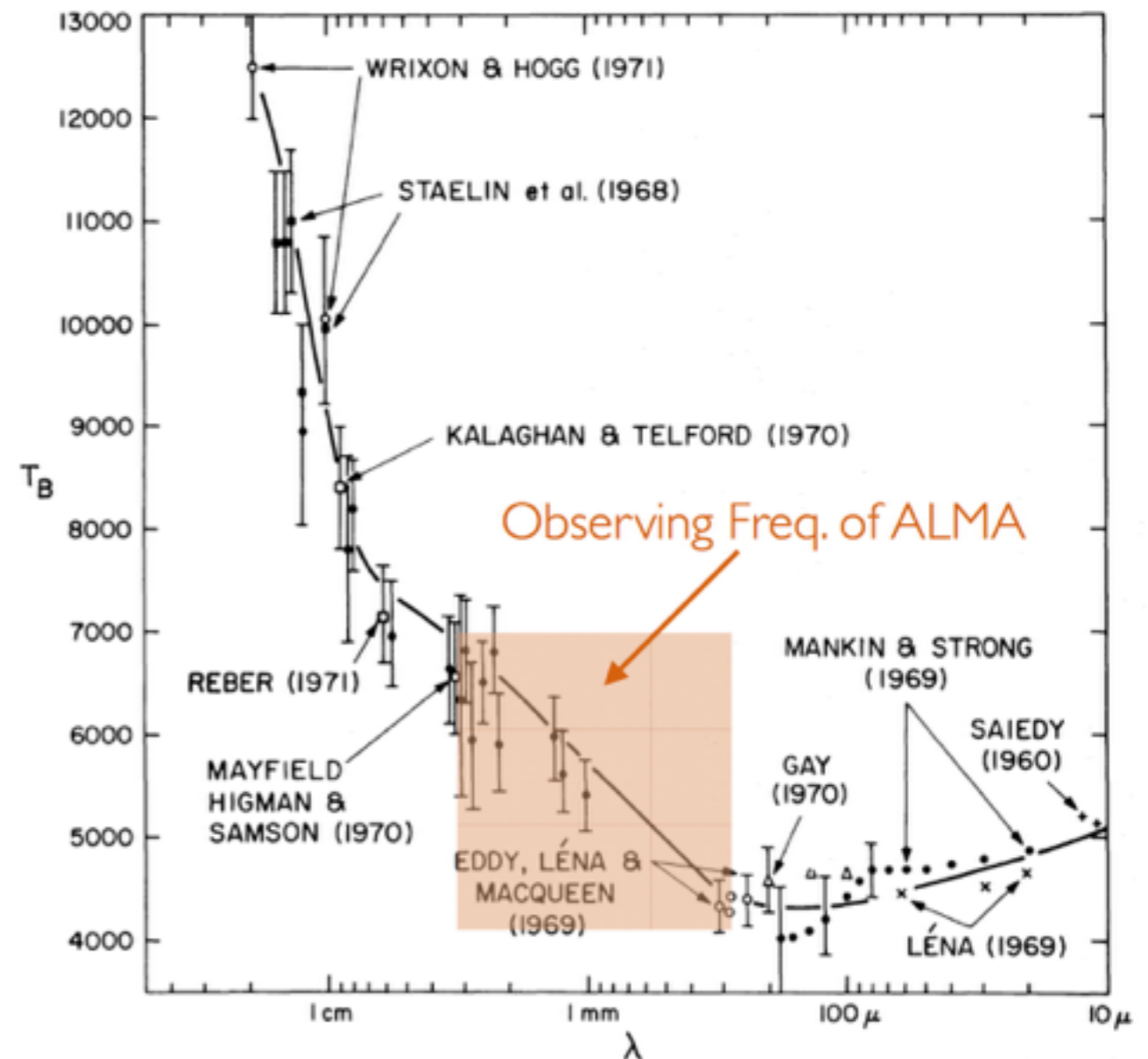


# Solar Observation using ALMA



# Science

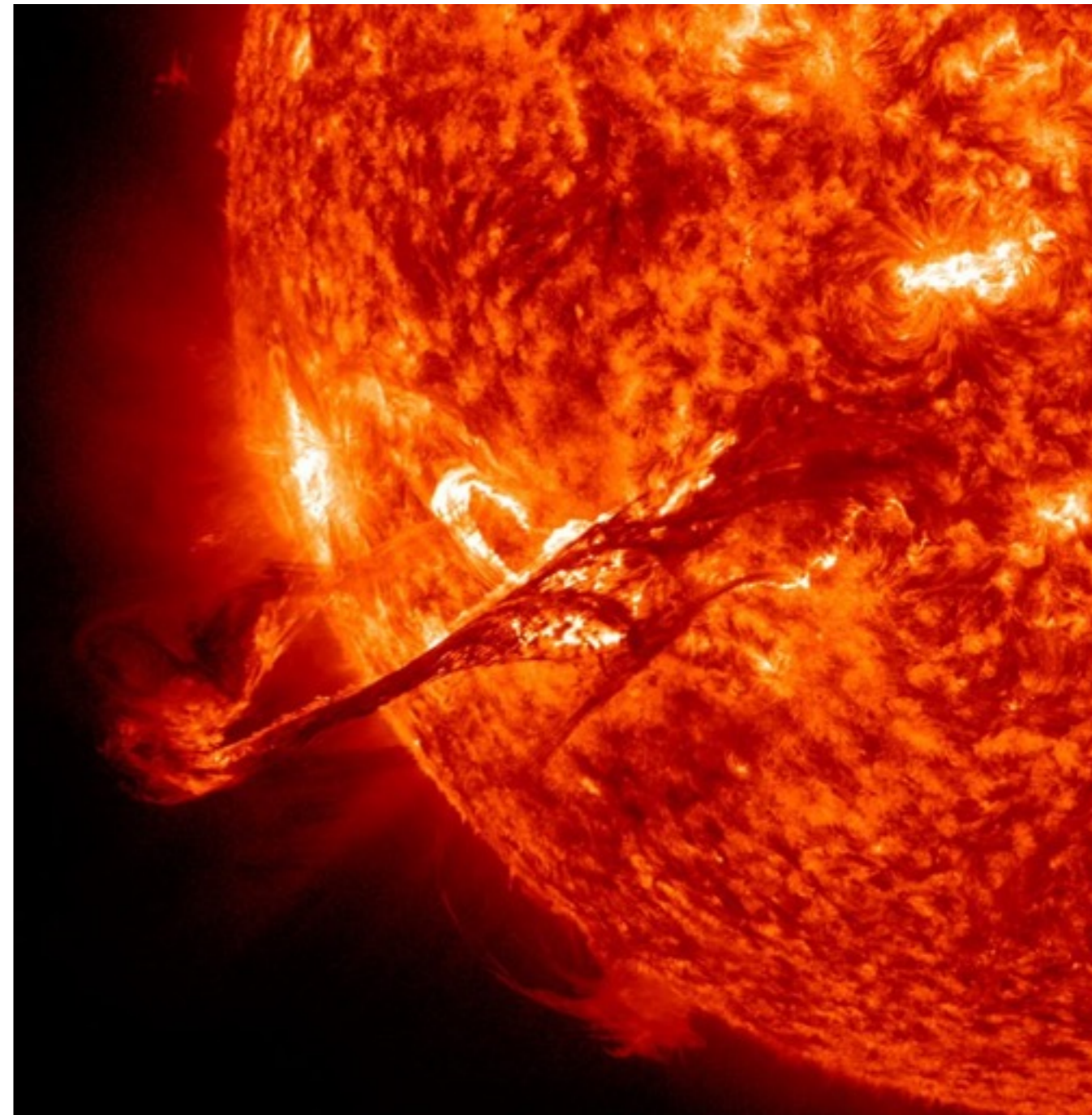
- Chromosphere
  - Continuum radiation at sub-mm, thermal f-f
- Flare physics
  - particle acceleration
  - new insight on spectrum
- Filament/Prominence
- New Diagnostics
  - Oscillation
  - Spectral lines  
(radio recombination line, Co)



