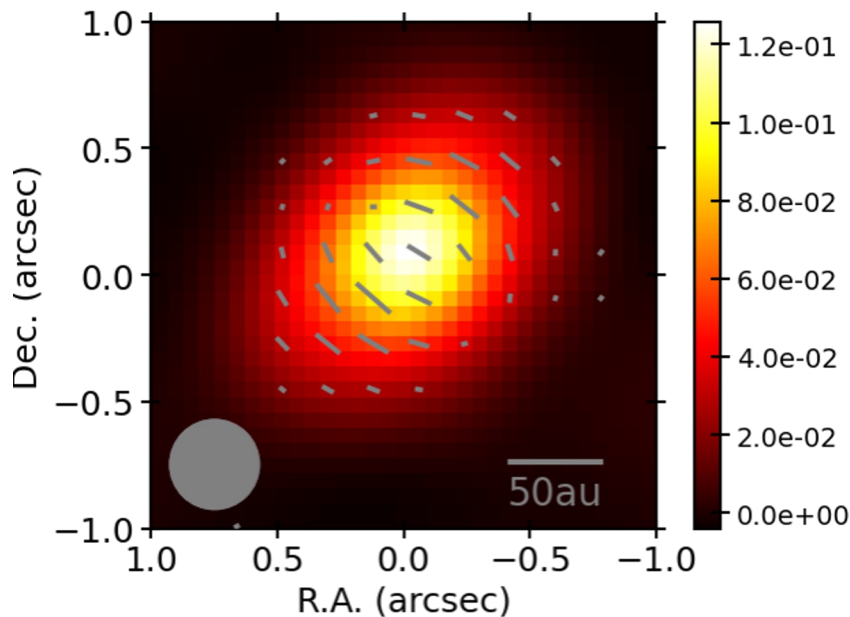
Intensity & Polarization angle map



Band 5

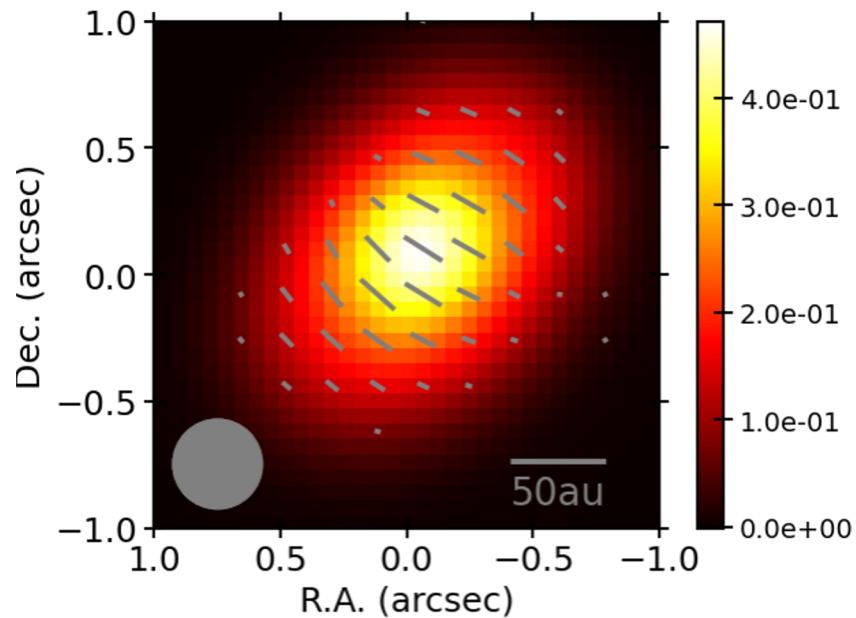
NE-SW (minor axis)

