

# Self-Calibration

Group1

Special thanks to

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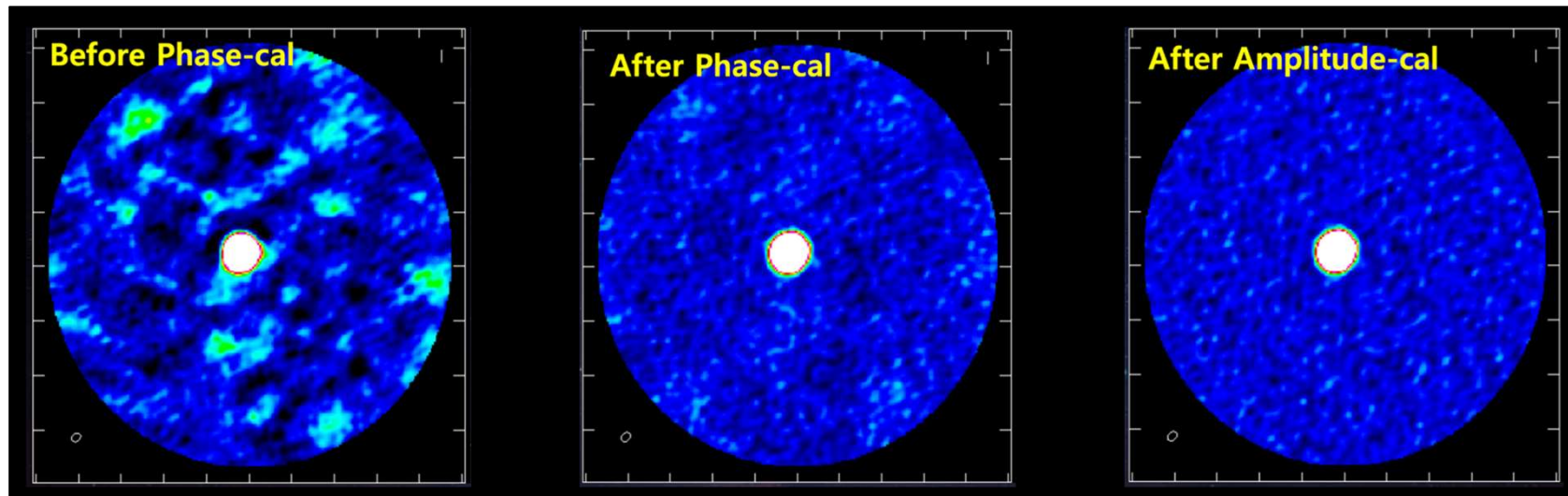
# Why do we need calibrations?

## Phase errors:

- Emission is smeared and astrometry is degraded.
- Flux is reduced and weak emission undetected
- Anti-symmetric artefacts in image

## Amplitude errors:

- Spotty or stripy emission
- Flux reduced and/or Noise increased
- Symmetric artefacts in image



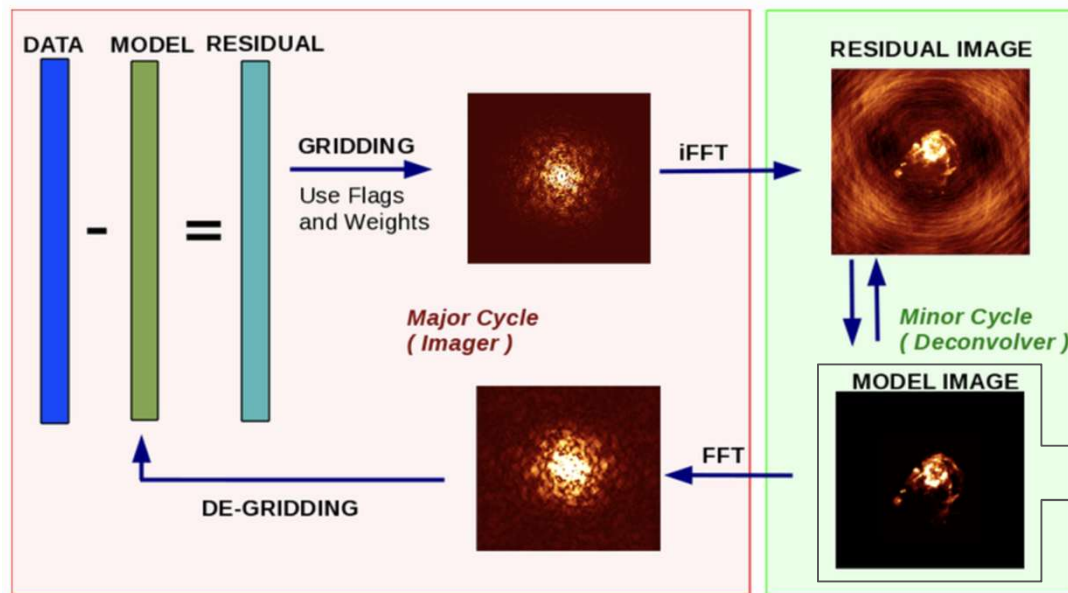
# Our Target - TW Hya

V\* TW Hya  
Type: TTauri\*  
Mag : 11.94  
Biblio: by [Ximbard](#)

$\theta \sim 5.6^\circ$

Atmospheric difference of position and time  
between Phase-calibrator and targets ( < a few degrees and times)

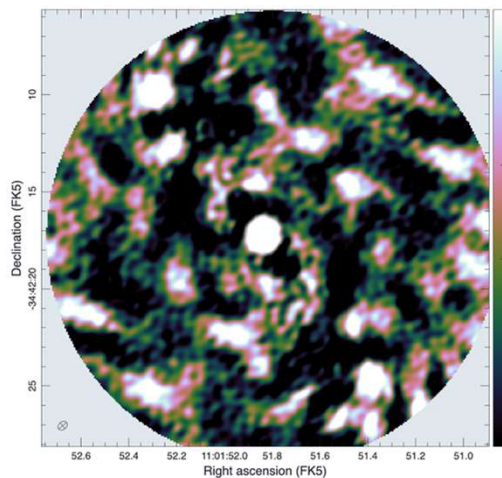
# What is self-calibration?



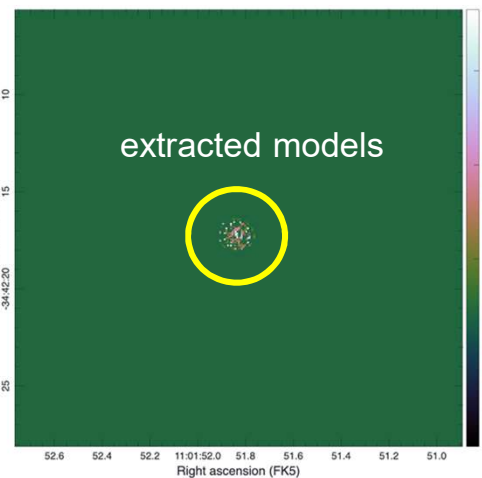
We could utilize a better calibrator  
→ which is **the target!**

# What is self-calibration?

Dirty Map



Model Map



$$V_{ij}^{obs} = \mathcal{G} V_{ij}^{true}$$

$$\left| V_{ij}^{obs} - \mathcal{G}_{ij}(t) V^{model} \right|^2$$

minimize!

→ *gaincal*

$$\frac{V_{ij}^{obs}}{\mathcal{G}} \approx V_{ij}^{true}$$

→ *applycal*

postulate the cleaned result as an **analog** of the phase calibrator at the source position!

# What is self-calibration?

## Sanity Check

inspection of sufficient S/N ratio must be preceded

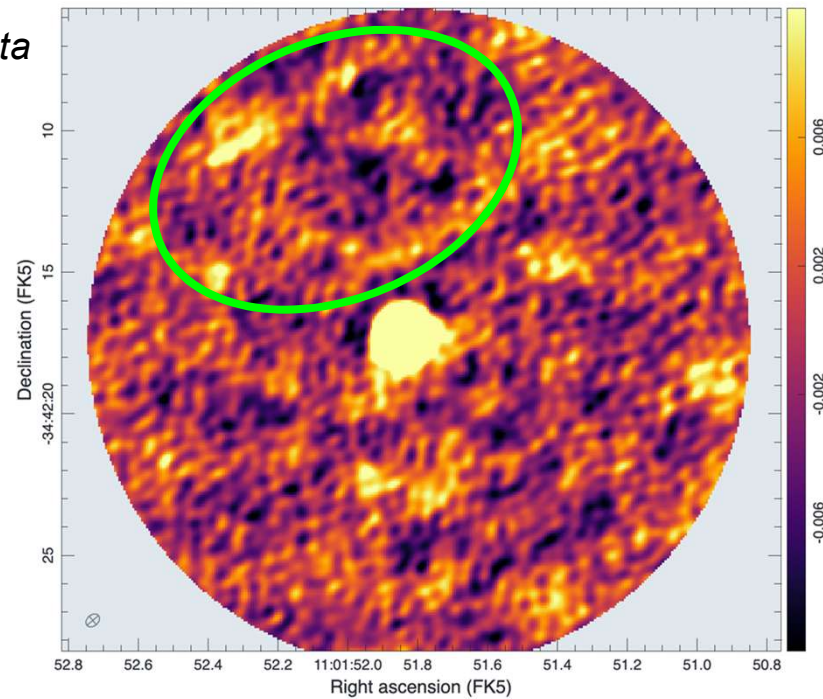
e.g.,  $S/N > 20$  for  $\sim 25$  antennas

- strong outlier included in the model
- severe variability of the data
- over-calibrated image

