

# Disk Substructures of AS 209 from Dust Continuum and 12CO Emission Line

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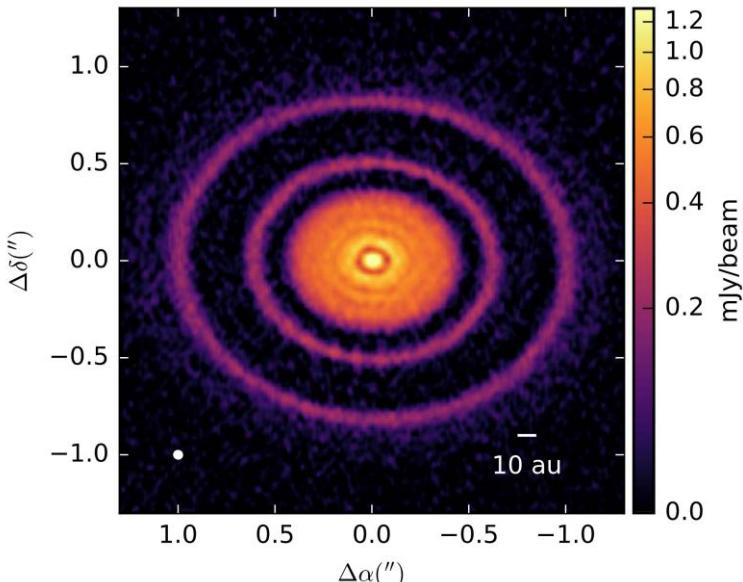
3 Astronomy Program, Department of Physics and Astronomy, Seoul National University

# Introduction

Guzmán et al.

## AS 209

- Ophiuchus,  $121 \pm 2$  pc
- Spectral type K5
- Mass  $\sim 0.9$  Msun
- Age  $\sim 1.6$  Myr
- Mdisk  $\sim 0.028$  Msun
- Inclination  $\sim 40$  deg
- Position angle  $\sim 85.8$  deg



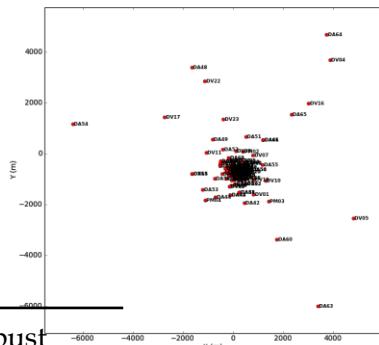
Dust continuum emission map

# AS 209 dust continuum

Dust continuum Band 6 (230 GHz), 157 antennas

Spectral window 0 ~ 8 (P.I. Oberg)

Weighting function natural, uniform, robust(Briggs)



Statistics	Natural	Uniform	Robust
Beam size (maj./min.) ["]	0.55 / 0.49	0.42 / 0.37	0.46 / 0.45
Total flux density [Jy]	0.274	0.25	0.27
Peak flux density [Jy]	0.08	0.056	0.07
Mean rms [mJy/beam]	0.18	0.5	0.2

Statistics	Oberg	Fedele	Andrew	Total
Angular resolution ["]	~ 0.5	~ 0.1	~ 0.03	~ 0.03
Beam size (maj./min.) ["]	0.46 / 0.45	0.17 / 0.16	0.06 / 0.04	0.06 / 0.04
Total flux density [Jy]	0.27	0.29	0.43	0.5
Peak flux density [Jy]	0.07	0.014	0.003	0.003
Mean rms [mJy/beam]	0.2	0.05	0.03	0.02

# AS 209 dust continuum

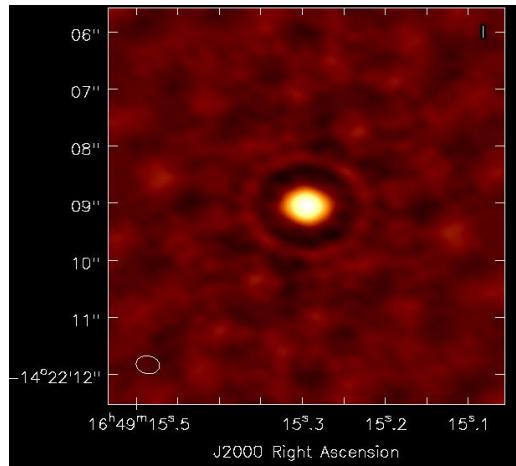
## self\_calibration trial

spw [ 0~8] split -> phase.caltable , amp.caltable  
spw [ 9~13] split -> phase.caltable , amp.caltable

-> apply them to spw[0~24] , but it didn't work

spw[0~24] -> phase.caltable, amp.caltable : self\_calibration

-> beam size =  $0.41''$ ,  $0.31''$  ( before  $0.064''$ ,  $0.039''$  )  
-> rms noise  $\sim 30$  mJy/beam