Color: Stokes I (Jy/beam)
Segment length \propto PI



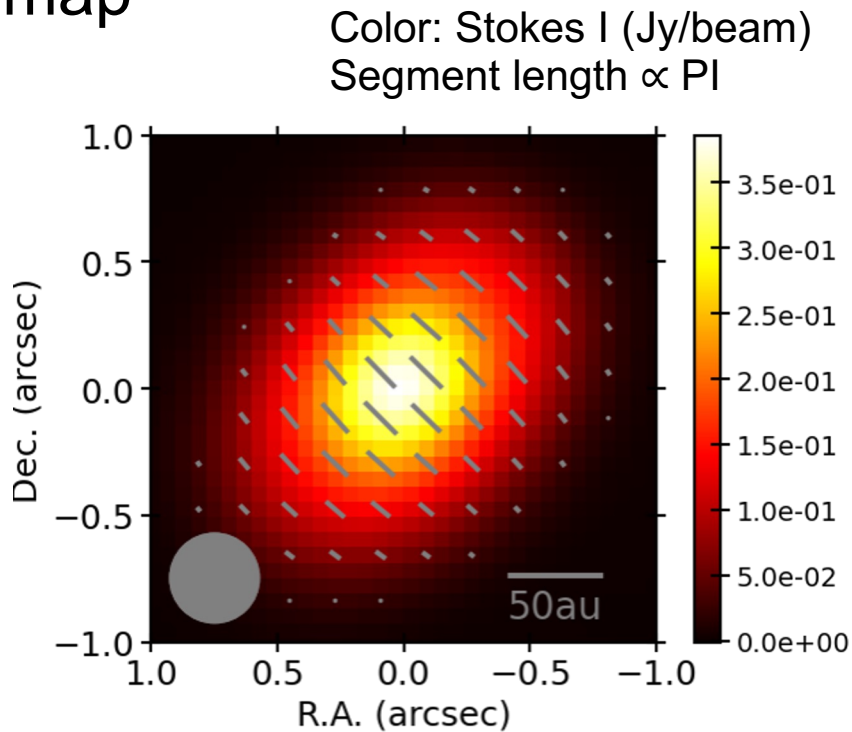
Band 6

Intensity & Polarization angle map



Band 7a

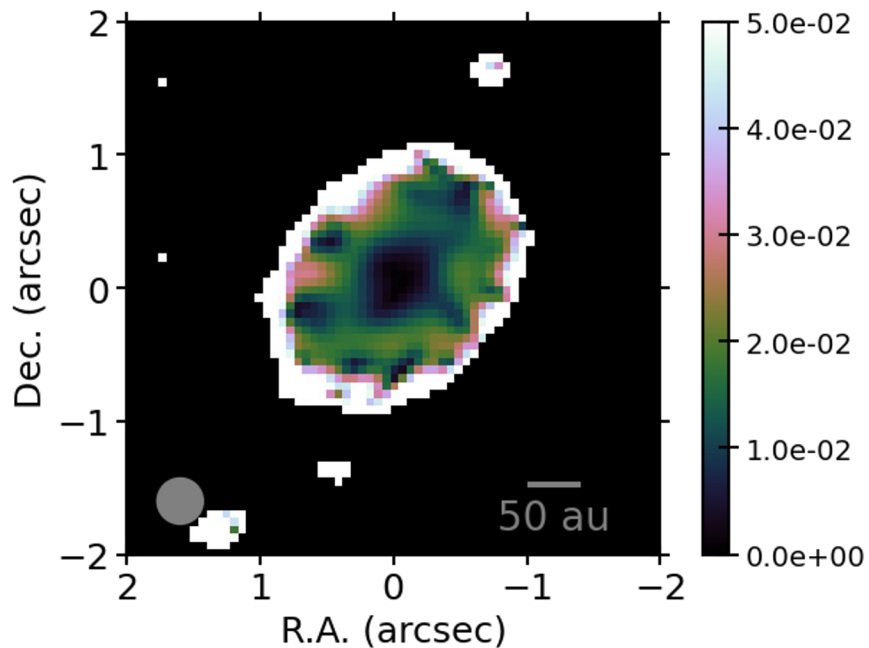
NE-SW (minor axis)