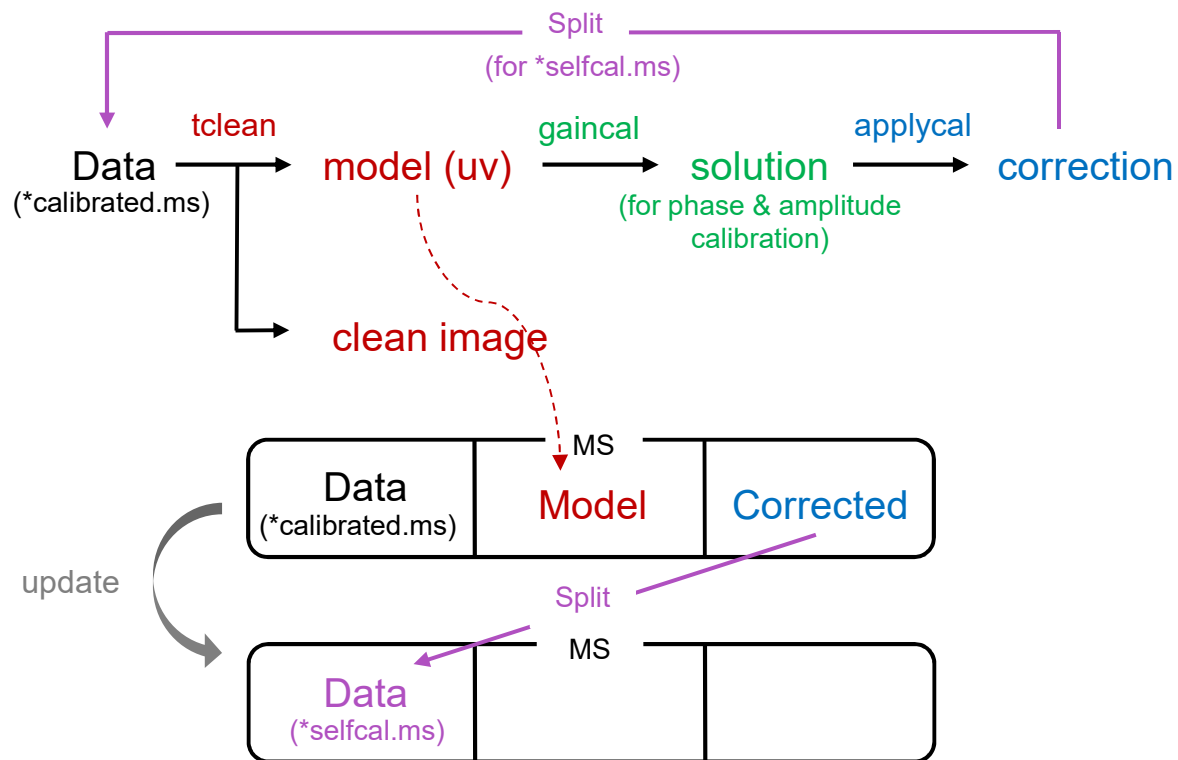
# What is self-calibration?

**significant** residuals still exist after phase calibration (with J1037-265)

*after cleaning the TW\_Hya data*

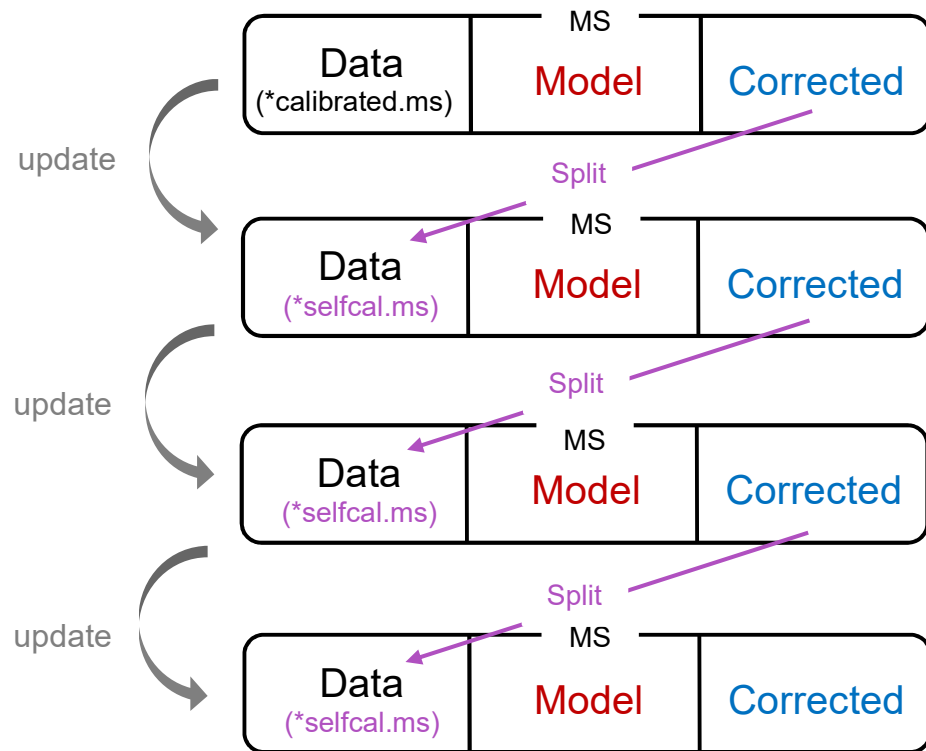


# Self-calibration process





# Self-calibration process



clean image  
(\*calibrated.image)

clean image

Self calibration round 1  
(**phase** calibration)

clean image

Self calibration round 2  
(**phase** calibration)

clean image

Self calibration round 3  
(**phase** calibration)

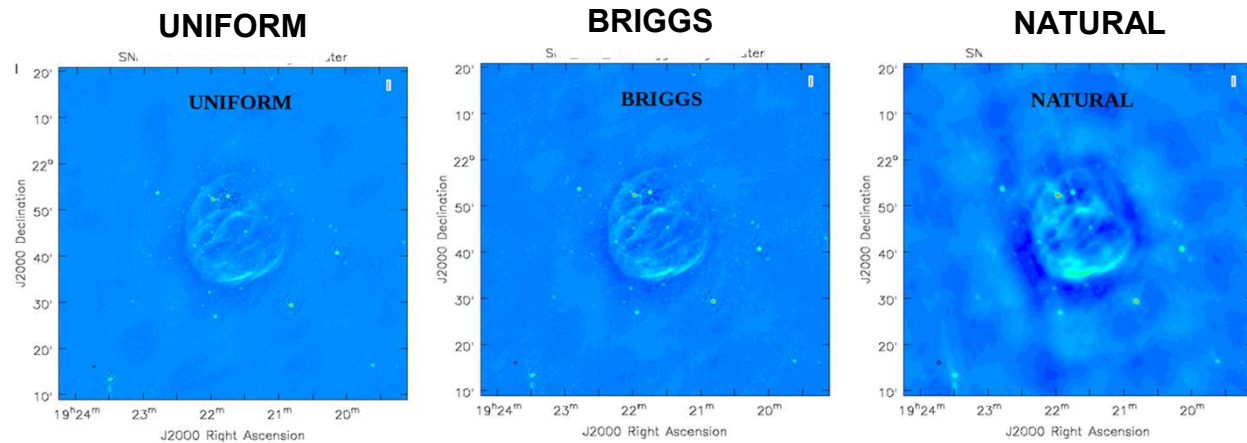
clean image

Self calibration round 4  
(**phase + amplitude** calibration)

# Adjusting parameter

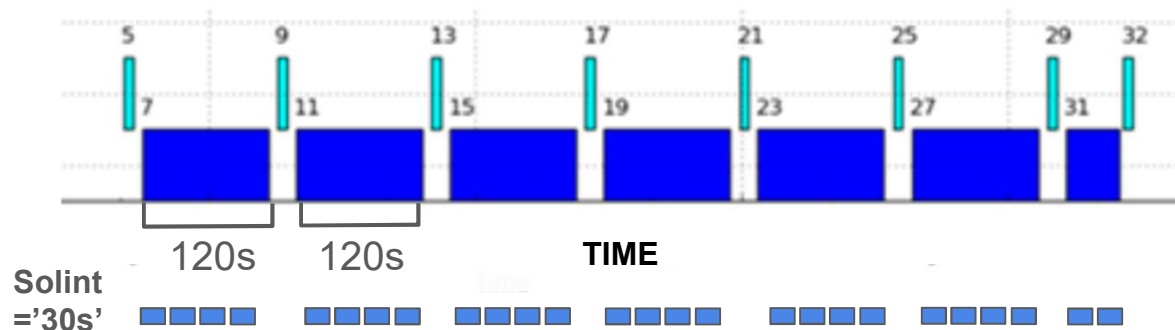
## [Adjusting weighting]

- case 1(briggs)
- case 2(uniform)
- case 3(natural)