Chromospheric spectrum calculated by VALC Model  
Vernazza, Averett & Loeser, 1973

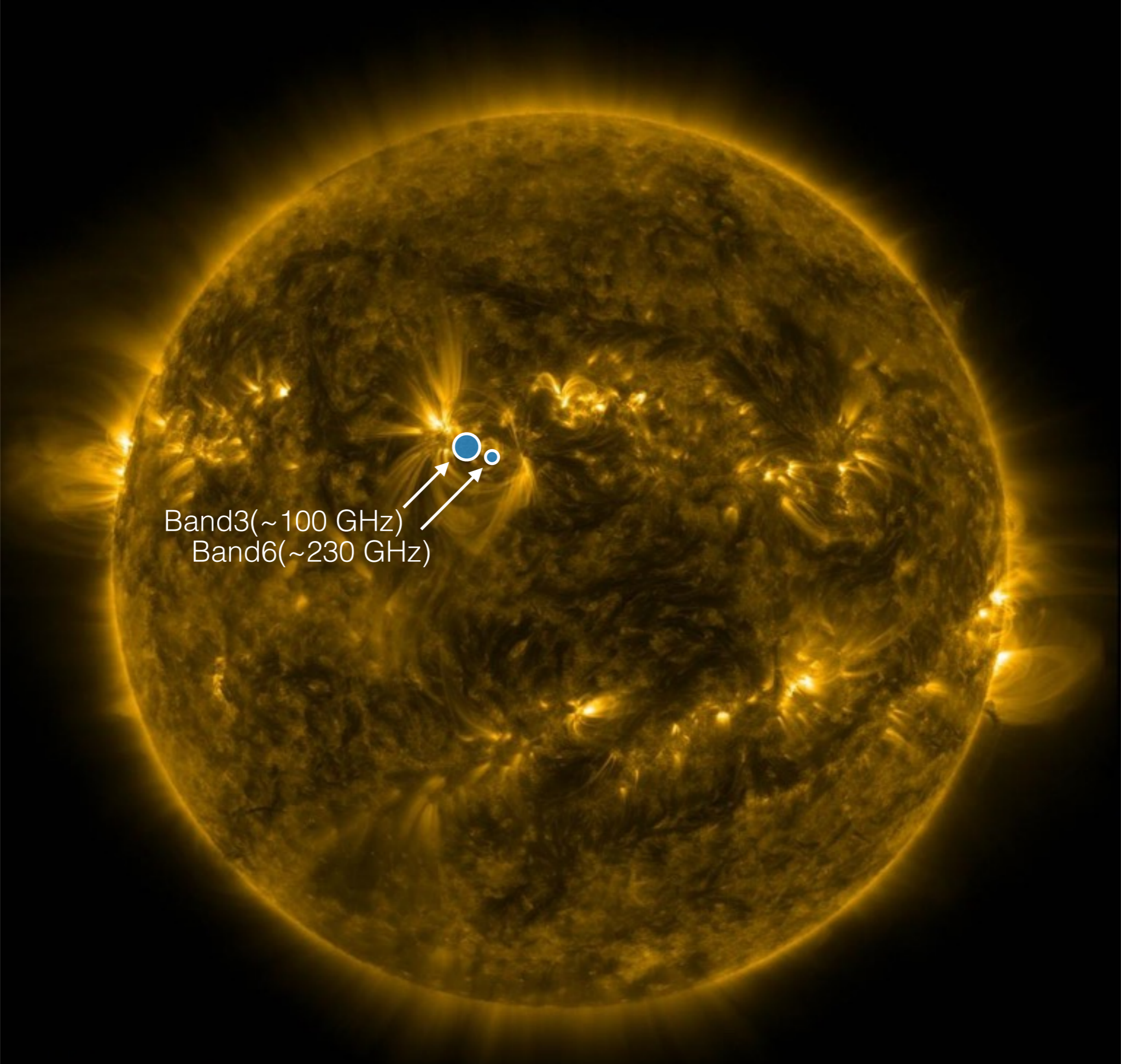
# Science

- Chromosphere
  - Continuum radiation at sub-mm~ thermometer
- Flare physics
  - particle acceleration
  - new insight on spectrum
- Filament/Prominence
- New Diagnostics
  - Oscillation
  - Spectral lines  
(radio recombination line, Co)



# Solar Observation ALMA Issues

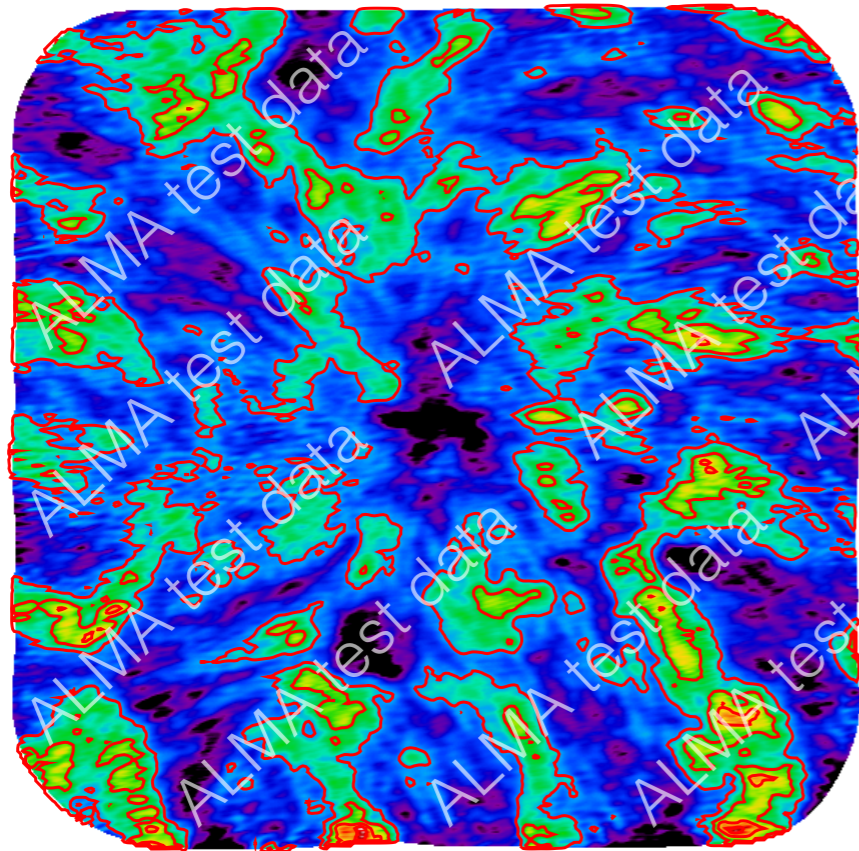
- Sun is big (  $\sim 0.5$  degree)
  - $\ll$  ALMA FoV
- Solar  $T_B \sim 6000$  K (Sis Mixer-Detuning Mode:MD1/2)
  - ALMA sensitivity  $\sim 800$  K
- Variable
  - time scale - ms  $\sim$  hours
  - spacial scale -  $0.1'' \sim$
- Support by coordinate observations
  - ground-based: NST, DKIST(2020), so on
  - satellites: IRIS, Hinode



Band3 (~100 GHz)  
Band6 (~230 GHz)