Band 7b

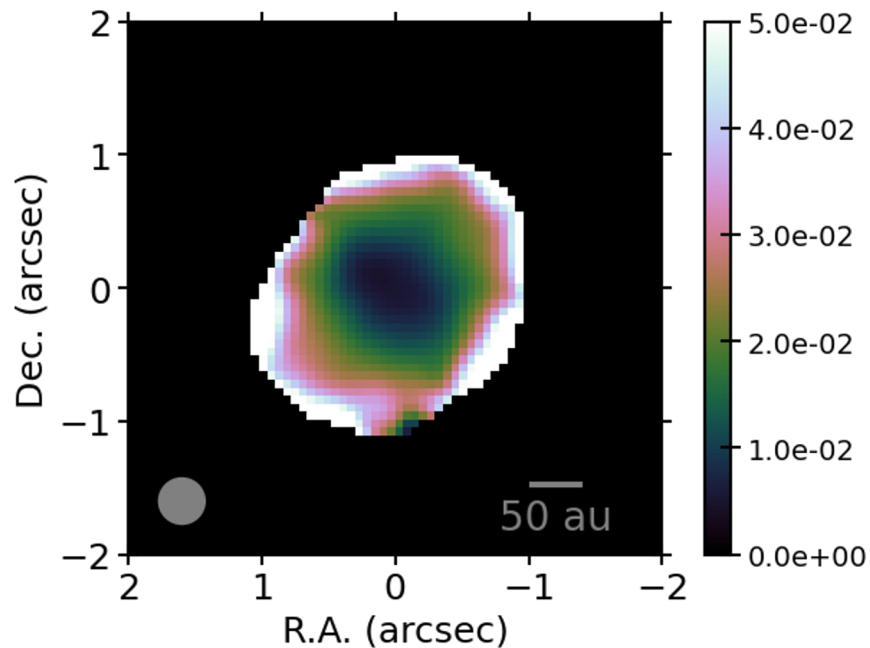
Polarization Fraction

$$p = \frac{PI}{I} = \frac{\sqrt{Q^2 + U^2 + V^2}}{I}$$



Band 3

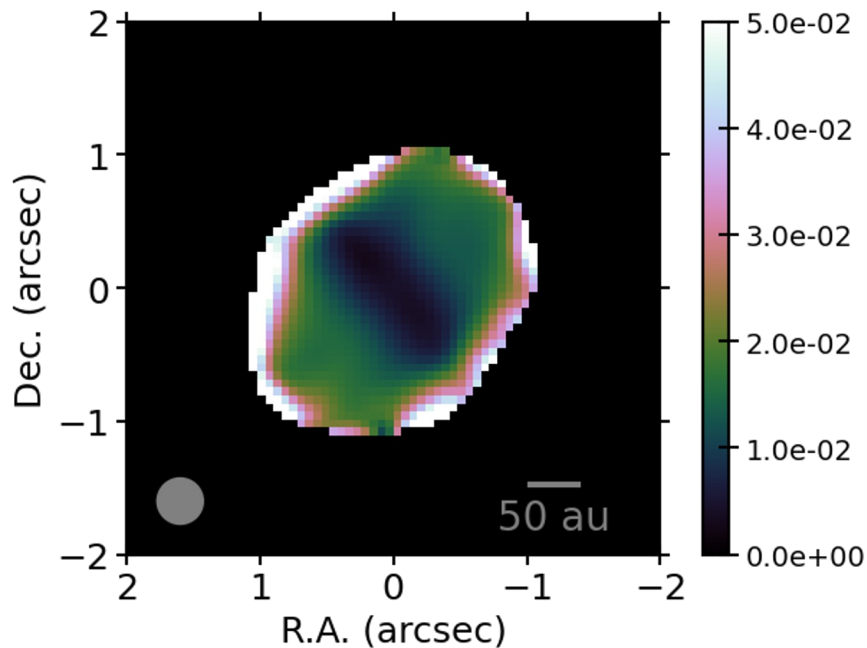
Ring-like distribution



Band 4

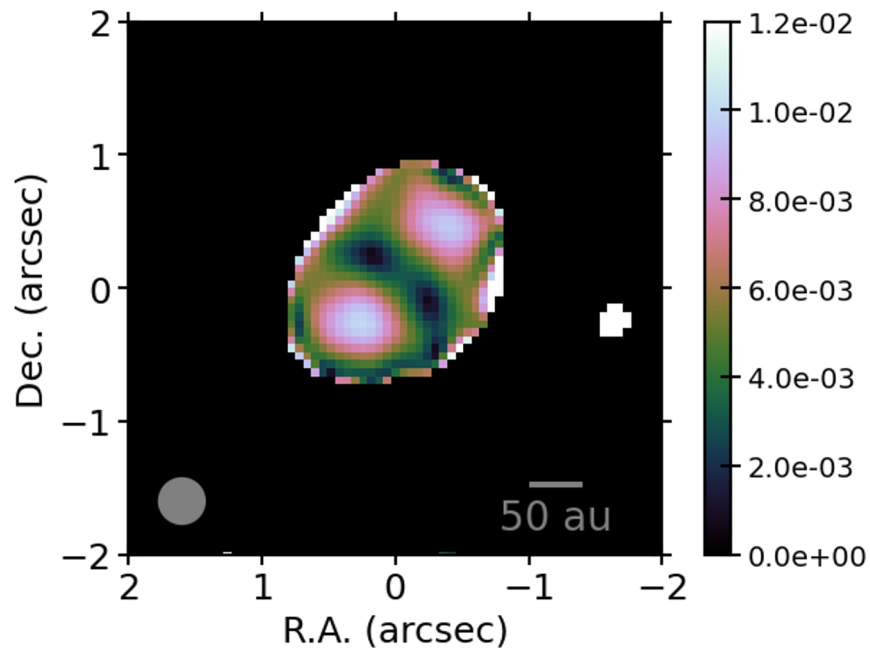