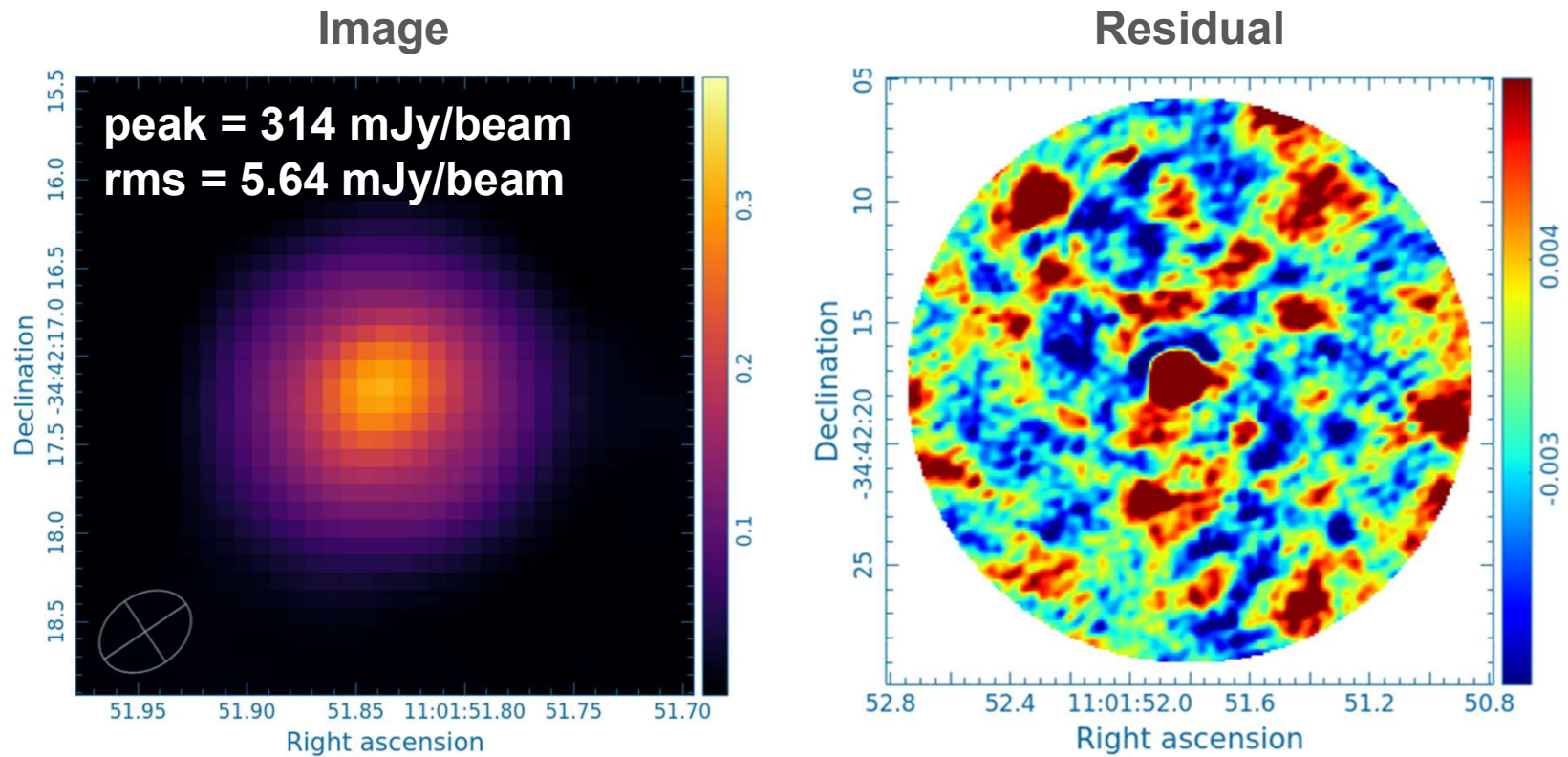


## [Adjusting solint]

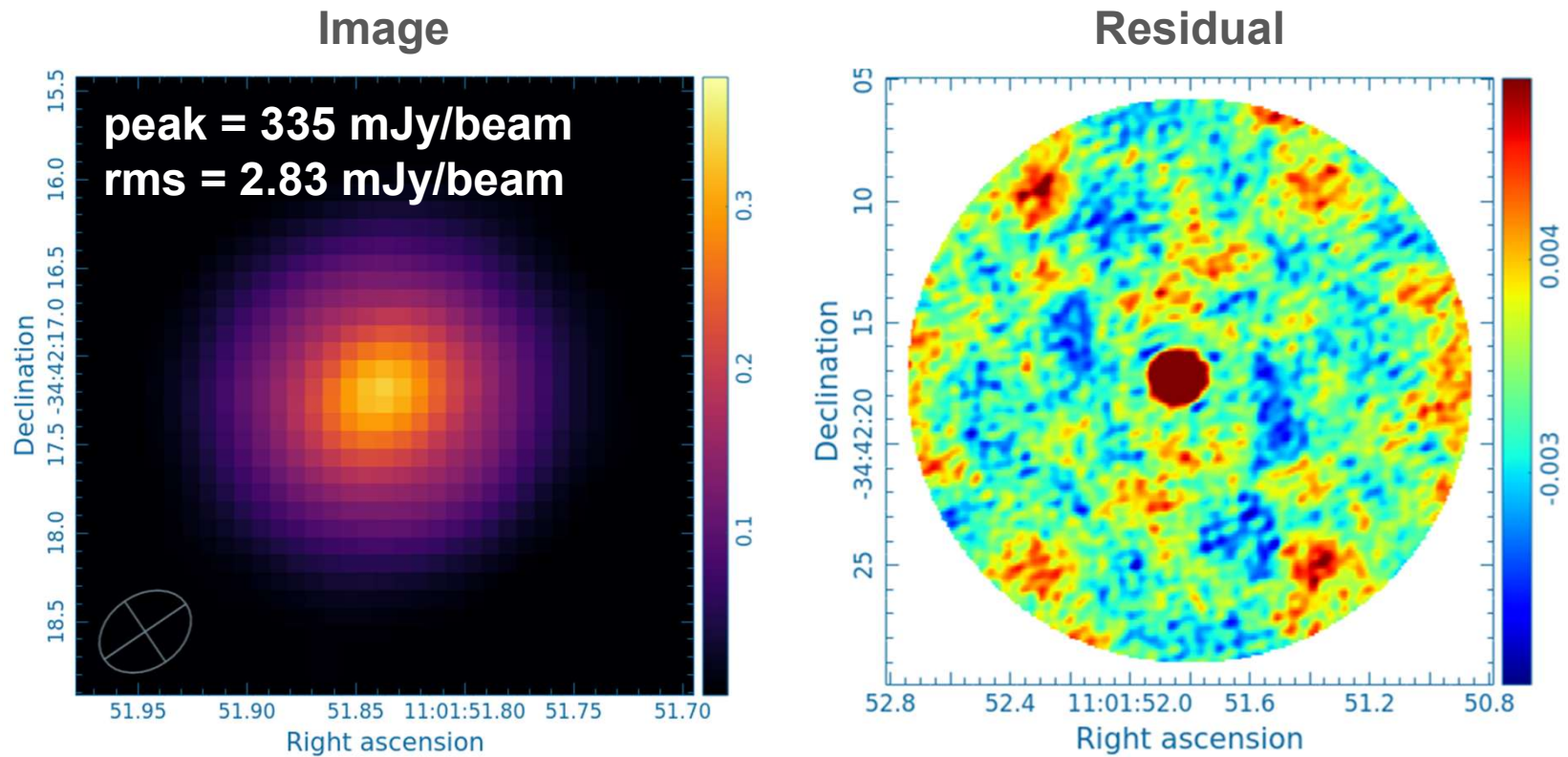
- what is solint
- solint result



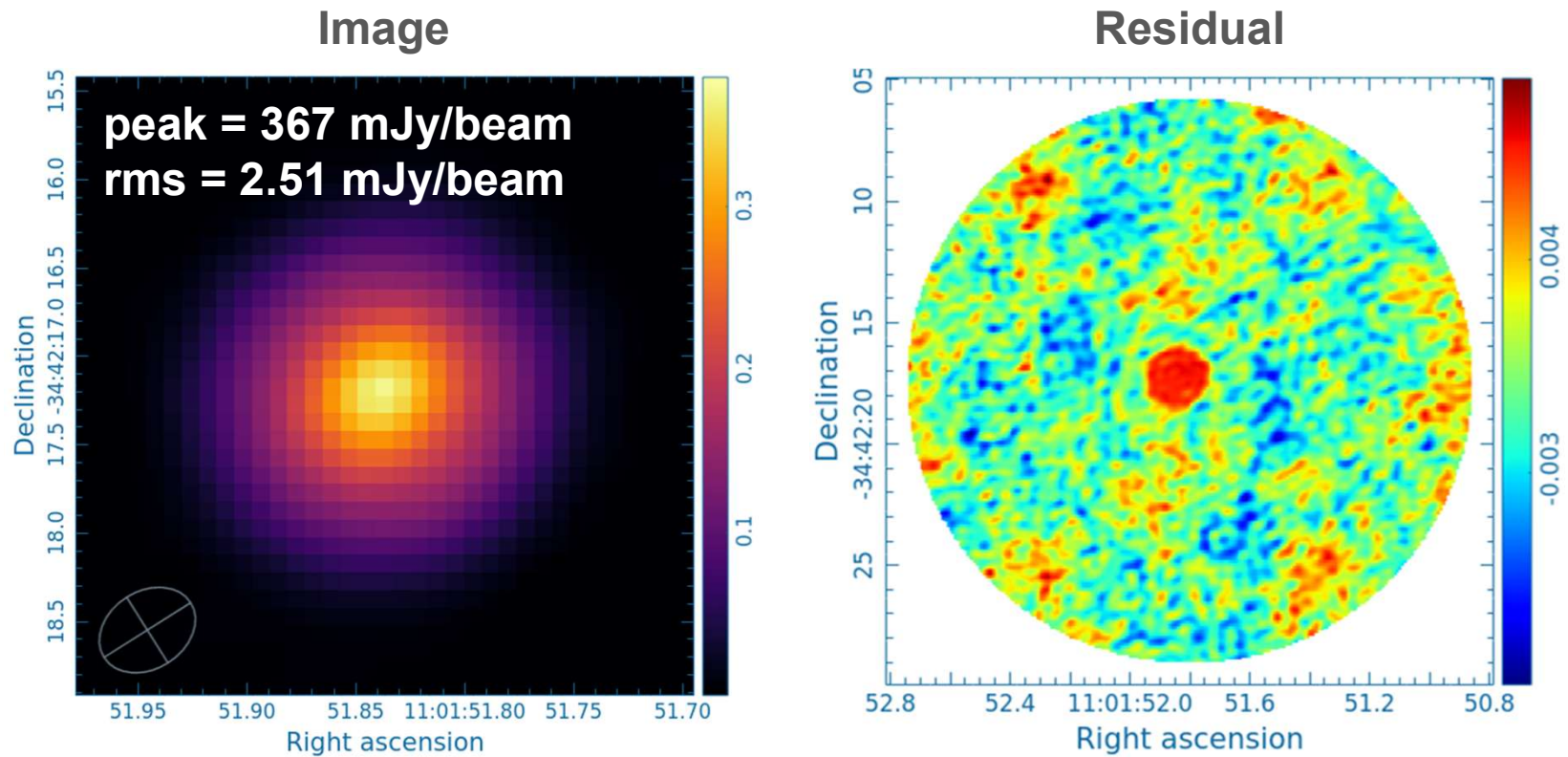
# Case 1: weighting “briggs” (robust = 0.5)



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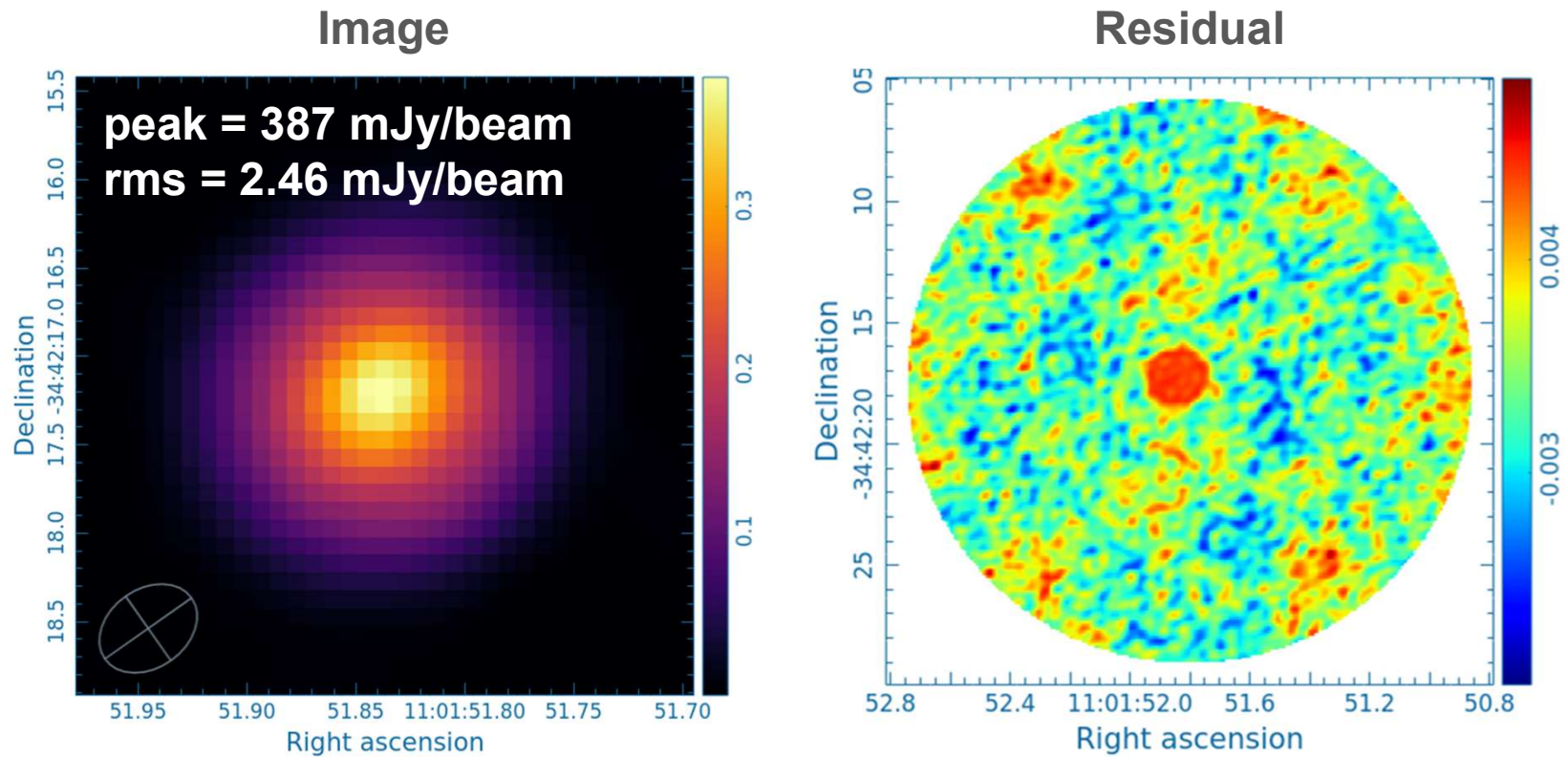