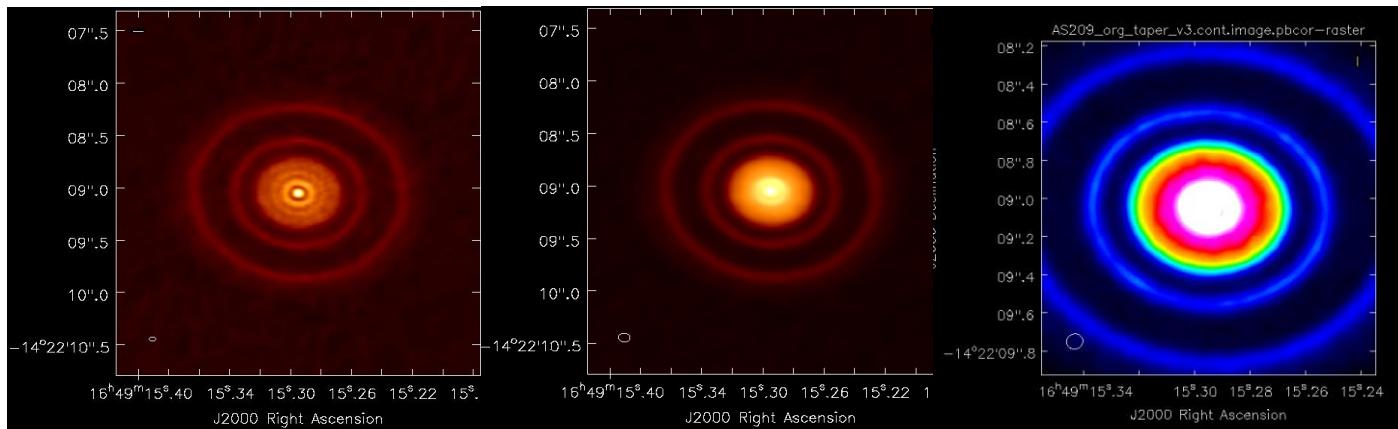
# AS 209 dust continuum

After cleaning : beam :  $0.064'', 0.041'', 85.3 \text{ deg}$

```
# uvtaper = [ '0.064arcsec','0.064arcsec','0deg']  
=> beam size :  $0.11'', 0.081'', 85.5 \text{ deg}$ 
```

```
# uvtaper = [ '0.065arcsec', '0.025arcsec', '-5deg']  
=> beam size :  $0.086'', 0.079'', -73.2 \text{ deg}$ 
```

```
# restoring beam = [ '0.064arcsec' ]  
=> beam size :  $0.064'', 0.039''$ 
```



# AS 209 dust continuum

## Estimating disk mass

$\kappa_\lambda$  (gas + dust) is taken to be  $0.01(1.3 \text{ mm}/\lambda) \text{ cm}^2 \text{ g}^{-1}$  from Ossenkopf & Henning (1994) and assumes  $M_{\text{gas}}/M_{\text{dust}} = 100$ . The disk mass (gas + dust) is then given by

$$M_{\text{disk}} = 0.06 M_\odot \frac{F_\lambda}{1 \text{ Jy}} \left( \frac{d}{100 \text{ pc}} \right)^2 \frac{50 \text{ K}}{\langle T \rangle} \frac{0.01 \text{ cm}^2 \text{ g}^{-1}}{\kappa_{1.3 \text{ mm}}} , \quad (6)$$

where  $F_\nu$  is the observed flux at 1.3 mm in Jy. The observational data and resulting masses are summarized in Table 7. The errors in the observed fluxes are taken to be  $\sim 30\%$ .

(Thi+2001)

adopting  $T_{\text{dust}} = 20K$

$$M_{\text{dust}} = \frac{F_\nu d^2}{\kappa_\nu B_\nu(T_{\text{dust}})} = 1.25 \times 10^{-6} M_\odot \left( \frac{F_{230\text{GHz}}}{1mJy} \right) \left( \frac{d}{121pc} \right)^2$$

$$F_{230\text{GHz}} = 386mJy$$



$$M_{\text{dust}} = 4.8 \times 10^{-4} M_\odot$$

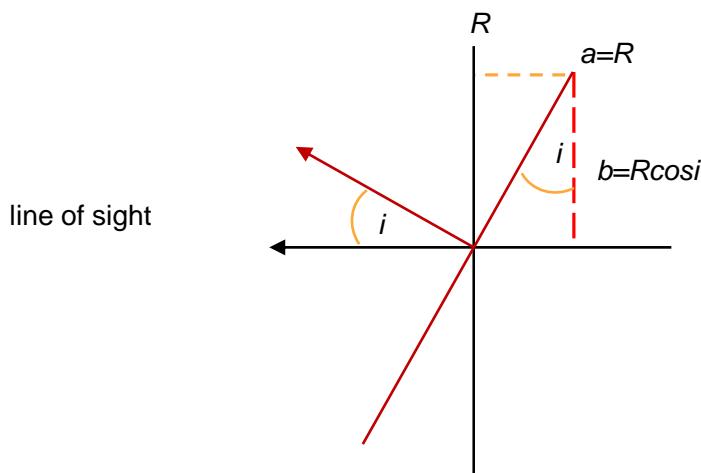
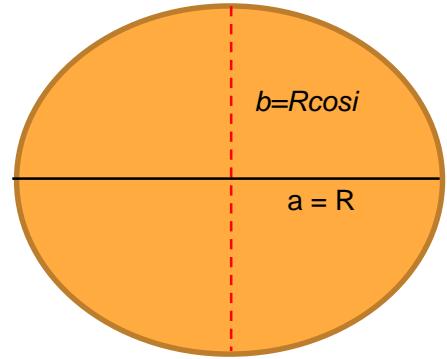
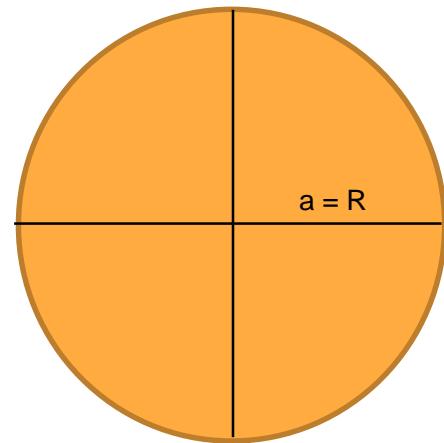
(using all spwIDs)