Polarization Fraction

$$p = \frac{PI}{I} = \frac{\sqrt{Q^2 + U^2 + V^2}}{I}$$



Band 5

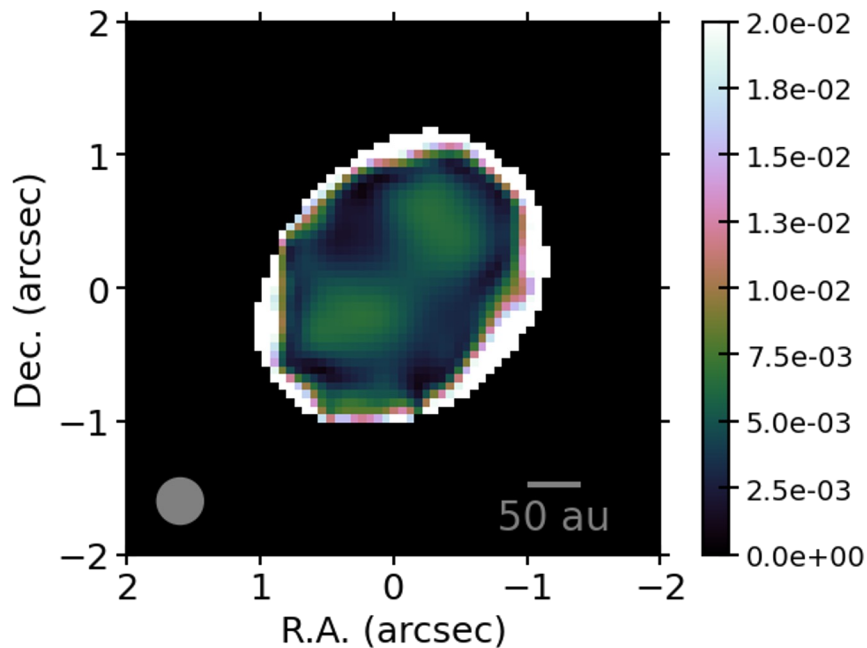
Double-peak distribution



Band 6

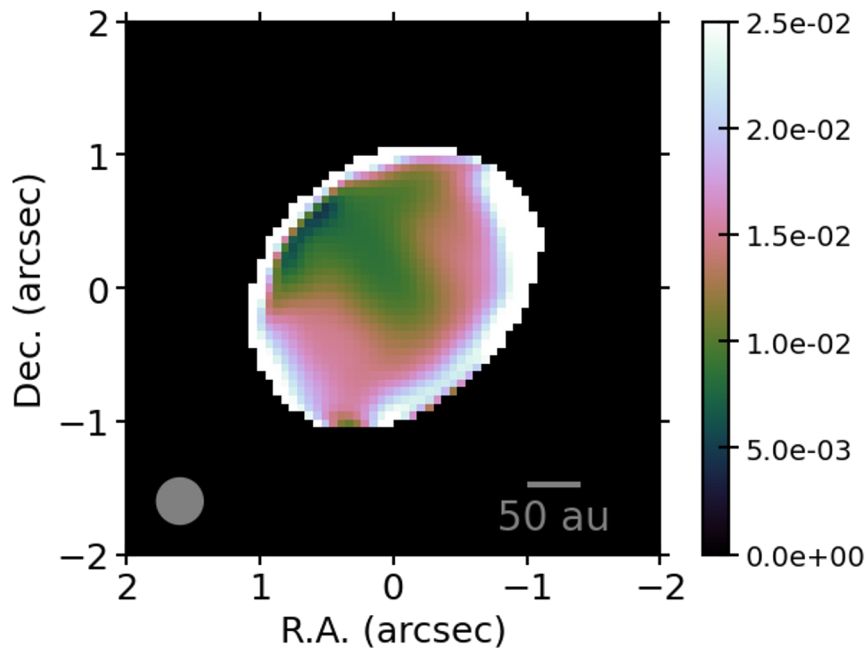
Polarization Fraction