# Case 1: weighting “briggs” (robust = 0.5)

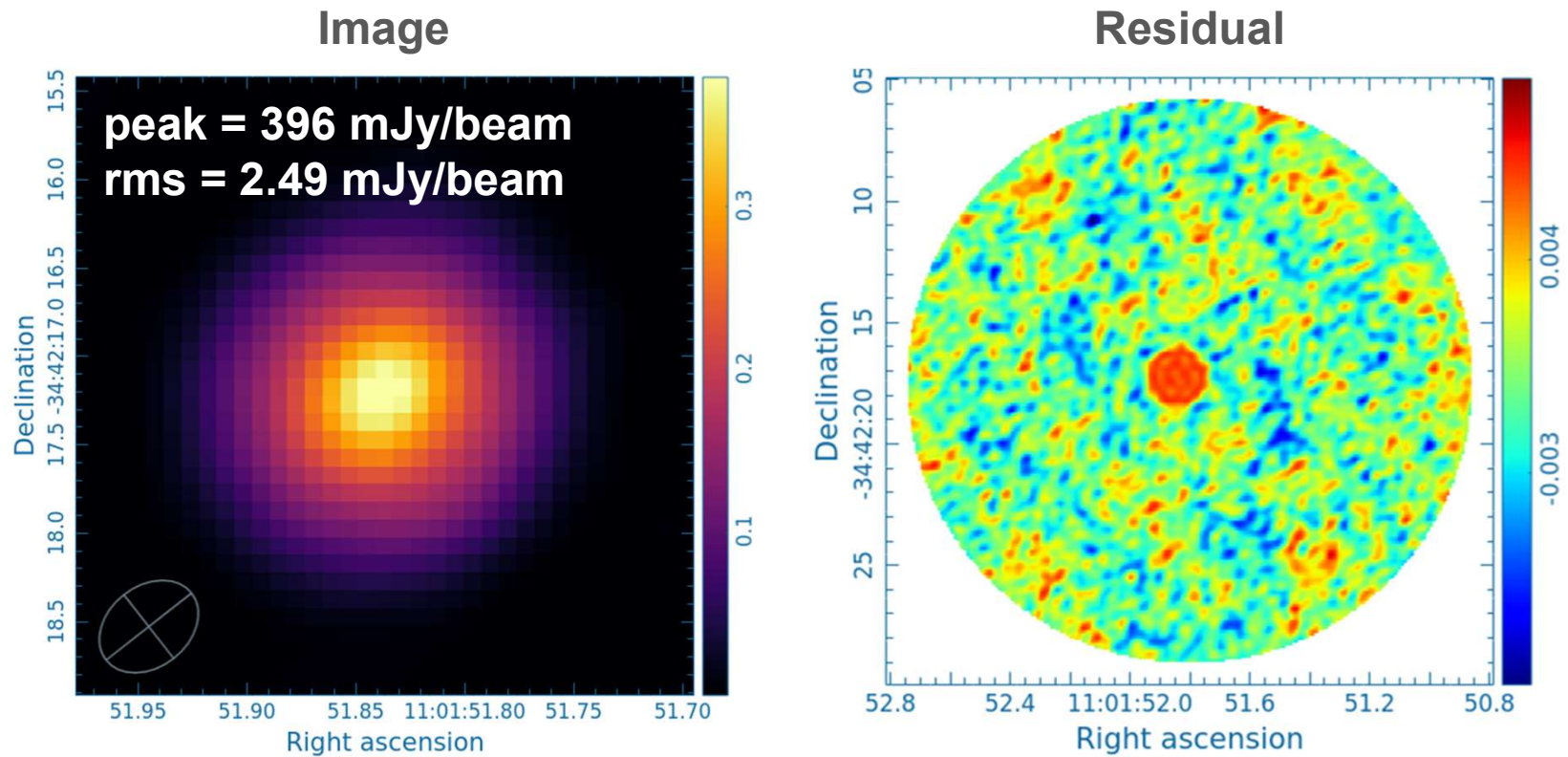




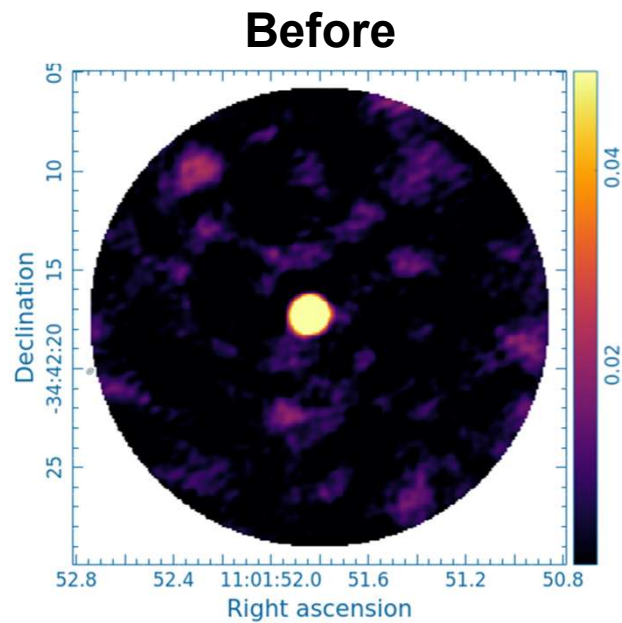
# Case 1: weighting “briggs” (robust = 0.5)



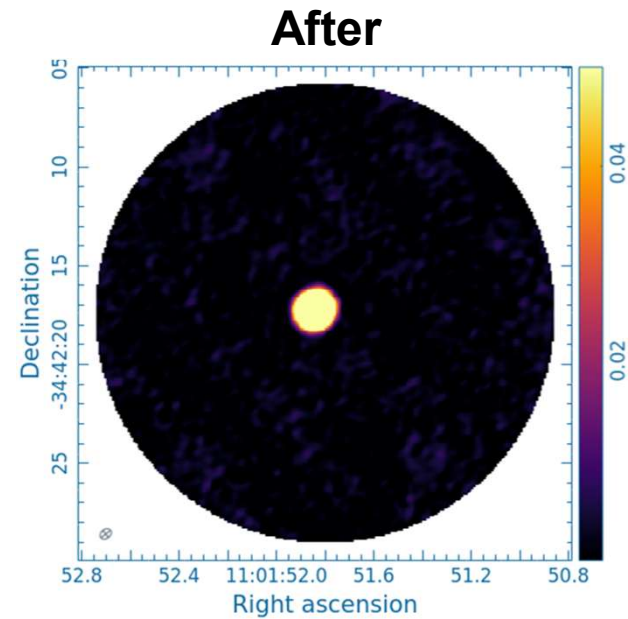
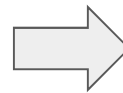
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SNR	56
beam	0.560" * 0.416"