$$M_{\text{gas+dust}} = 4.8 \times 10^{-2} M_\odot$$

$$M_{\text{gas+dust}} = 2.8 \times 10^{-2} M_\odot$$

(Andrews+2009)

Inclination angle  $i$



$$i = \cos^{-1} \left( \frac{b}{a} \right)$$

$$\begin{aligned} a &= 2.73'' \\ b &= 2.12'' \end{aligned}$$

$$i = \cos^{-1} \left( \frac{2.12}{2.73} \right) \approx 40^\circ$$

$$i = 34.9^\circ$$

(Guzmann+2019)

# AS 209 12CO emission line

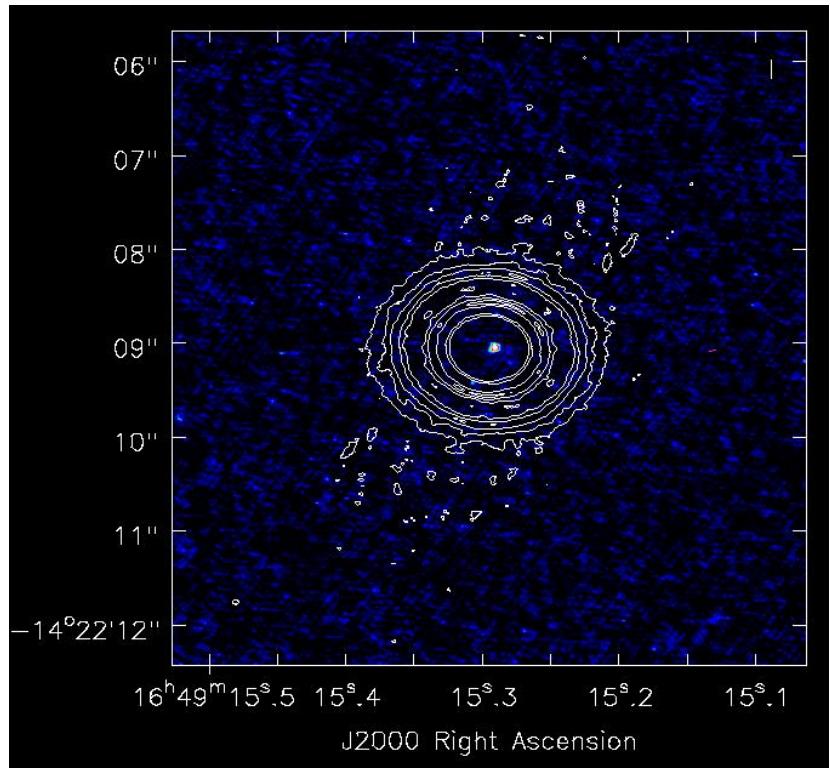
12CO line emission

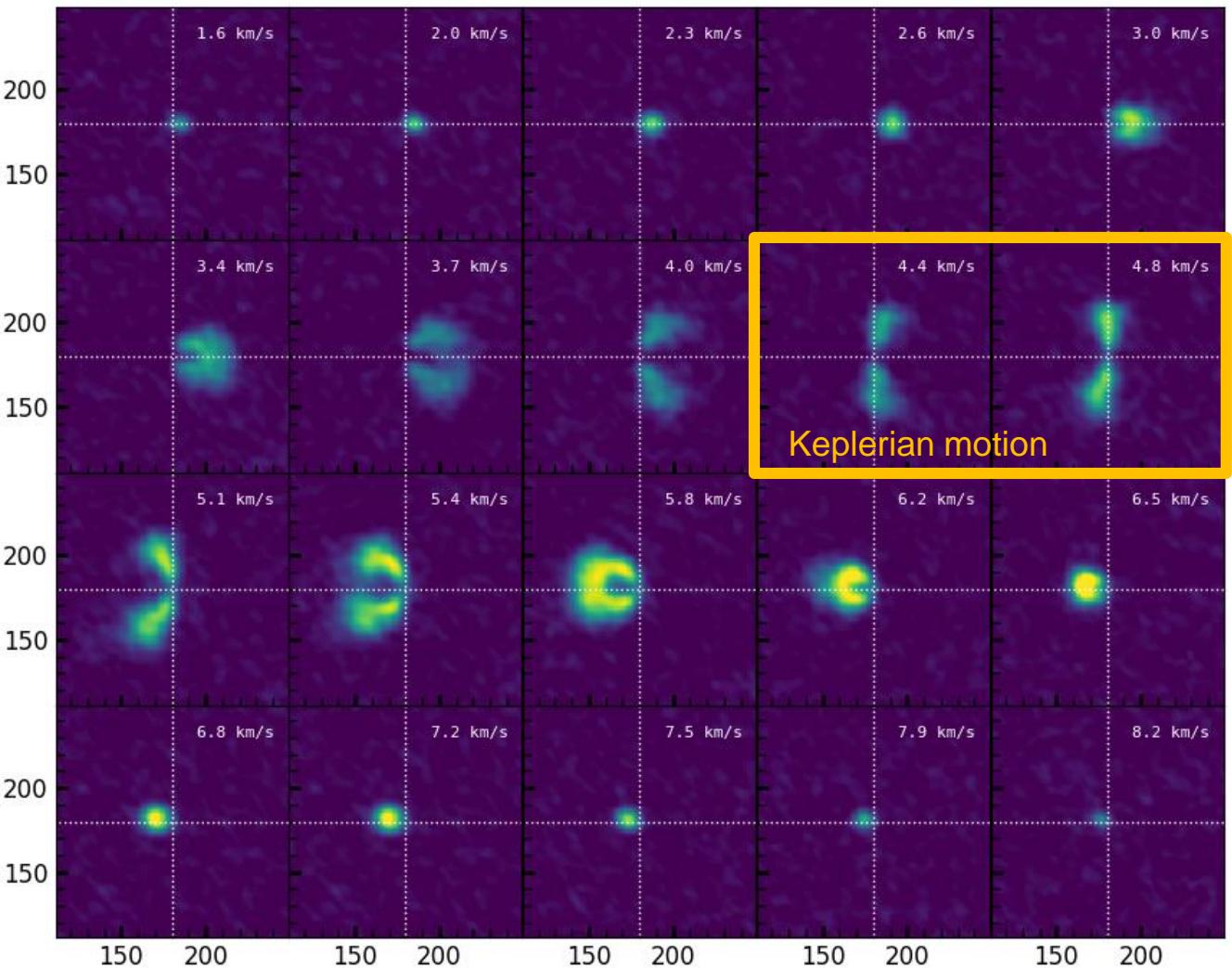
spw [0~2]

ChanWid : 269.147kHz

CtrFreq : 230.533 GHz

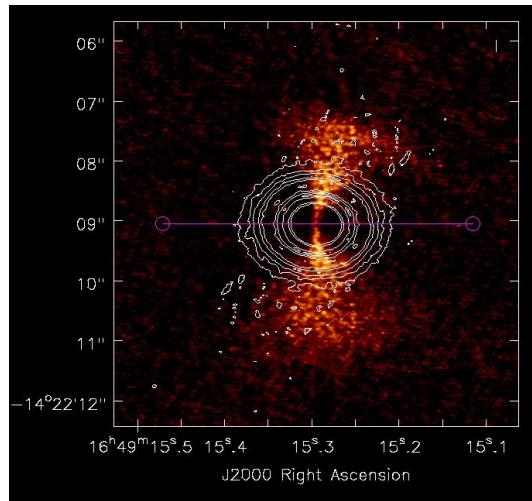
vel ~ 0.35 km/s





Keplerian motion

# AS 209 12CO emission line



PV diagram

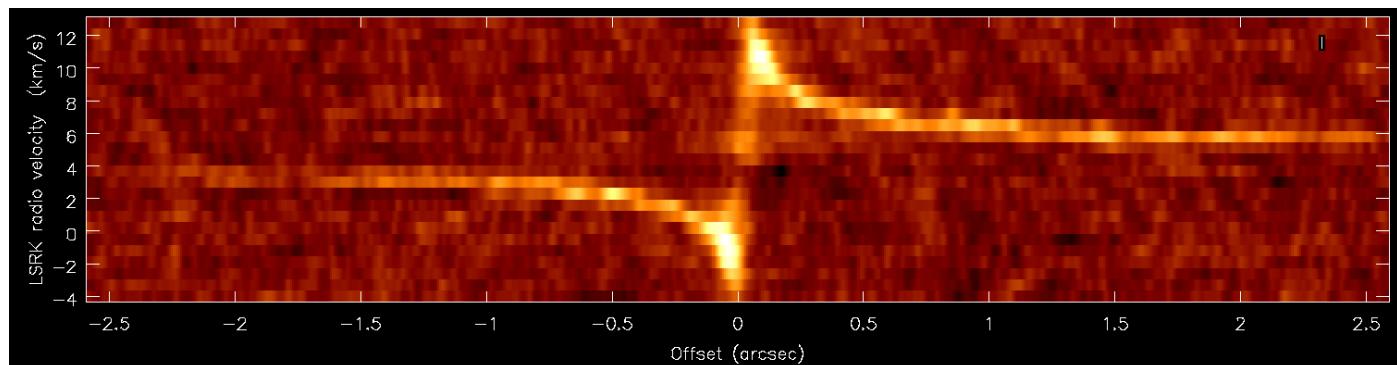
$$\frac{mv^2}{r} = \frac{GM_\star m}{r^2}$$