$$p = \frac{PI}{I} = \frac{\sqrt{Q^2 + U^2 + V^2}}{I}$$



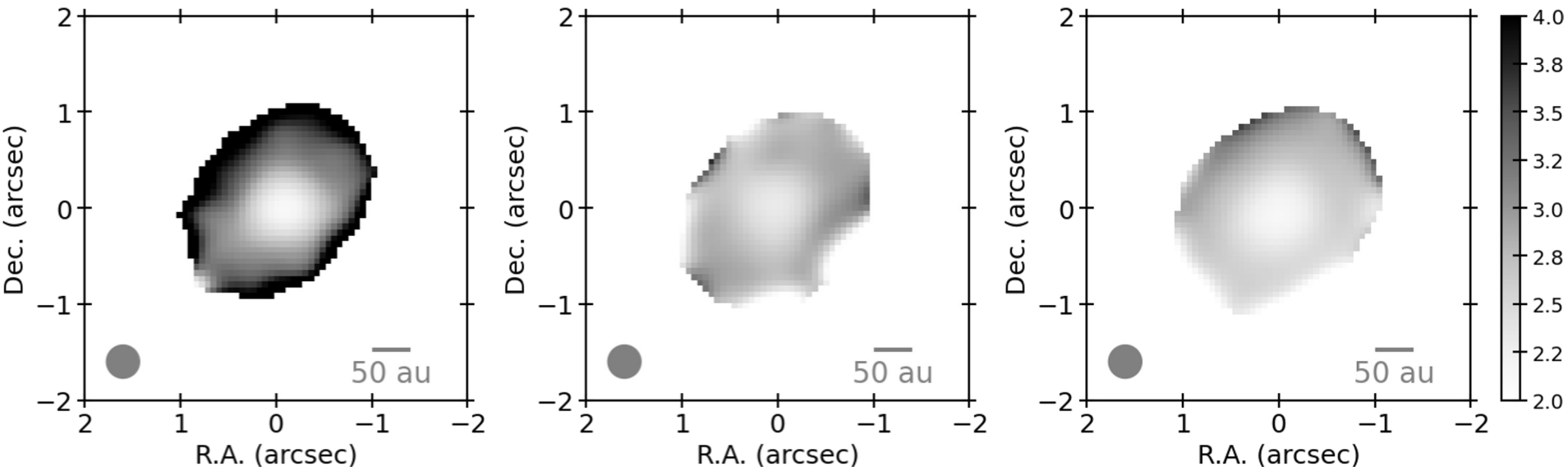
Band 7a

Double-peak distribution



Band 7b

Spectral index α of Band 3-4, Band 4-5, Band 5-7a



Band 3-4

Band 4-5

Band 5-7b

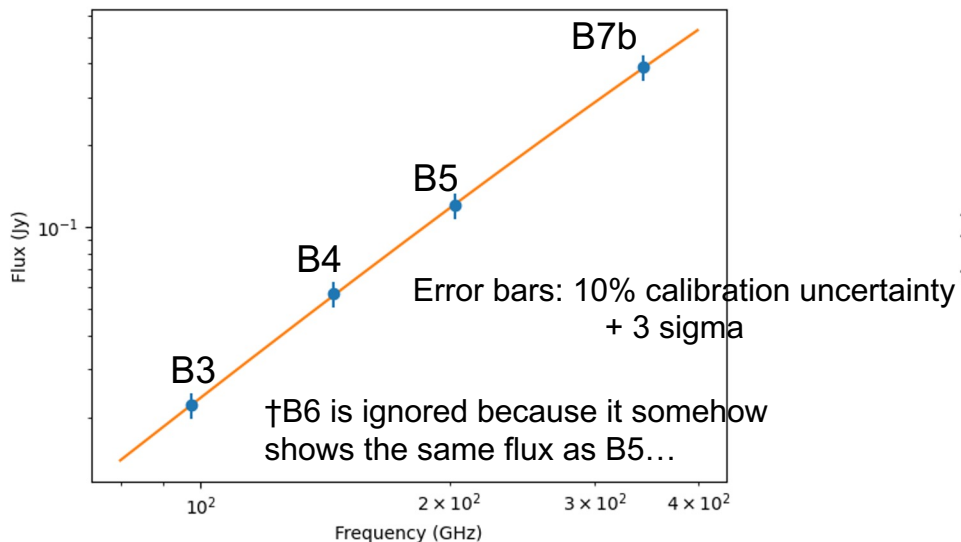