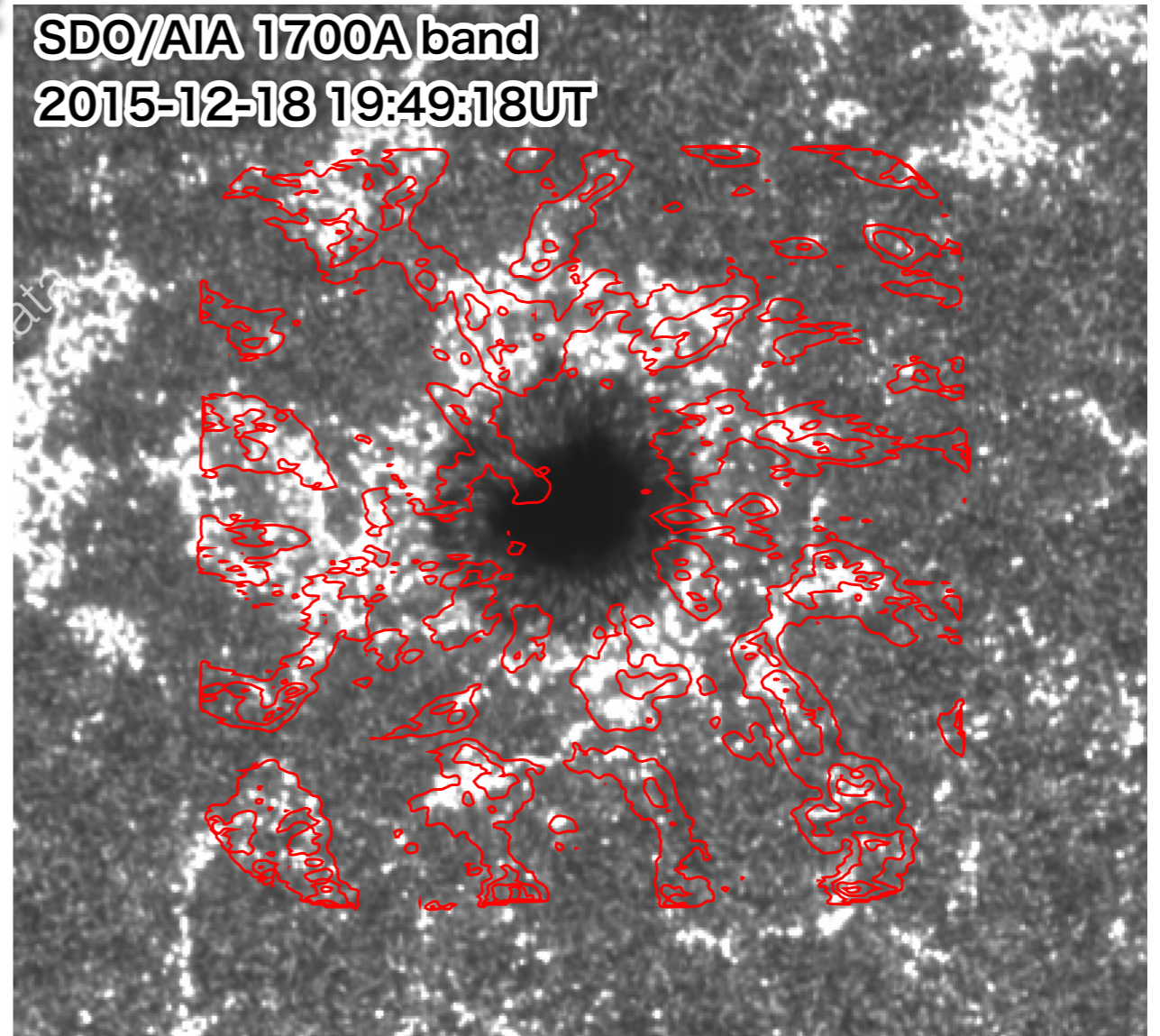
# Mosaic Mode

ALMA 150pts-Mosaic Image  
NOAA 12470 Preceding Sunspot  
2015-12-18 19:39 -- 20:03UT

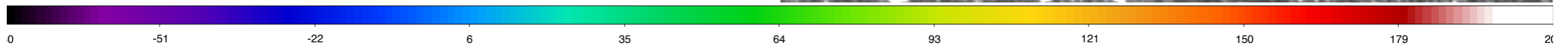


Band6 230 GHz (Synth. Beam 2.6" x 0.7")

SDO/AIA 1700A band  
2015-12-18 19:49:18UT



Red Contours : ALMA 230GHz



# Solar Observation ALMA Issues

- Sun is big (  $\sim 0.5$  degree)
  - $\ll$  ALMA FoV
- Solar  $T_B \sim 6000$  K (Sis Mixer-Detuning Mode:MD1/2)
  - ALMA sensitivity  $\sim 800$  K
- Variable
  - time scale - ms  $\sim$  hours
  - spacial scale -  $0.1'' \sim$
- Support by coordinate observations
  - ground-based: NST, DKIST(2020), so on
  - satellites: IRIS, Hinode

# Coordinate Observations(Groundbased)





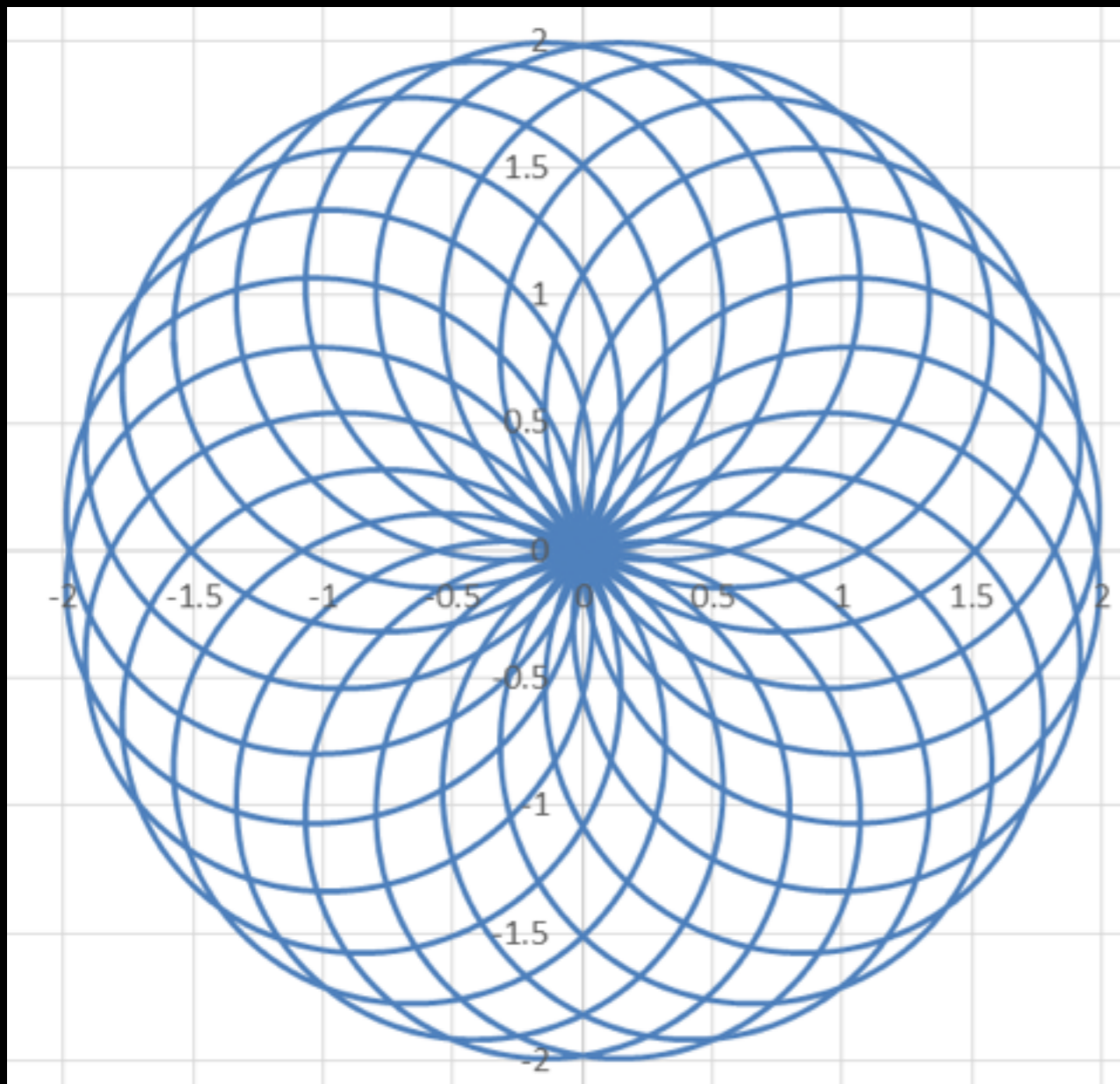
# ALMA capability for Cycle4

- Band3 & Band6 : 4 windows with 128 channels/2 GHz
  - Band3: 93, 95, 105, 107 GHz
  - Band6: 230, 232, 246, 248 GHz
- Spatial Resolution: 1.5"~3.7" @Band3 / 0.63"~1.6"@Band6
  - C40-1(150m), C40-2(273m), C40-3(460m) +ACA
- Single Pointing(2 s), Mosaic Mode(>20 m, max 150-points)
- Period of solar observations in a day: 13:00UT ~ 20:00UT
- Observing Season: 23 December 2016 ~ 27 April 2017
- Single-Dish(12m) array observation support Interferometric observation