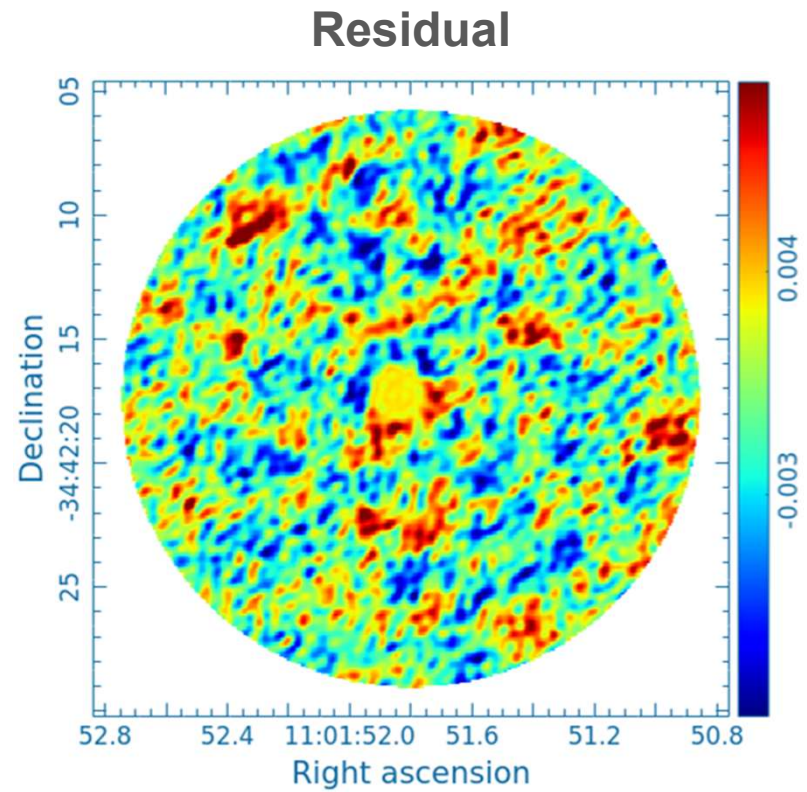
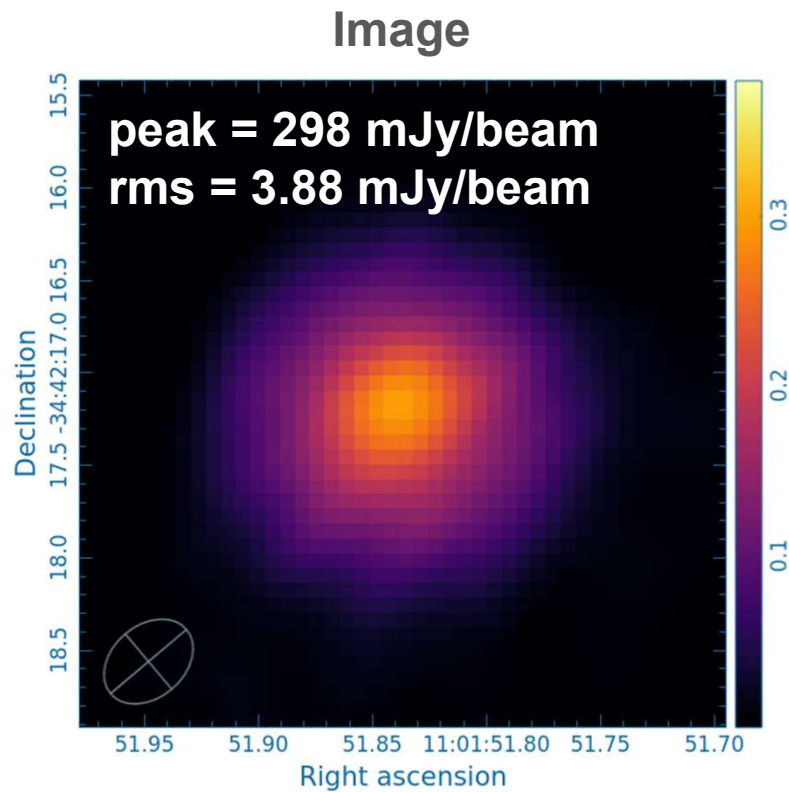
SNR	159
beam	0.614" * 0.457"

**+184%**

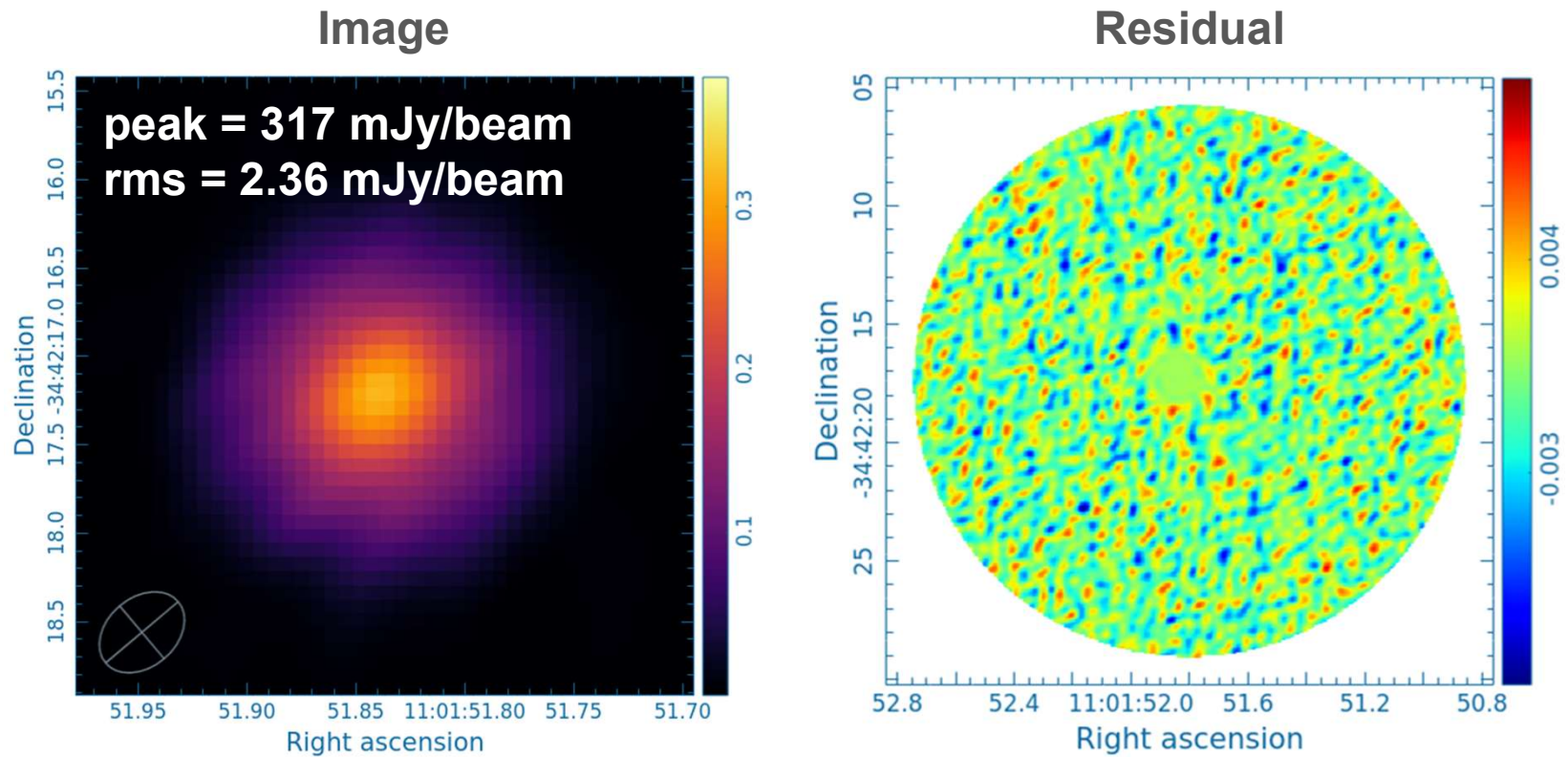
**+9.64%**



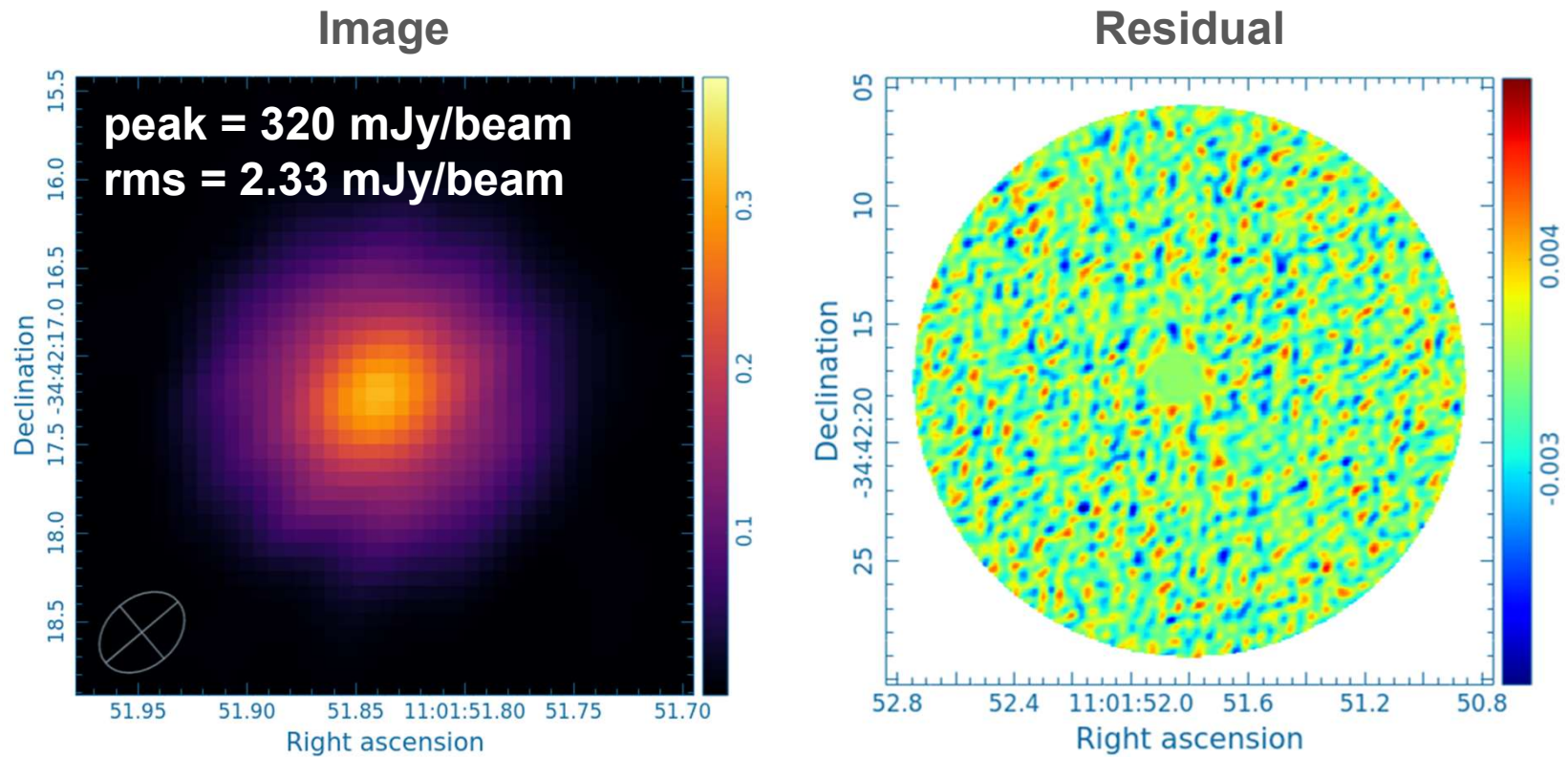
## Case 2: weighting “uniform”