$\alpha \sim 2$ at the center, while $\alpha \sim 3$ in the outer region

Discussion

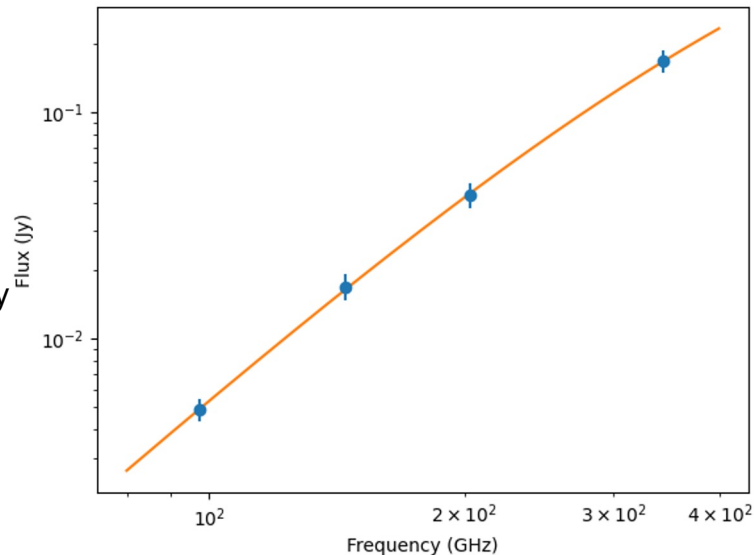
Grain size from SED

$$F_\nu = (B_\nu(T) - B_\nu(T_{bg}))(1 - e^{-\tau_0(\nu/200 \text{ GHz})^\beta})\Omega$$