# Array

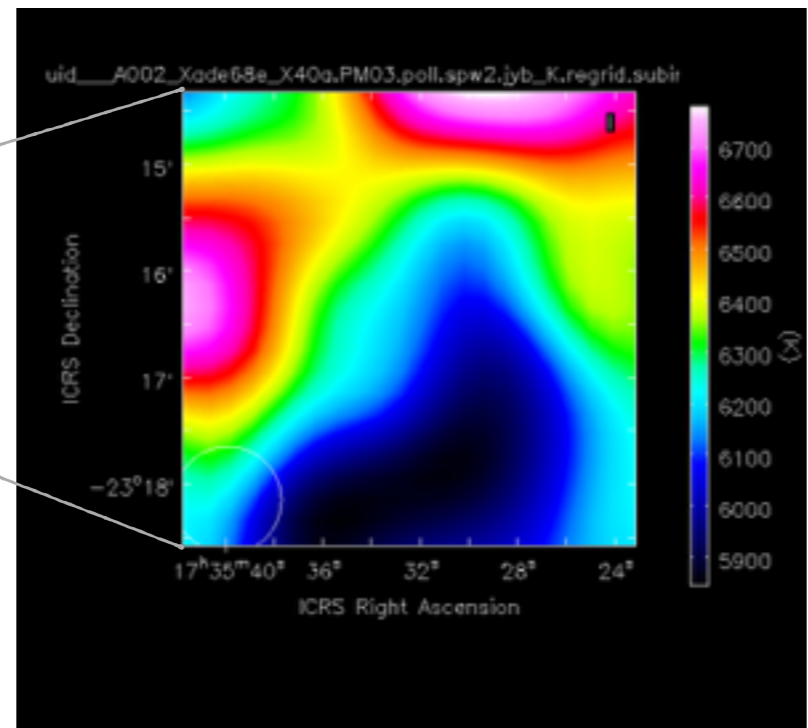
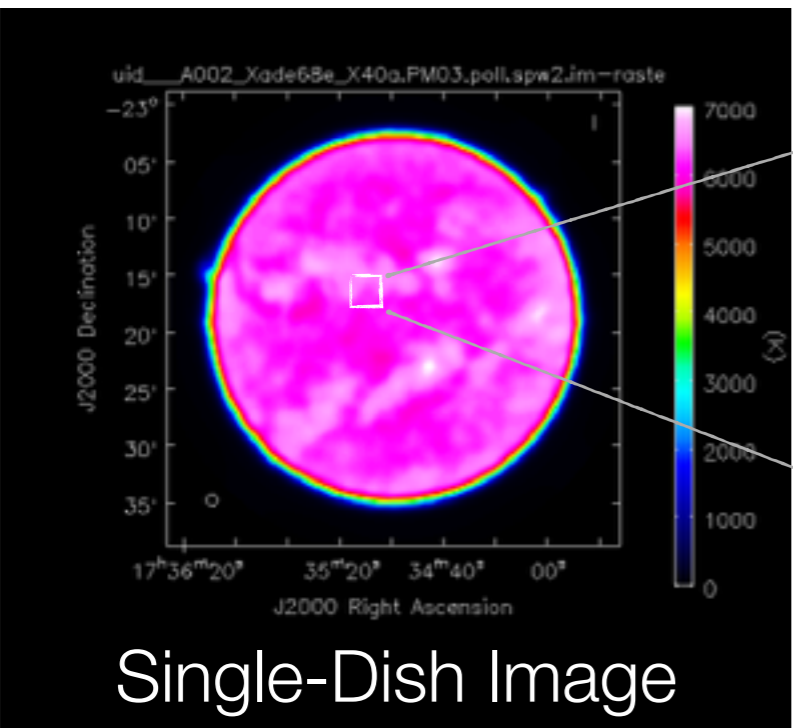
Config	Lmax	Band	Band 3	Band 4	Band 6	Band 7	Band 8	Band 9	Band 10
	Lmin	Freq	100 GHz	150 GHz	230 GHz	345 GHz	460 GHz	650 GHz	870 GHz
7-m Array	45 m	AR	12.5"	8.4"	5.4"	3.6"	2.7"	1.9"	1.4"
	9 m	MRS	66.7"	44.5"	29.0"	19.3"	14.5"	10.3"	7.7"
C40-1	155 m	AR	3.7"	2.5"	1.6"	1.1"	0.80"	0.57"	0.42"
	15 m	MRS	29.0"	19.4"	12.6"	8.4"	6.3"	4.5"	3.3"
C40-2	273 m	AR	2.4"	1.6"	1.0"	0.69"	0.52"	0.37"	0.27"
	15 m	MRS	22.1"	14.8"	9.6"	6.4"	4.8"	3.4"	2.5"
C40-3	460 m	AR	1.5"	0.97"	0.63"	0.42"	0.32"	0.22"	0.17"
	15 m	MRS	13.7"	9.1"	5.9"	4.0"	3.0"	2.1"	1.6"
C40-4	704 m	AR	0.93"	0.62"	0.40"	0.27"	0.20"	0.14"	0.11"
	15 m	MRS	8.9"	5.9"	3.9"	2.6"	1.9"	1.4"	1.0"
C40-5	1.1 km	AR	0.54"	0.36"	0.23"	0.16"	0.12"	0.083"	0.062"
	17 m	MRS	6.0"	4.0"	2.6"	1.7"	1.3"	0.93"	0.69"
C40-6	1.8 km	AR	0.35"	0.23"	0.15"	0.10"	0.076"	0.054"	0.040"
	15 m	MRS	3.1"	2.1"	1.3"	0.90"	0.67"	0.48"	0.36"
C40-7	3.7 km	AR	0.21"	0.14"	0.090"	0.060"	0.045"	0.032"	0.024"
	81 m	MRS	1.8"	1.2"	0.77"	0.52"	0.39"	0.27"	0.20"
C40-8	6.8 km	AR	0.12"	0.079"	0.052"	0.034"	N/A	N/A	N/A
	168 m	MRS	1.3"	0.87"	0.57"	0.38"			
C40-9	12.6 km	AR	0.066"	0.044"	0.029"	N/A	N/A	N/A	N/A
	271 m	MRS	0.78"	0.52"	0.34"				

# Single-Dish scanning (Band9)



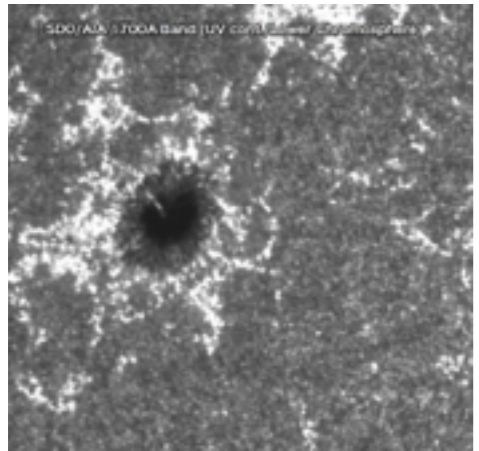
# Combine interferometric image with single-dish image