Velocity deprojection using inclination

0.5 arcsec → 1.2 Msun

1.0 arcsec → 0.88 Msun

2.0 arcsec → 0.95 Msun

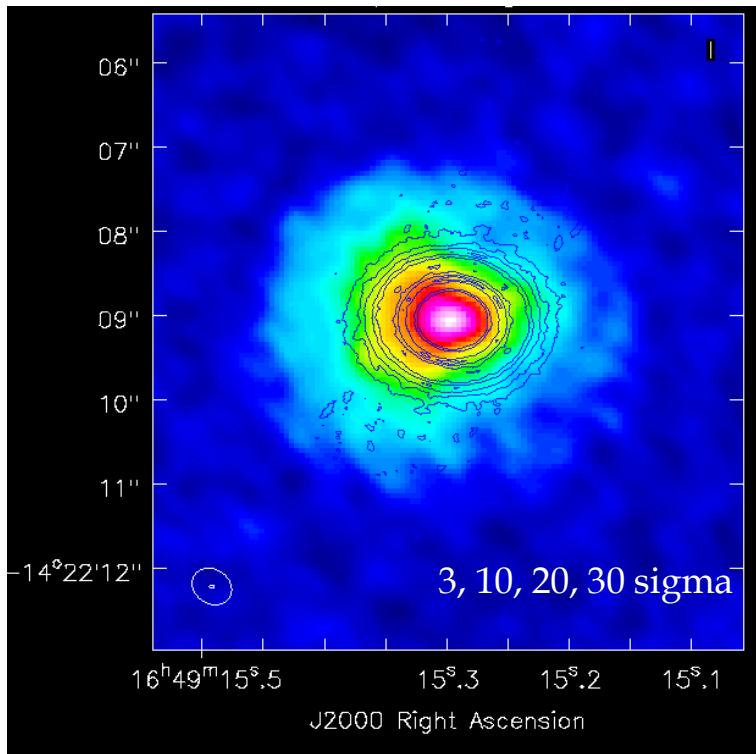


# AS 209 12CO emission line

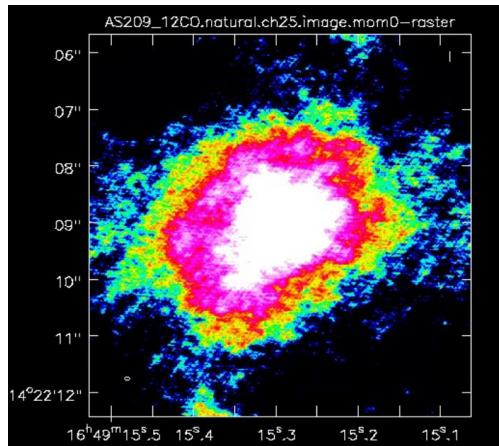
Moment map

Spectral window 0 ~ 2

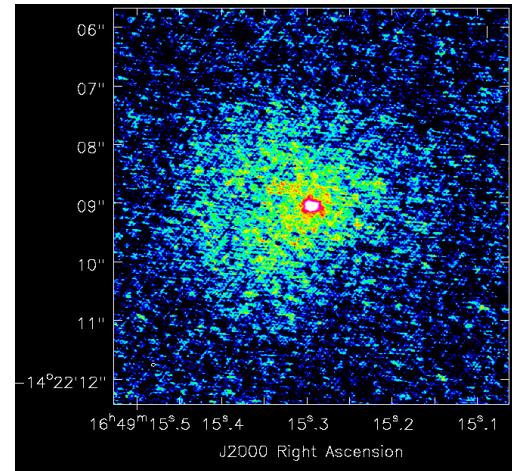
Gas disk is more extended toward NE direction than the dust disk