Central region



Outer region



T	τ_0	β
$55 \pm 16 \text{ K}$	0.8 ± 0.4	0.5 ± 0.1

T	τ_0	β
$25 \pm 2.6 \text{ K}$	0.7 ± 0.1	1.3 ± 0.1

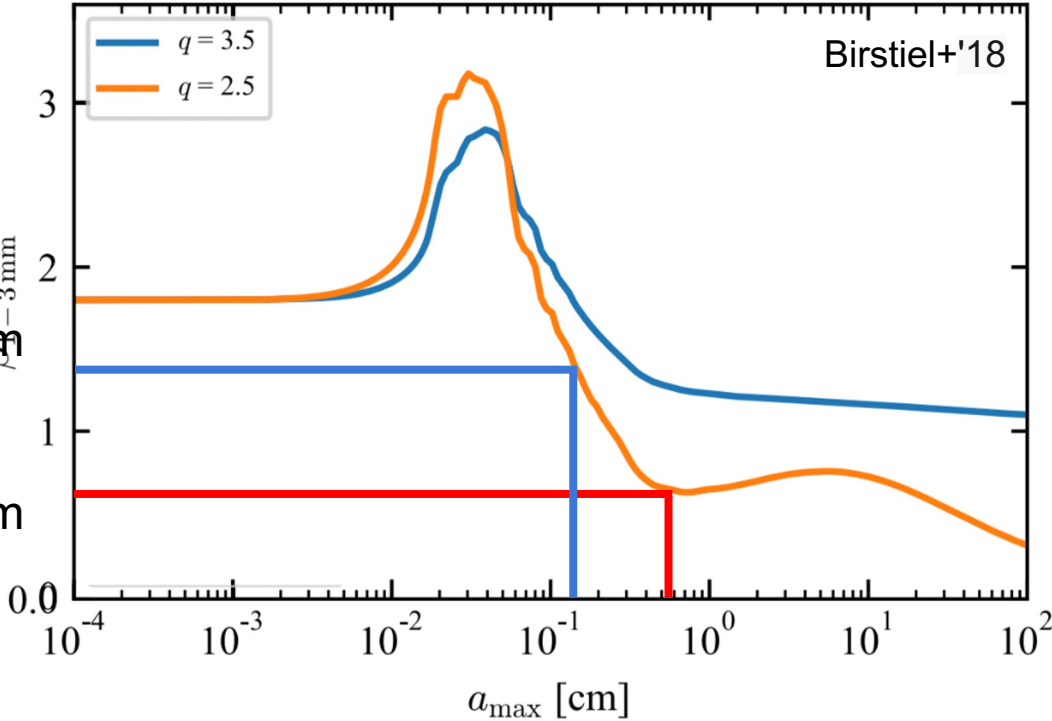
Grain size from SED

Central region

$\sim 4-5$ mm

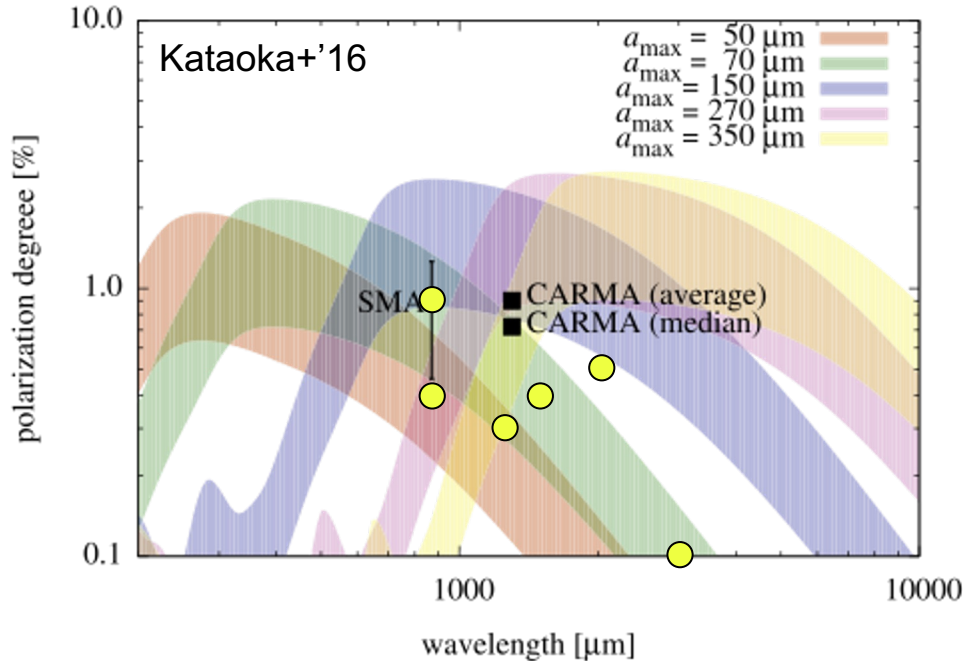
Outer region

$\sim 1-2$ mm



Birstiel+'18

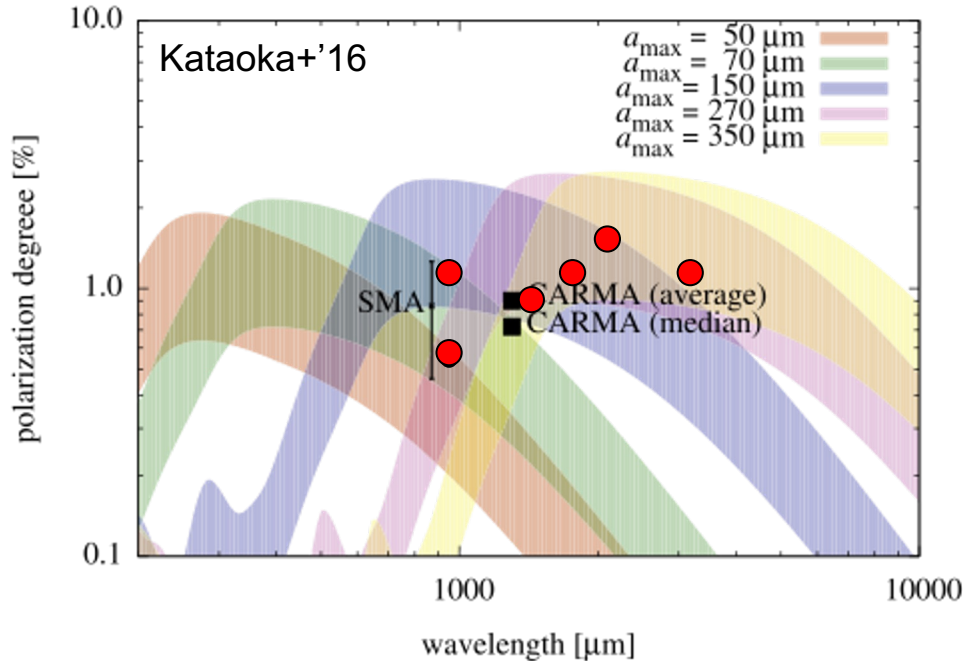
Grain size from polarization



	Wavelength (μm)	Center PF
Band 3	3,000	0.1%
Band 4	2,000	0.5%
Band 5	1,470	0.4%
Band 6	1,280	0.3%
Band 7a	870	0.4%
Band 7b	870	0.9%

Estimated grain size from polarization in Center: $70\mu\text{m}$ (0.07mm)

Grain size from polarization



	Wavelength (μm)	Outer PF
Band 3	3,000	1.3%
Band 4	2,000	1.5%
Band 5	1,470	1.3%
Band 6	1,280	0.9%
Band 7a	870	0.6%
Band 7b	870	1.3%

Estimated grain size from polarization in outer region: $270 \mu\text{m}$ (0.27 mm)

Discrepancy between SED and polarization

SED provided the sizes of 1 - 5 mm.

Polarization fraction provided the sizes of 0.1 - 0.3 mm

Possible reasons:

- Circular patterns cancel polarization at the center.
- Polarization may be more sensitive to surface smaller grains.
- PF model depends on dust properties (Yang & Li 2020).

Summary

Summary

We reduced data of the Class I/II YSO HL Tau to study the usage of polarization. The main results are followings.

- Polarization direction appears circular patterns at Band 3&4, alined to NE-SW at Band 5,6 and 7
- Polarization fraction map shows ring structure at Band 3&4, double-peak structure at Band 6&7a. (Weak in band 5&7b)
- Spectral index is 2 (central), ~ 3 (outer region).
- Grain size estimated from polarization fractions is $70 \mu\text{m}$ (central), $270 \mu\text{m}$ (outer region).
- Grain size estimated from SEDs is 4~5 mm (central), 1~2 mm (outer region).

Thanks!