## Example: 149pts-Mosaic Observation with Band3

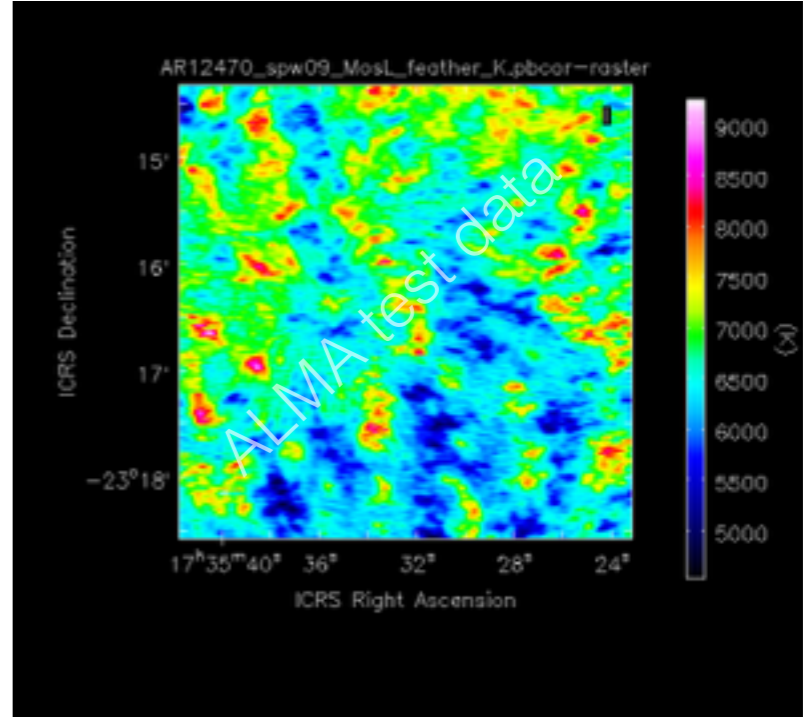
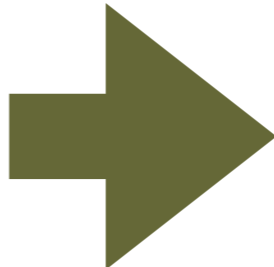
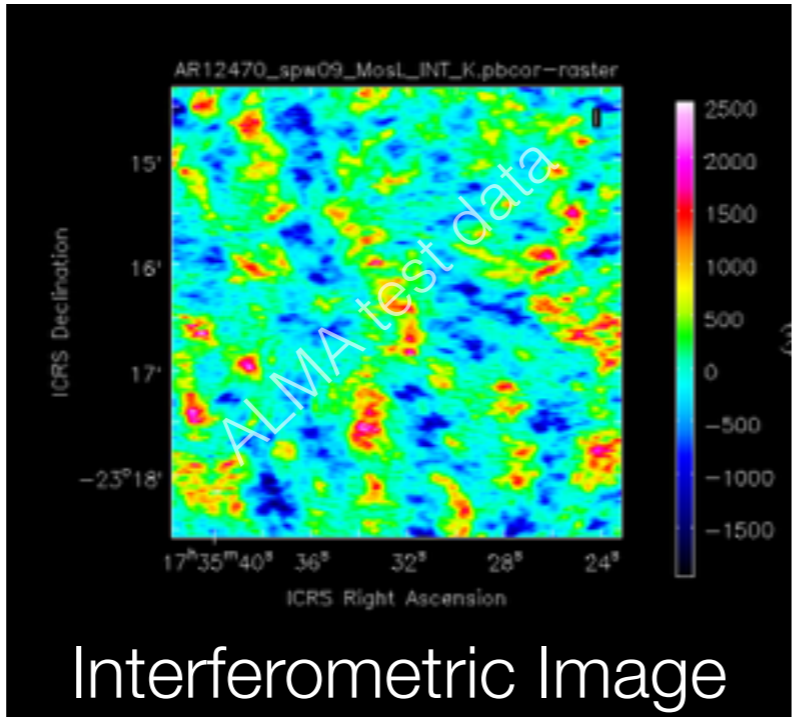


Unit: Brightness Temperature [K]