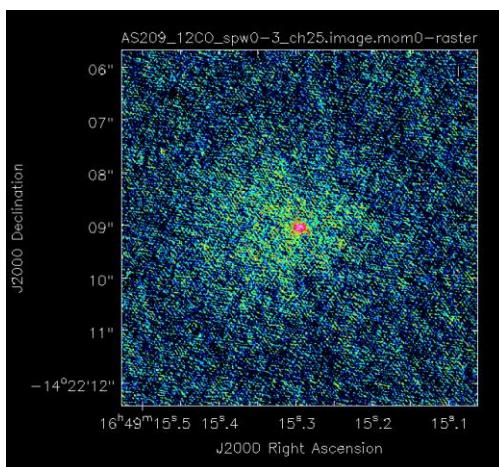
## 12CO line momen-0 map



Natural weighting



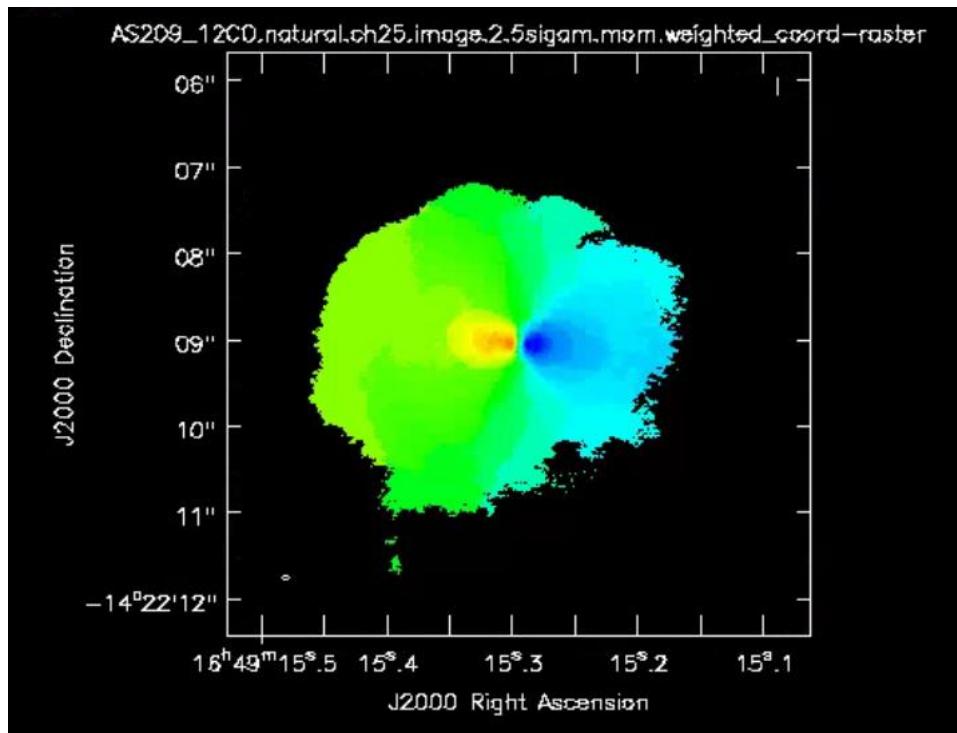
Robust weighting



Uniform weighting

Natural weighting

12CO line momen-1 map