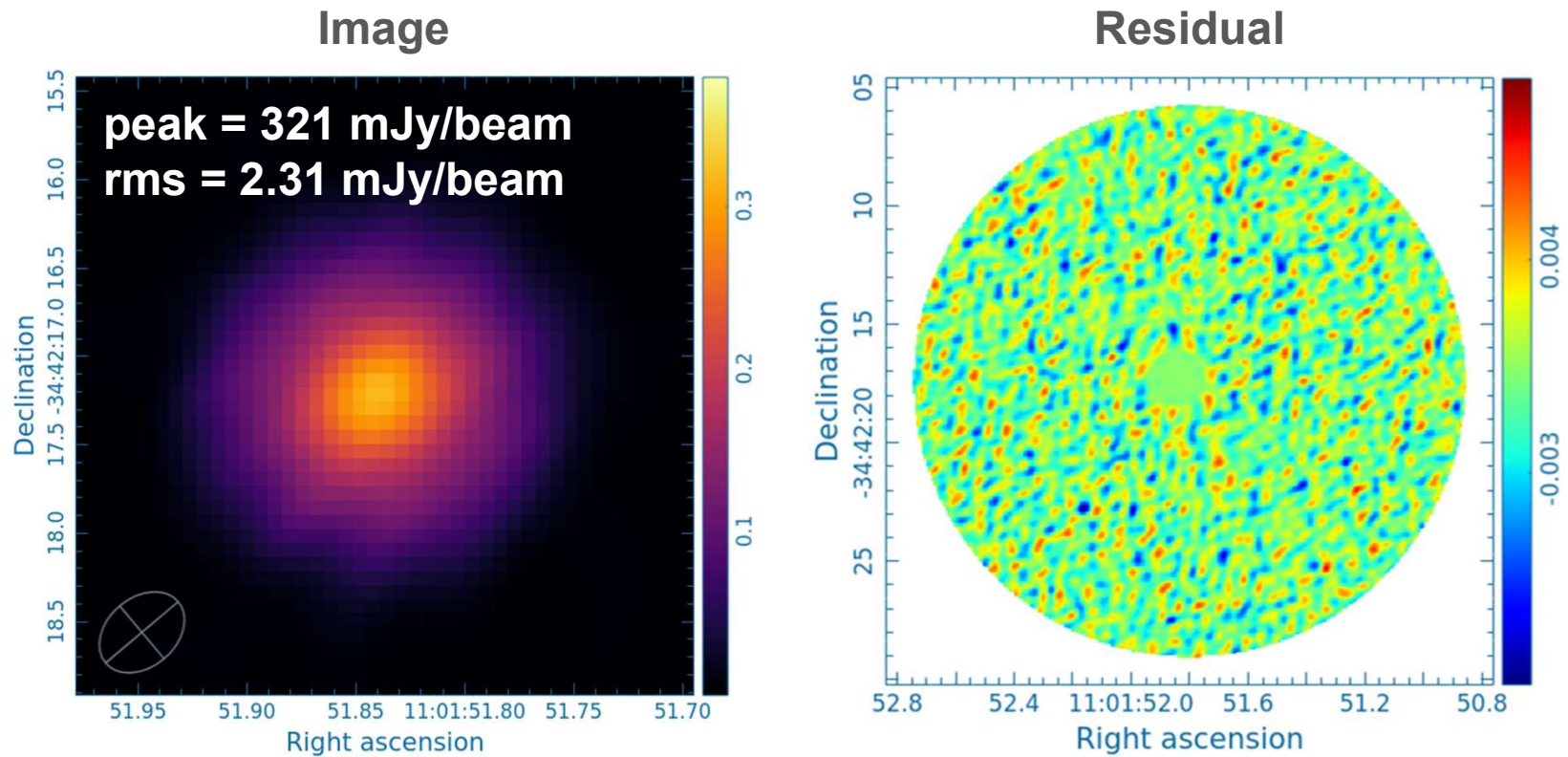
## Case 2: weighting “uniform”



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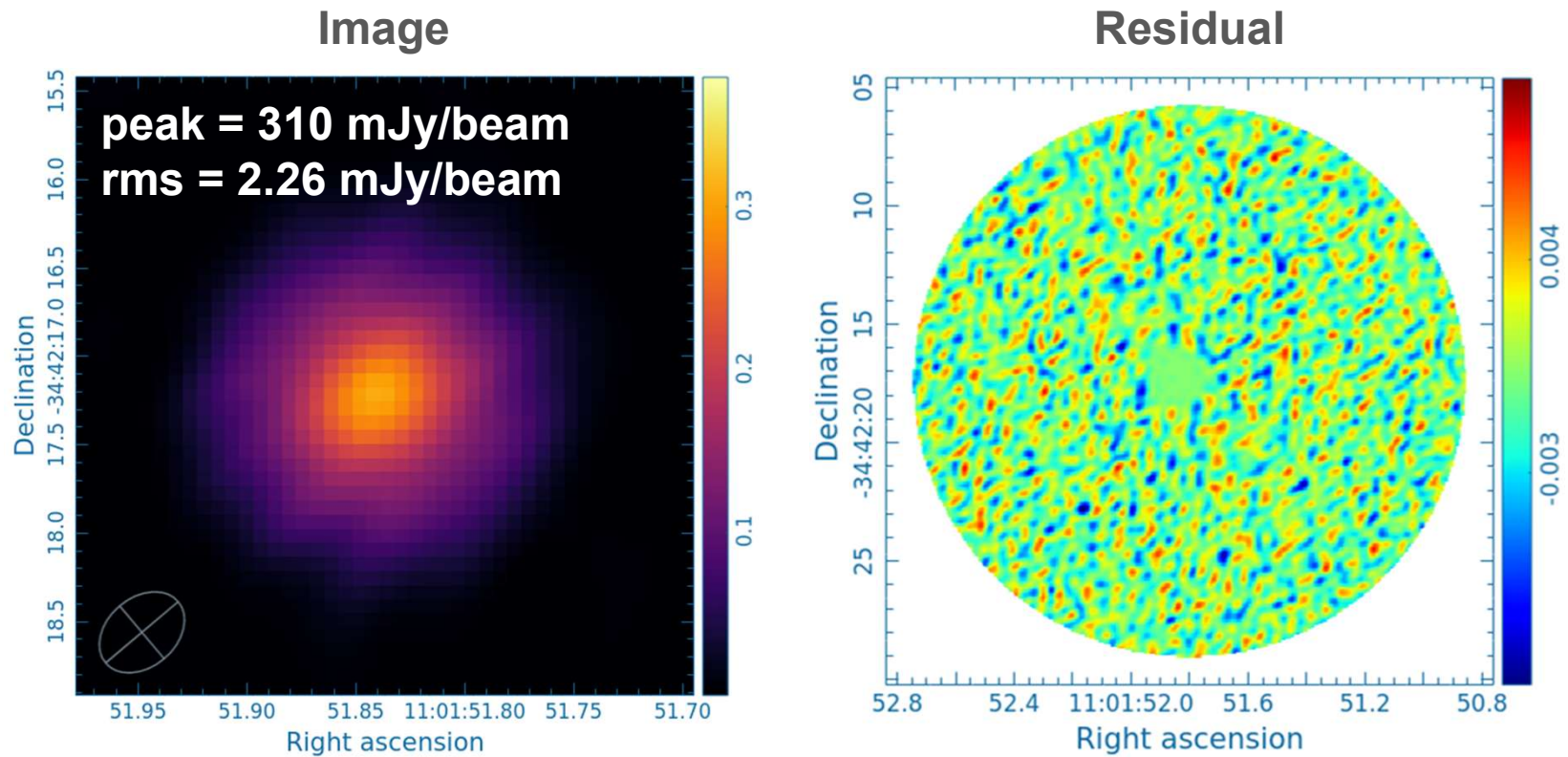