**Feathering**



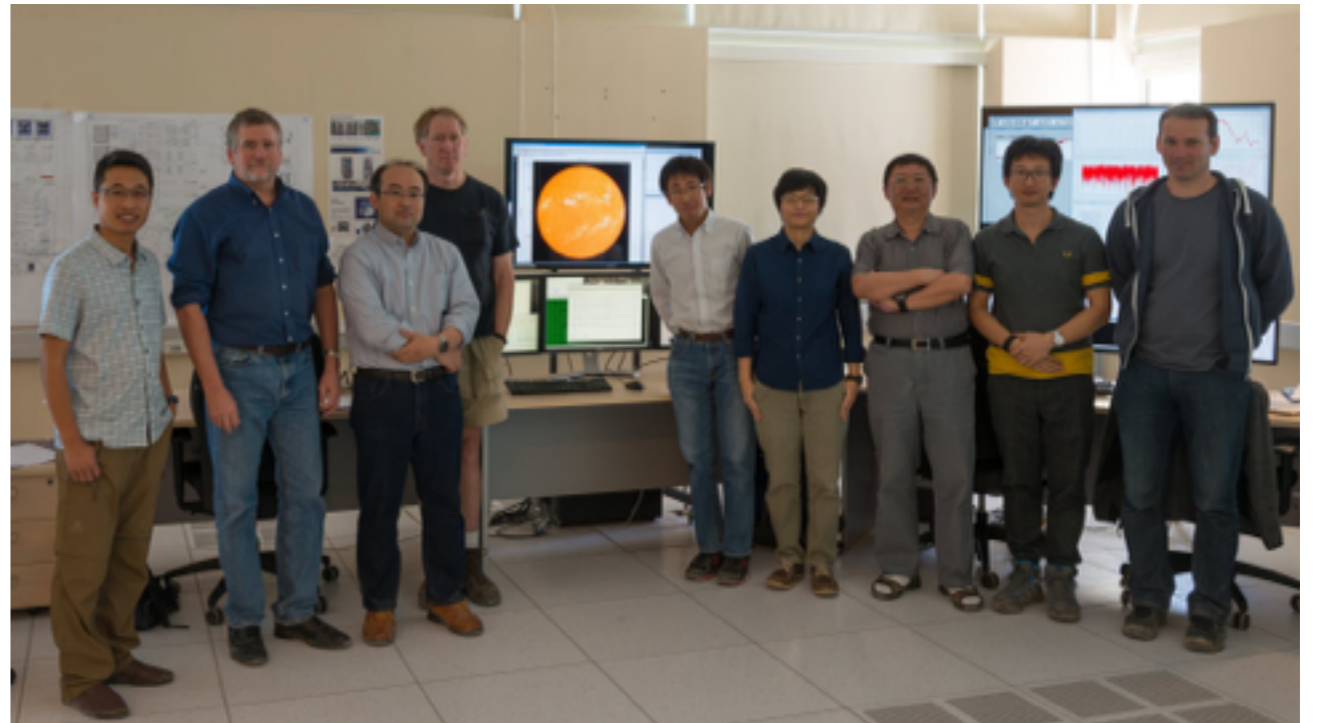
SDO/AIA 1700A Band



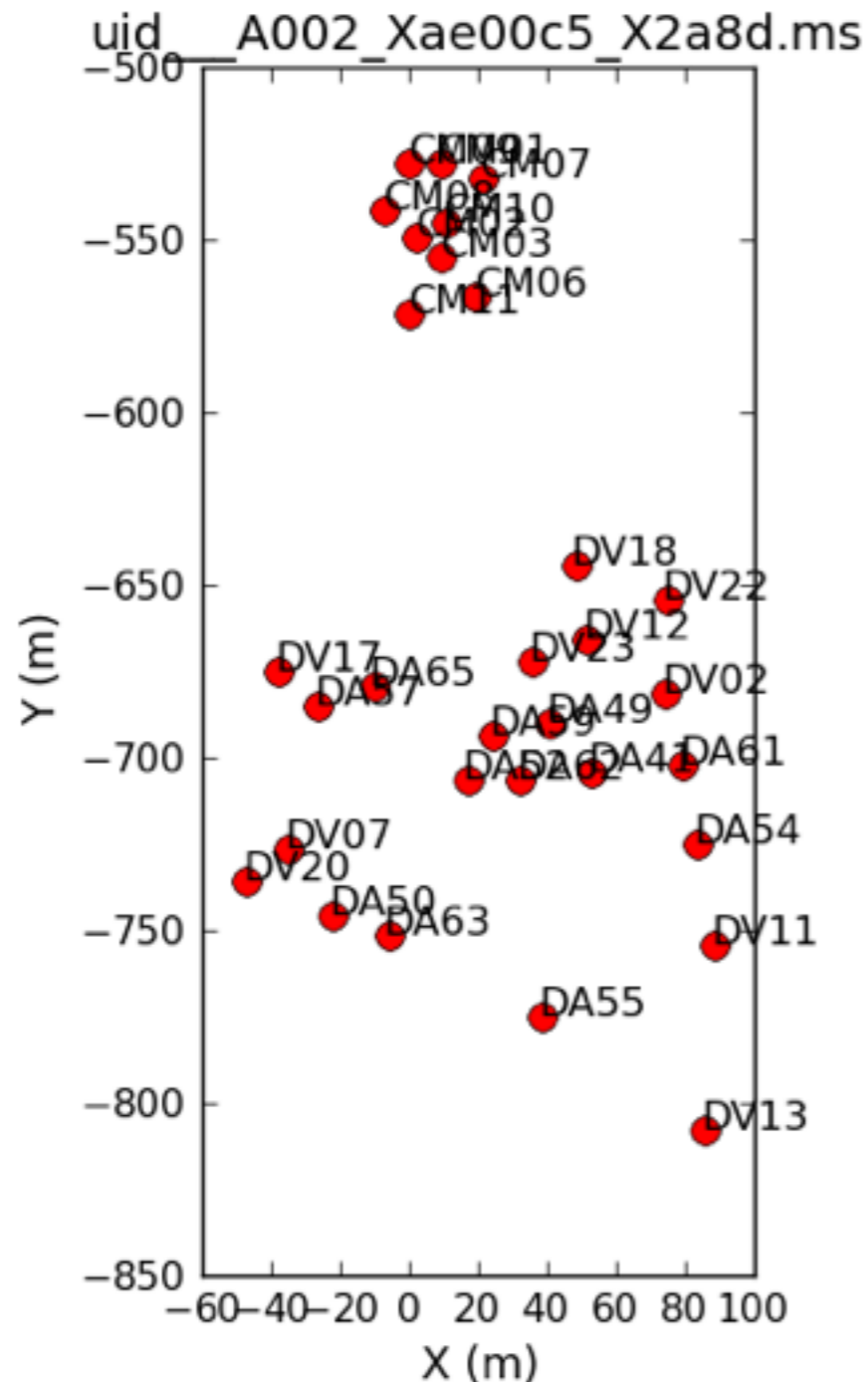
# Solar Commissioning Campaign 2015

- 2015/12/14 - 2015/12/21
  - We could not carry out observations in afternoon 19-21 Dec, because the wind in the antenna site is too strong mainly.
- Goals: **Dress-Rehearsal of Cycle4** / New function of Cycle5
- # of antennas: 20~23 12-m + 7~8 7-m antennas. Total: **27~31 antennas**... with C36-1 (Most compact configuration)
- Lead of The Campaign: Tim Bastian (NRAO) Cycle4 : 12m x 40 + 7m x 10 = 50 antennas
  - Lead of the interferometer: Masumi Shimojo (NAOJ)
  - Lead of the single-dish: Stephen White (AFRL)
  - Supporters from Joint ALMA Observatory: Antonio Hales, Neil Philips, Akihiro Hirota (Remote)
- EastAsia (EA)
  - Masumi Shimojo (NAOJ)
  - Kazumasa Iwai (NICT)
  - Sujin Kim (KASI)
- Europe (EU)
  - Ivica Skokie (Ondřejov Observatory)
- NorthAmerica (NA)
  - Tim Bastian (NRAO)
  - Stephen White (AFRL)
- China (as participants of NA development team)
  - Yihua Yang, Sijie Yu, Donghao Liu (NAOC)

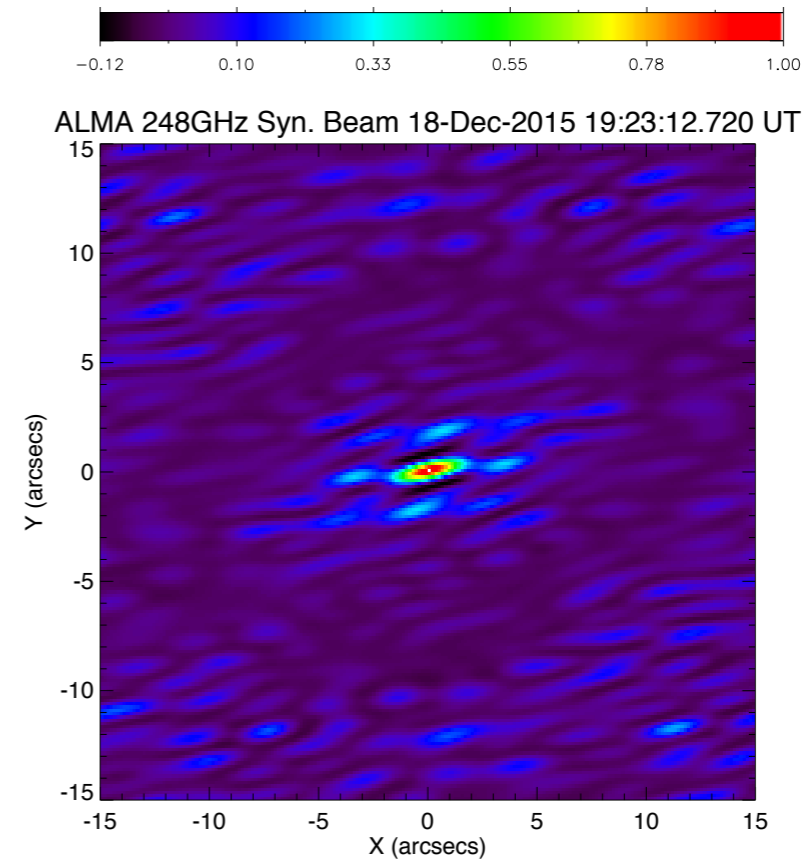
Joint Observations with  
Hinode, IRIS, NST/BBSO



# Typical antenna configuration & synthesized beam during solar development campaign 2015



Pseudo C36-1 configuration  
(The number of antennas is shortage.)



- HPFW of Synthesized Beam
  - Band3: 4.8" x 1.4"
  - Band6: 2.8" x 0.7"

### Spatial resolution in Cycle 4

	Band3	Band6
C40-1+7m	3.4"	1.5"
C40-2+7m	1.8"	0.8"
C40-3+7m	1.2"	0.5"

# Observed Targets of the 2015 solar campaign

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- 149pts-Mosaic Observations
  - **Band3 / Band6: AR12470**
  - **Band6: South Pole**
- 39pts-Mosaic Observations
  - **Band3: Prominence**
  - Band6: Quiet Sun near the Limb
  - ~~Band3: Quiet Sun on the disk~~
- Single-Pointing
  - **Band3 / Band6: AR12470**

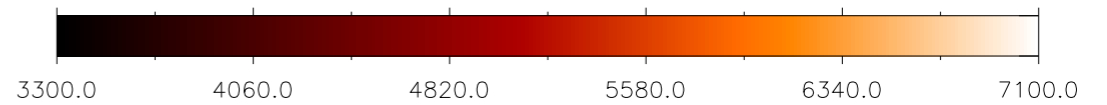
## Preliminary Results!!