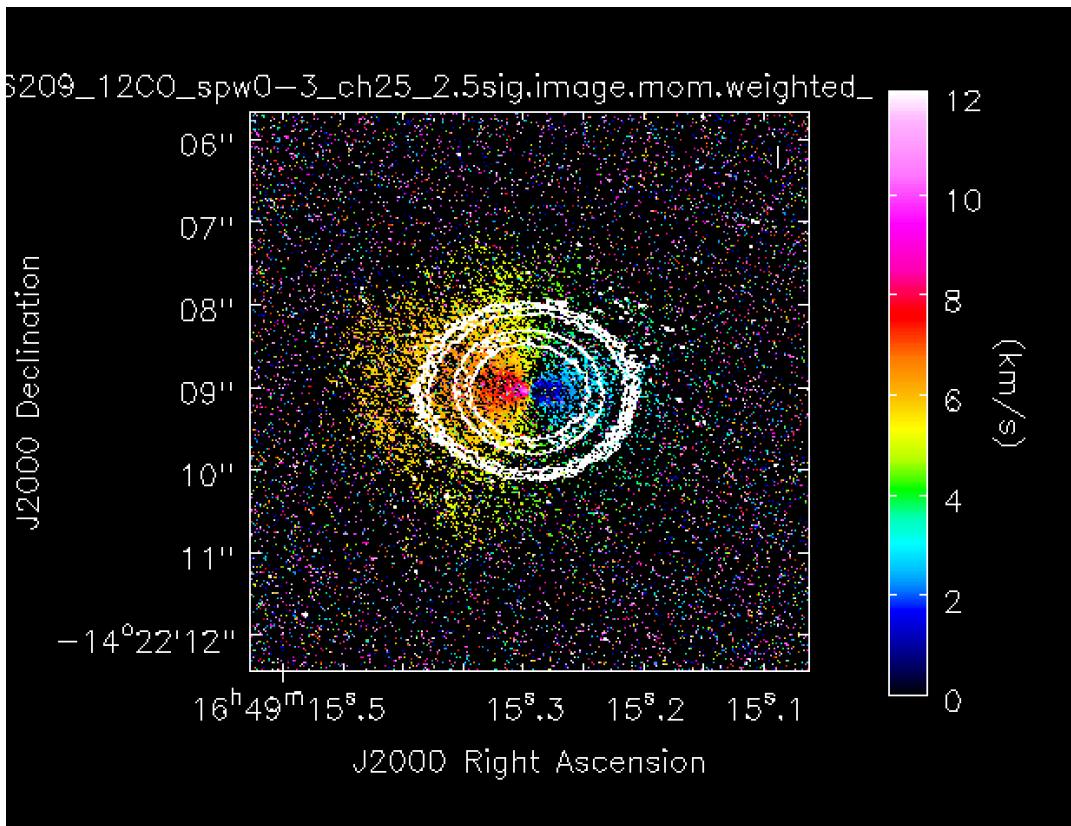


# Uniform weighting

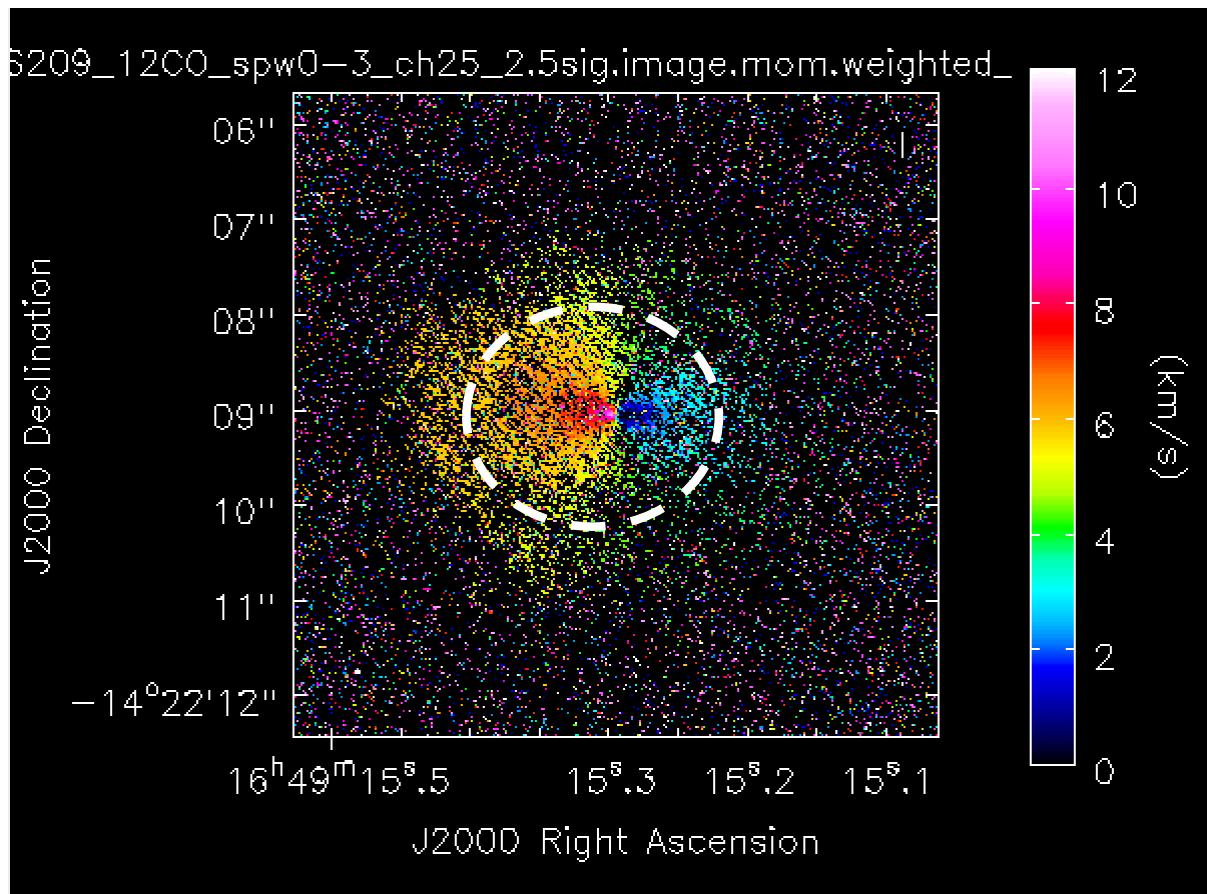
moment 1 vs continuum

moment 1  
ch2-ch23  
include pixel > 0.00375 Jy/beam (2.5sigma)