## Case 2: weighting “uniform”

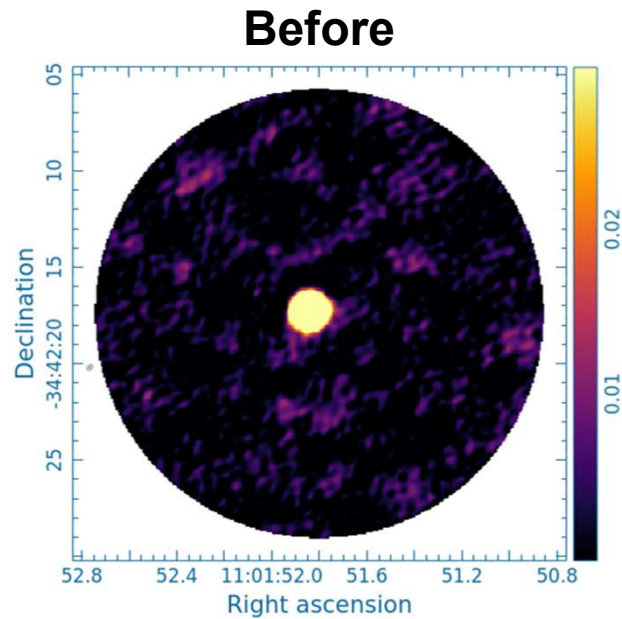




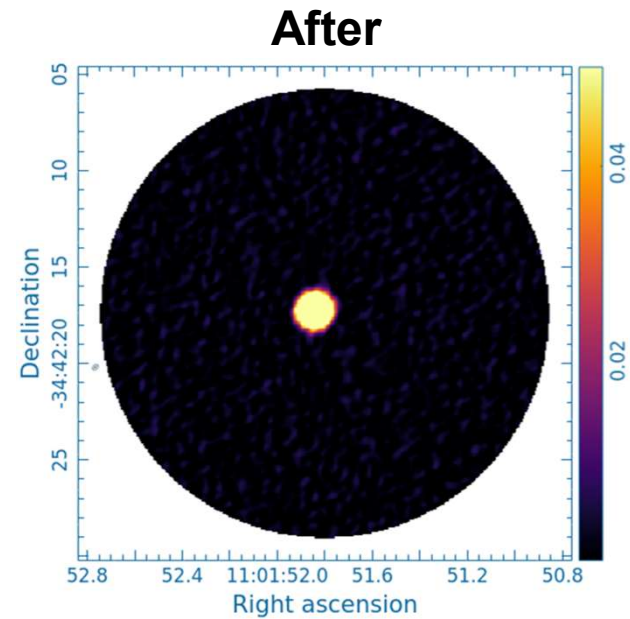
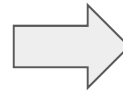
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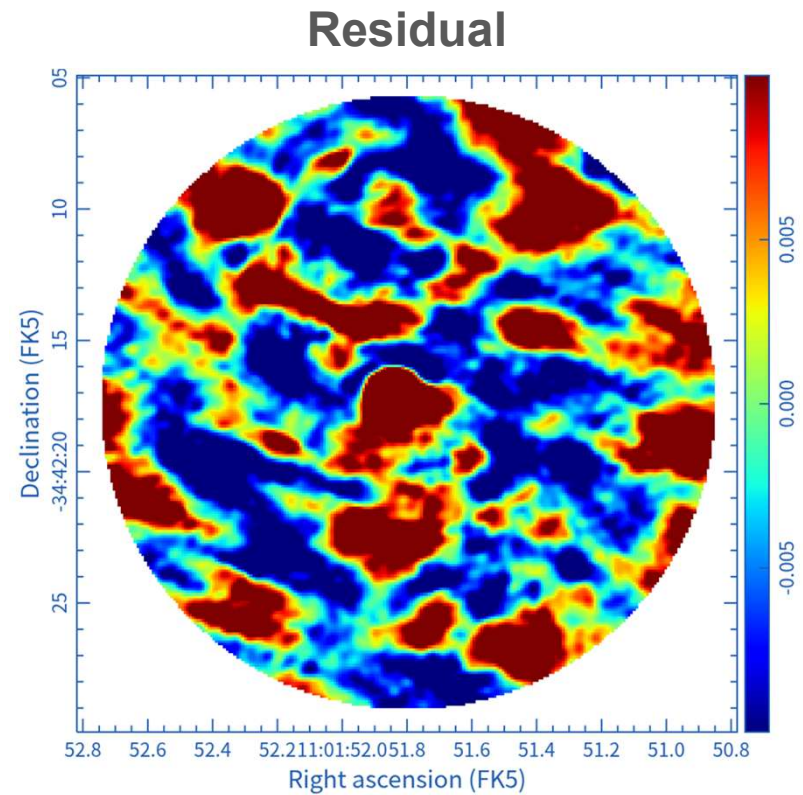
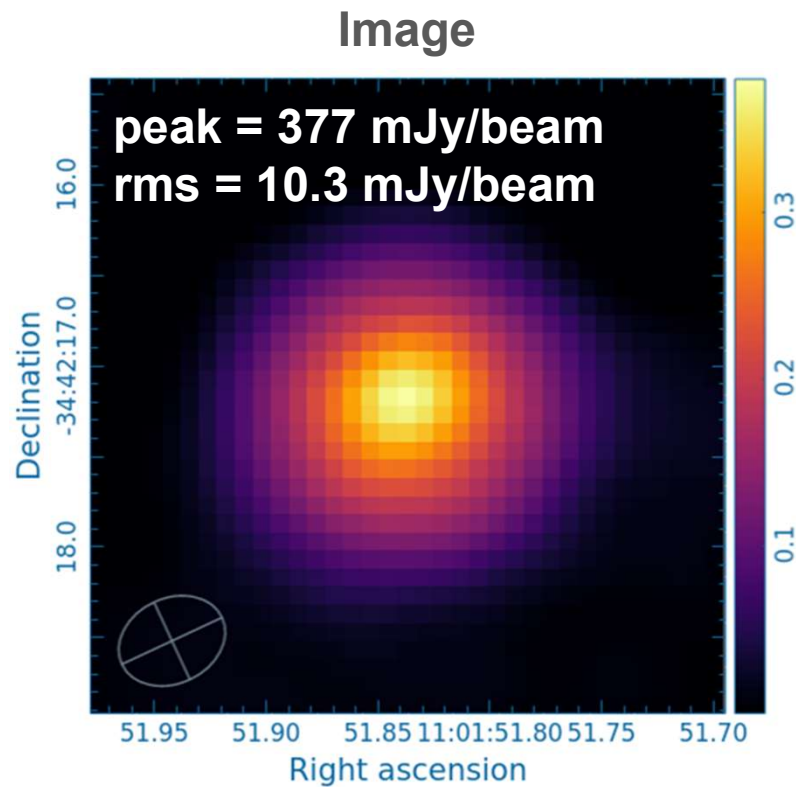
SNR	77
beam	0.537" * 0.390"



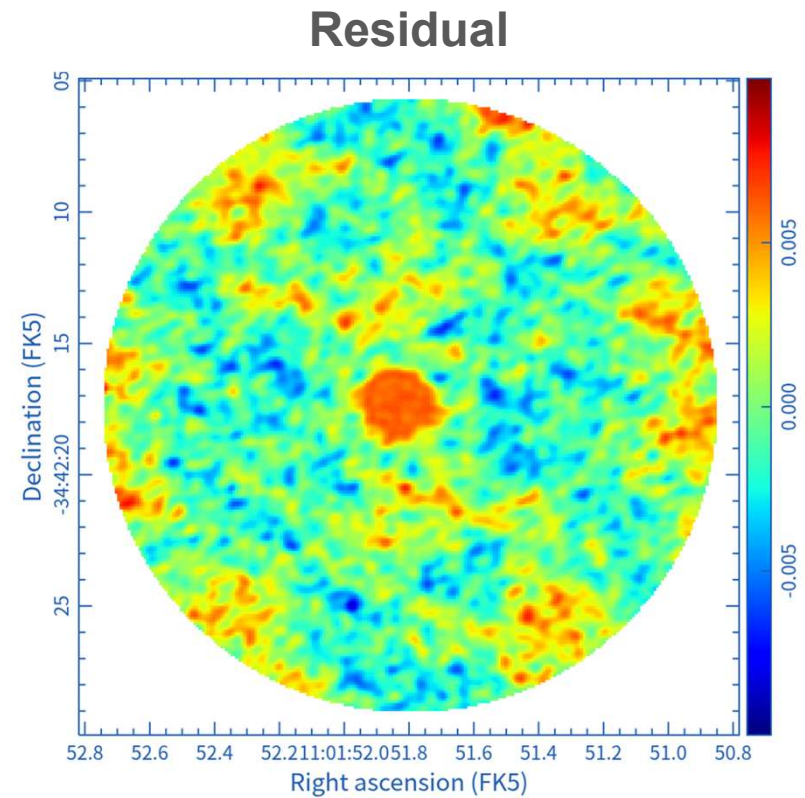
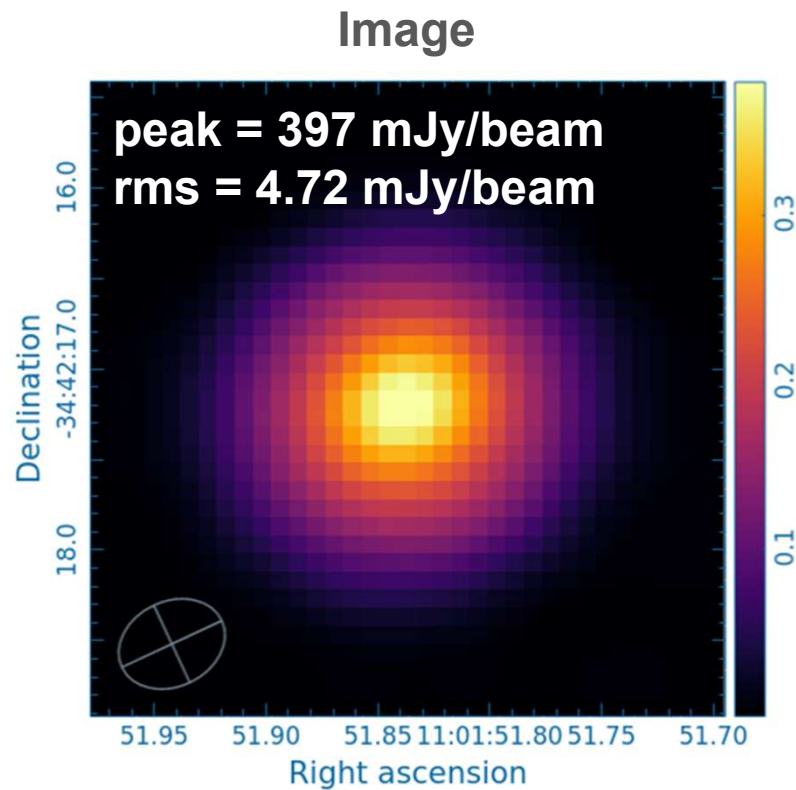
SNR	137
beam	0.537" * 0.390"

**+77.9%**  
**0%**

## Case 3: weighting “natural”

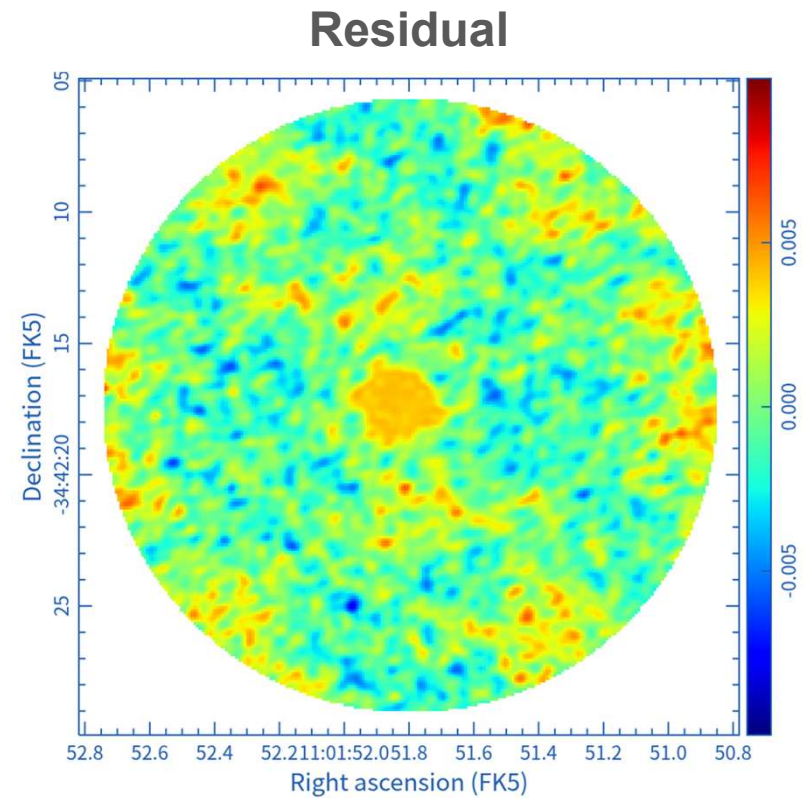
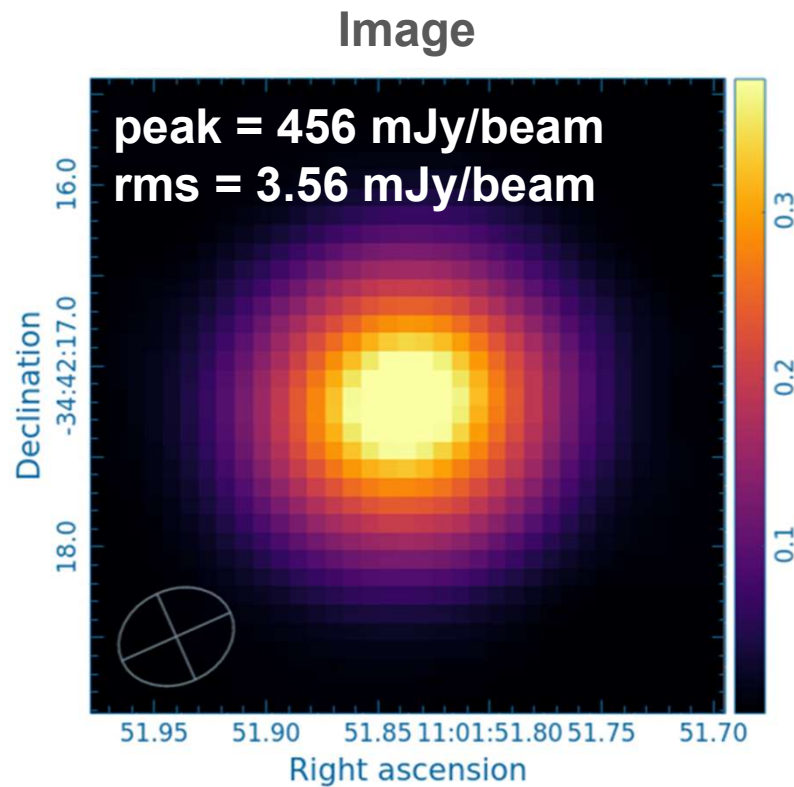