- **FoV of the 149pts MOSAIC**
  - **350"x350" for Band3**
  - **150"x150' for Band6**
- **FoV of Single-points:**
  - **60" for Band3/ 25" for Band6**
- **The coordinate of most images is the RA/DEC coordinate.**

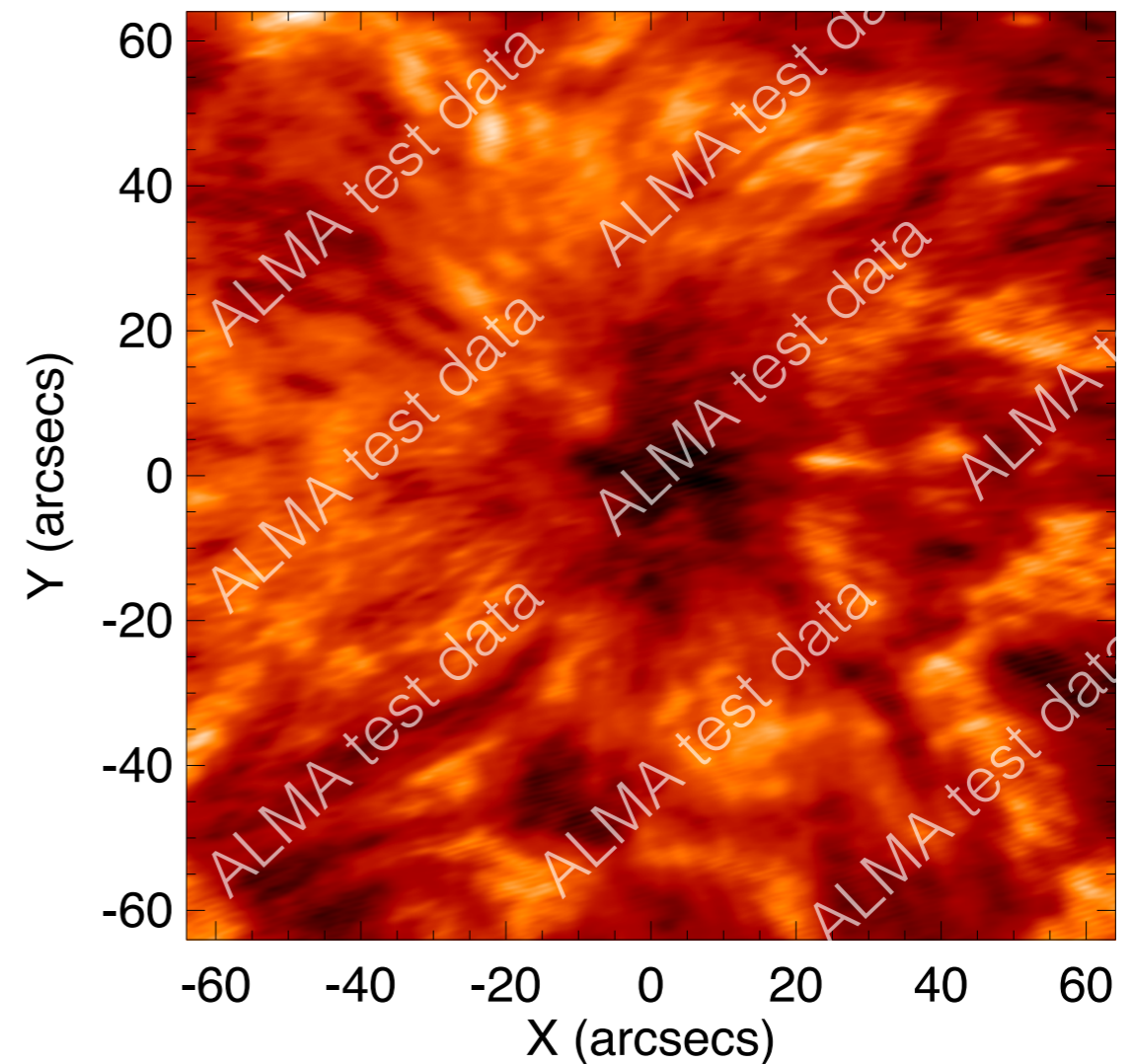
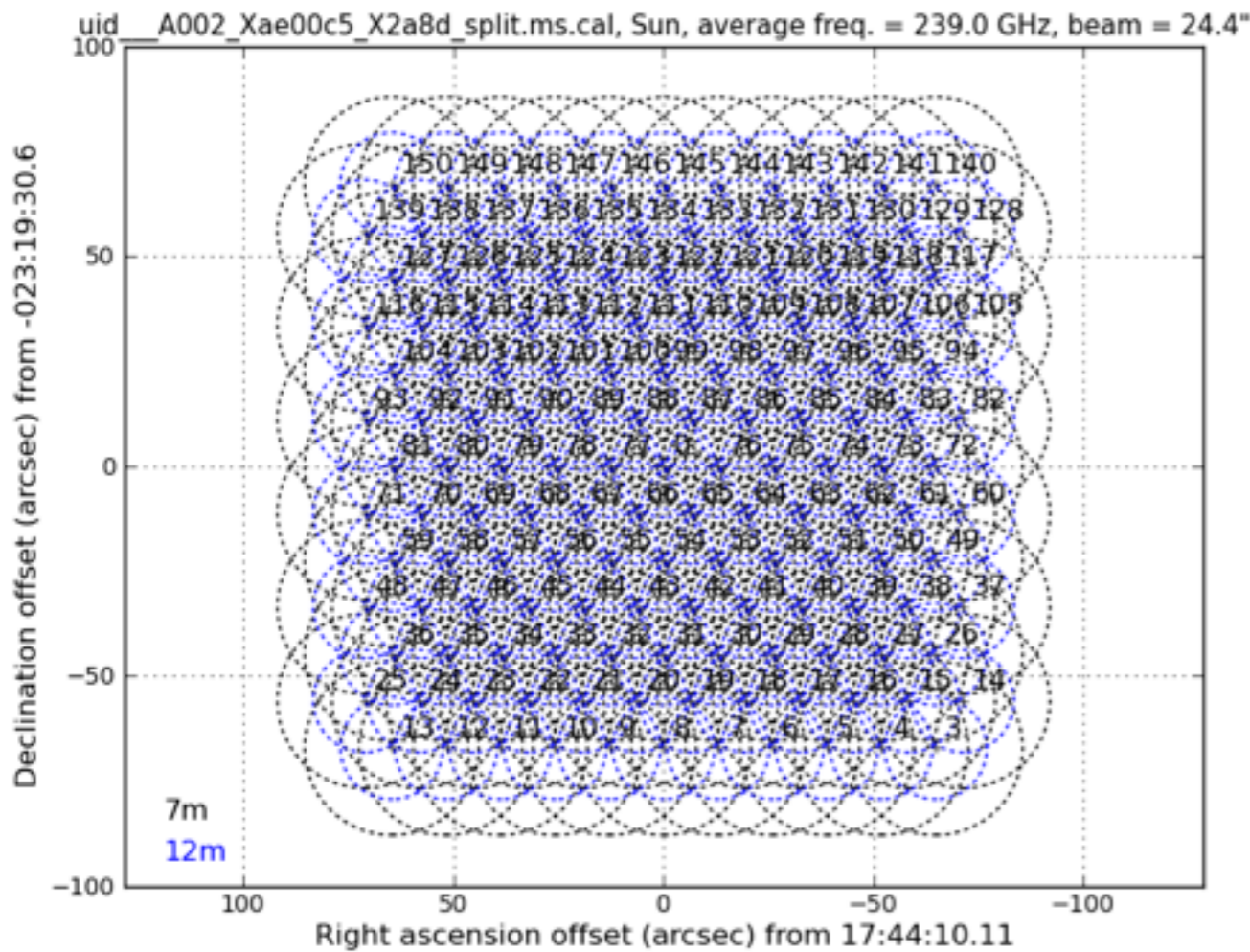
# NOAA12470 Preceding Sunspot on 18-Dec-2015

## Band6-248GHz 149pts-MOSAIC observation

Unit: Brightness Temperature [K]