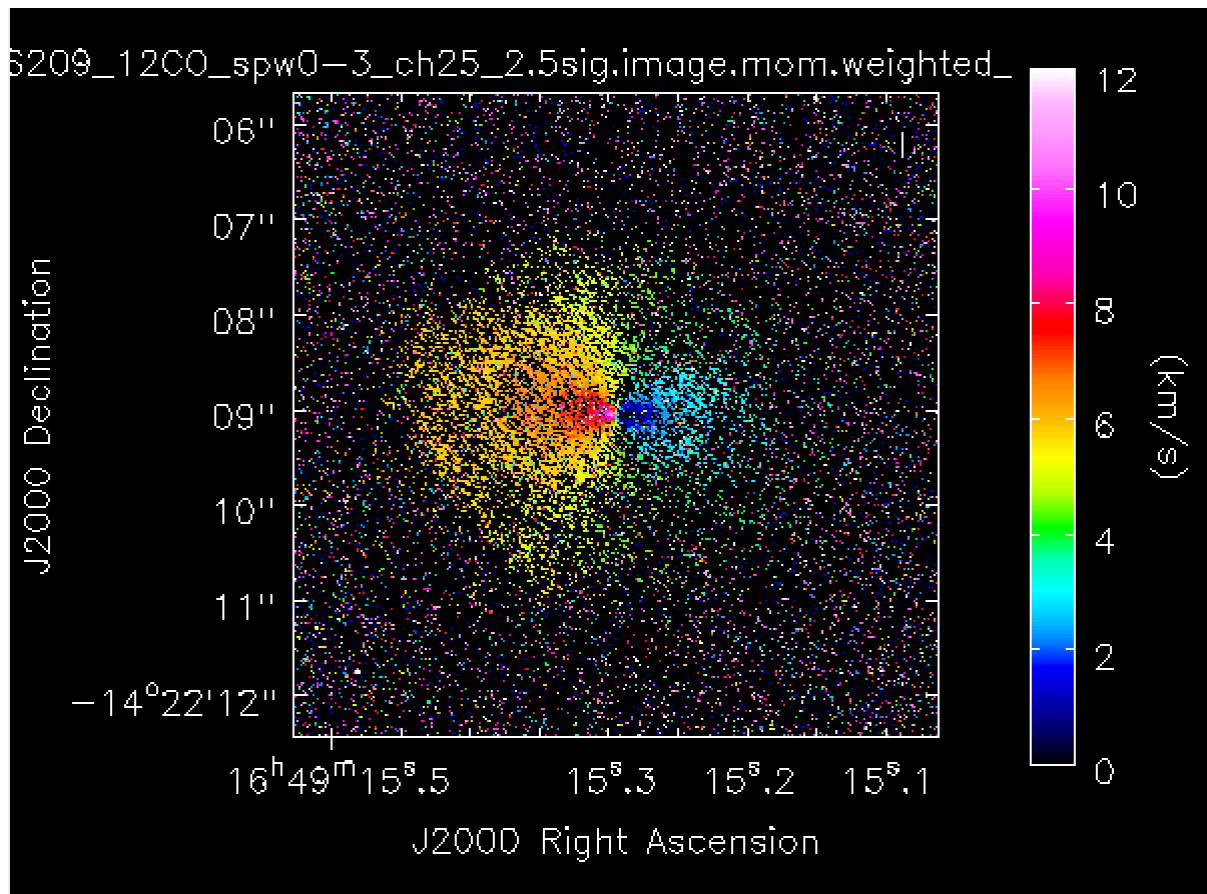
contour: dust continuum.  
3, 5, 9-sigmas  
(1sig = 1.35e-5Jy/beam)



moment 1

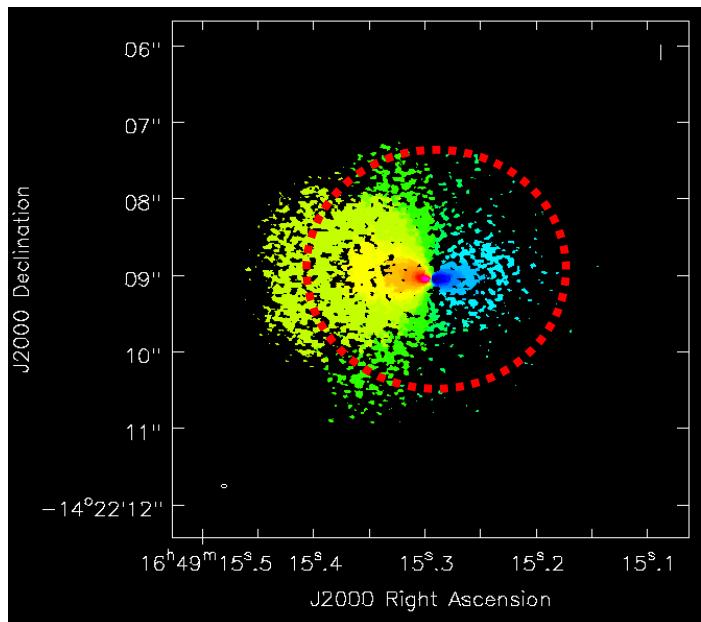


moment 1

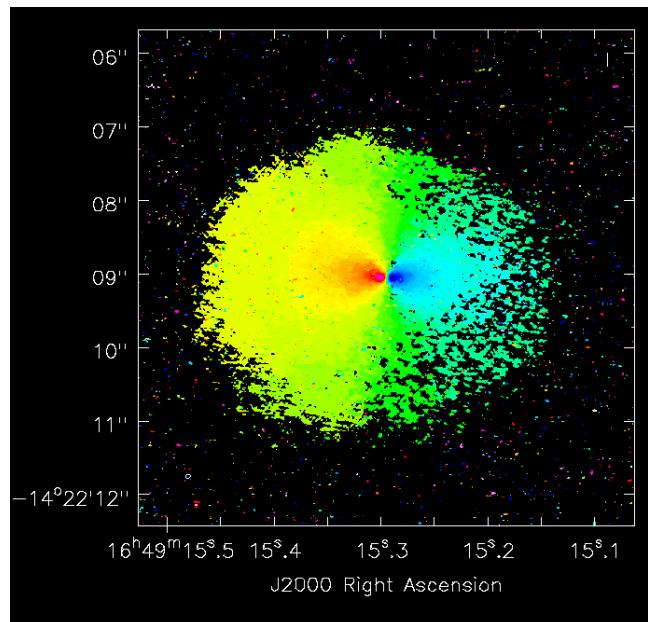


## Robust weighting

## 12CO line momen-1 map



5 sigma



2.5 sigma