ALMA 248 GHz 18-Dec-2015 19:23:12.720 UT

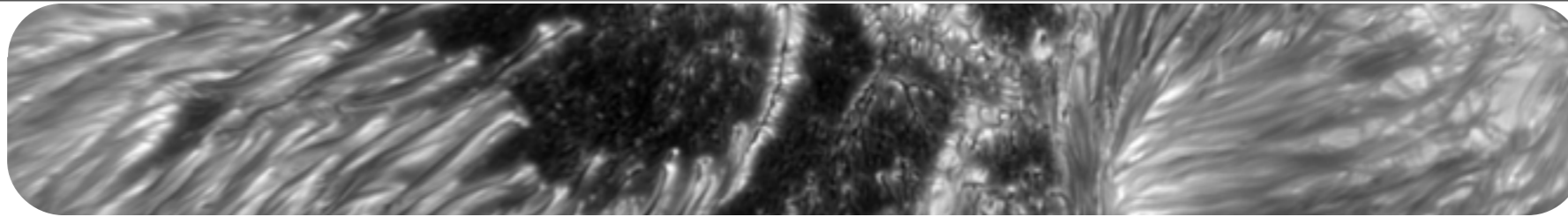




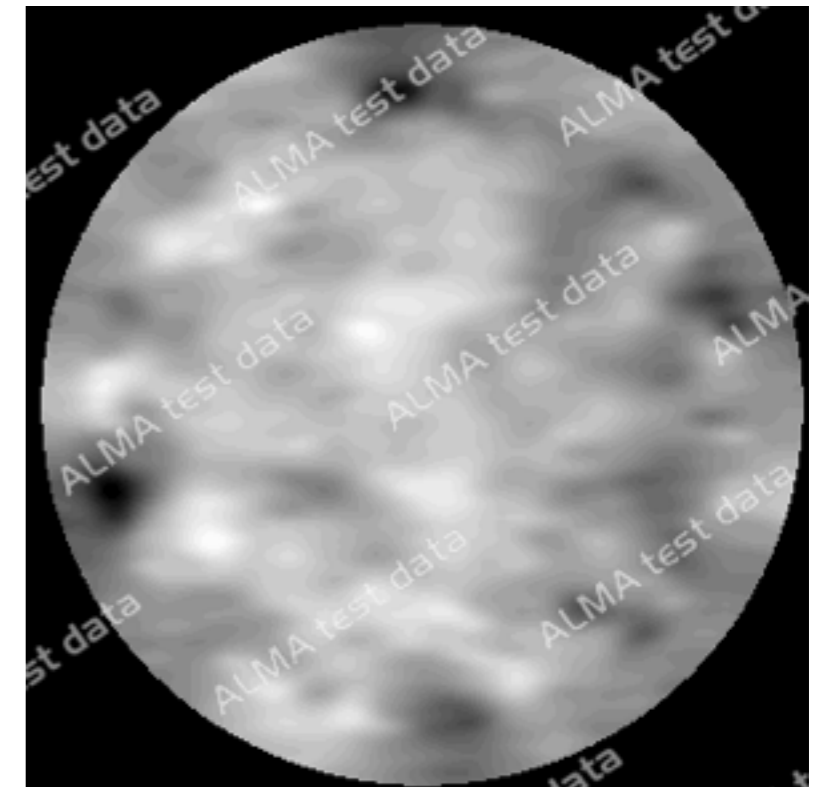
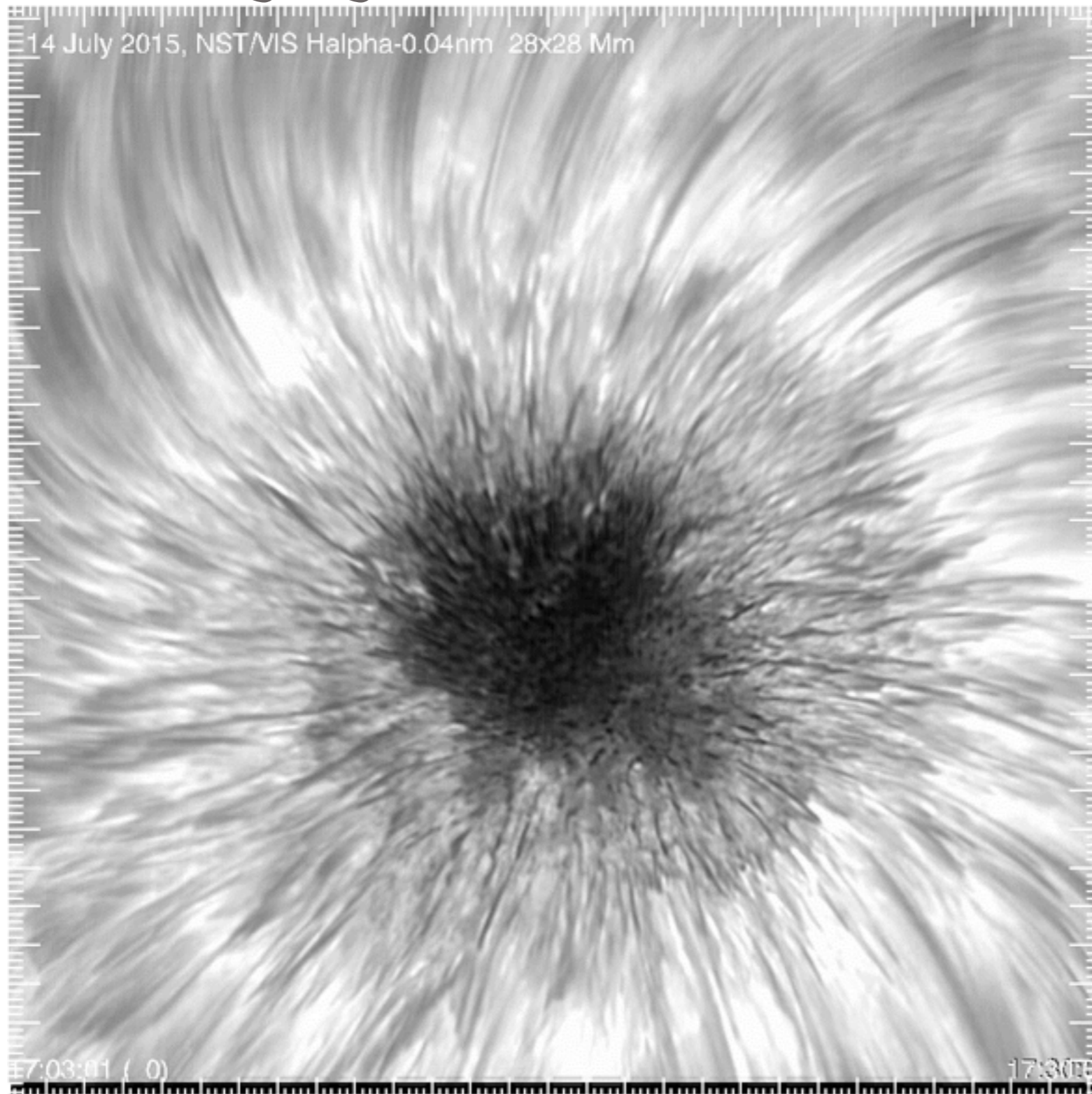
# Time cadence of MOSAIC images

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- The required time for one point: 7.6 sec
  - [Integration Time : 6 sec] + [Overhead: 1.6 sec]
- Time cadence of the MOSAIC image
  - **# of point in one MOSAIC image x 7.6 sec**
  - **Exception**
    - When the phase calibration is carried out during scanning of one map, the required time for one map becomes to **# of point x 7.6 sec + 3mins.**
    - **Ex. 149-pt: 7.6s x 149 = 19 mins / Actual: ~22 mins**



# Science highlights—umbral waves



Example ALMA movie  
*Dec. 2016 campaign (2 s)*  
*(Yi Chai)*

Yurchyshyn et al., in  
preparation

# Cycle5

- Band7 & Band9
- Spectral line in Band3/6/7/9
- Polarimetry
- Subarray
- Fast TP mapping
- Integration time <2s