## Case 3: weighting “natural”

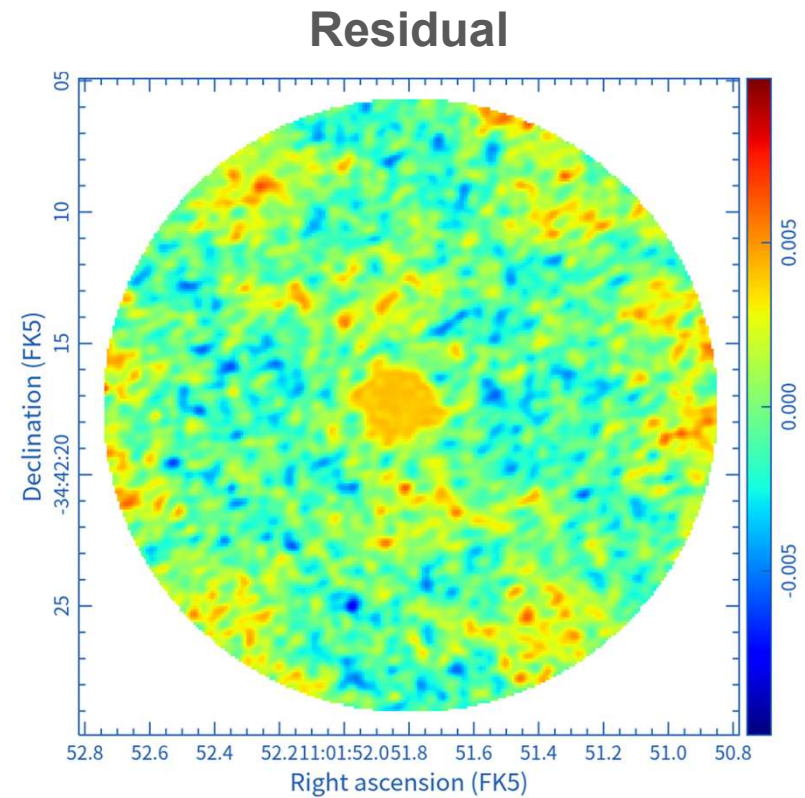
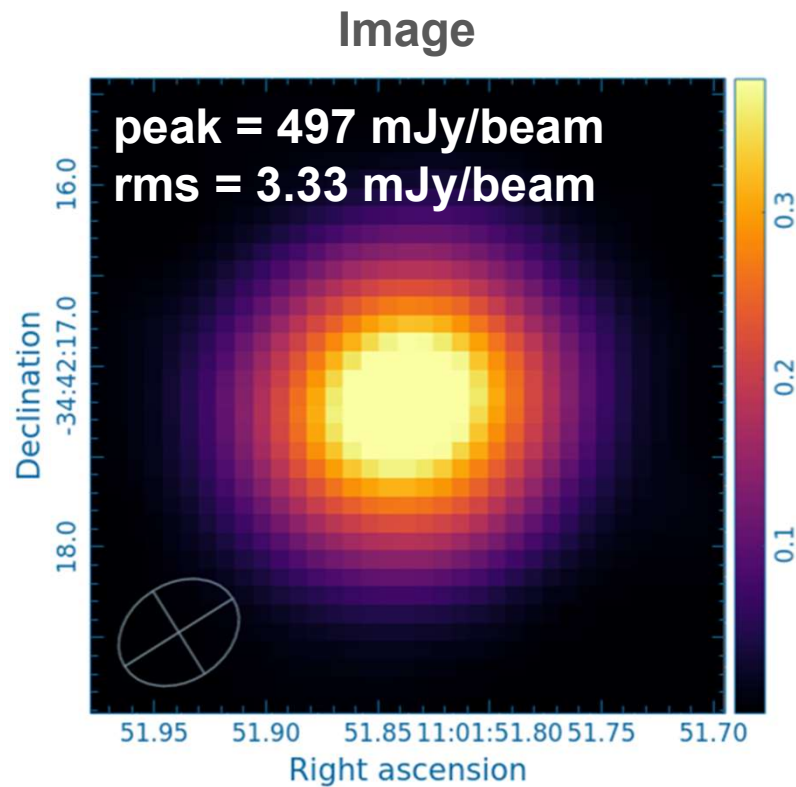




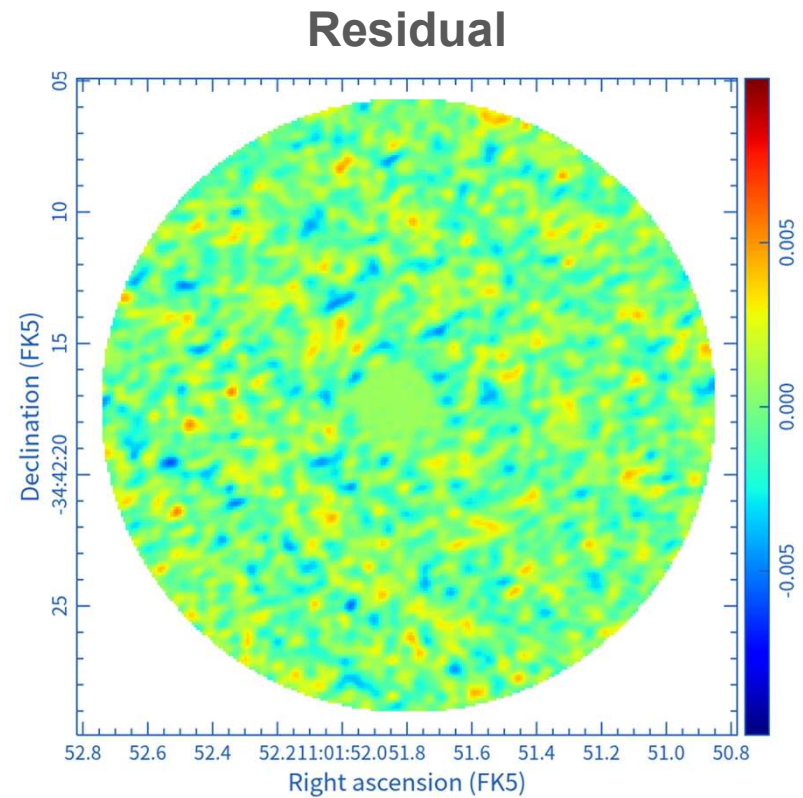
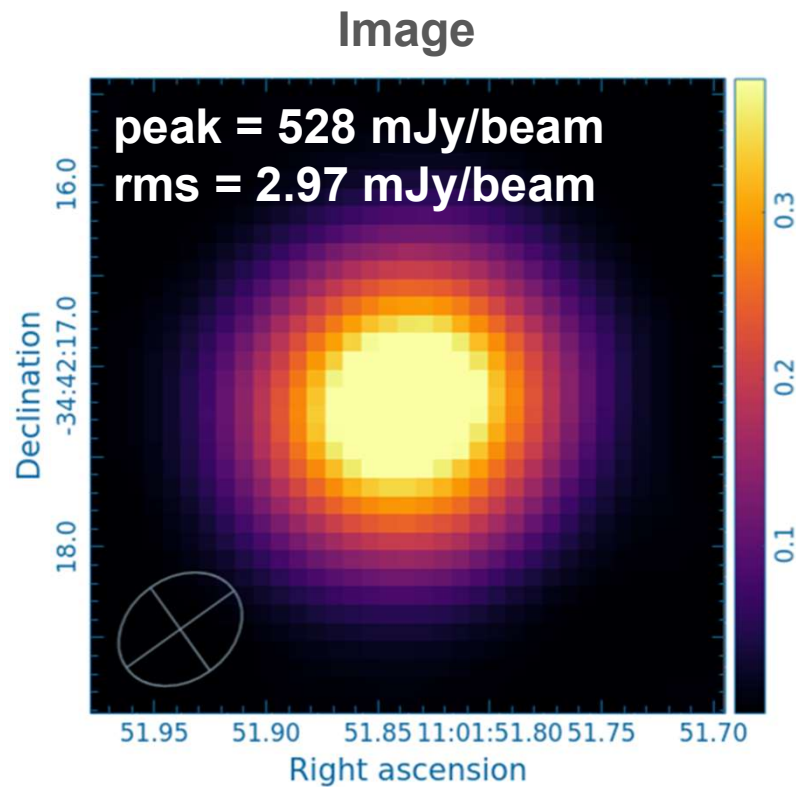
## Case 3: weighting “natural”



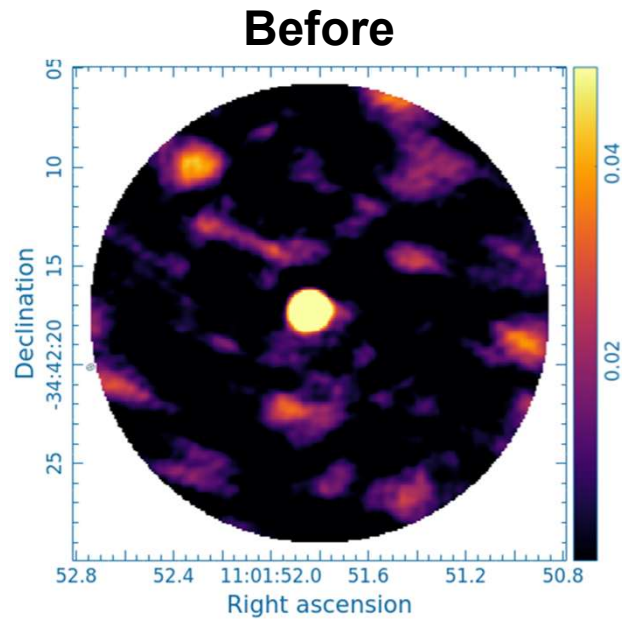
## Case 3: weighting “natural”



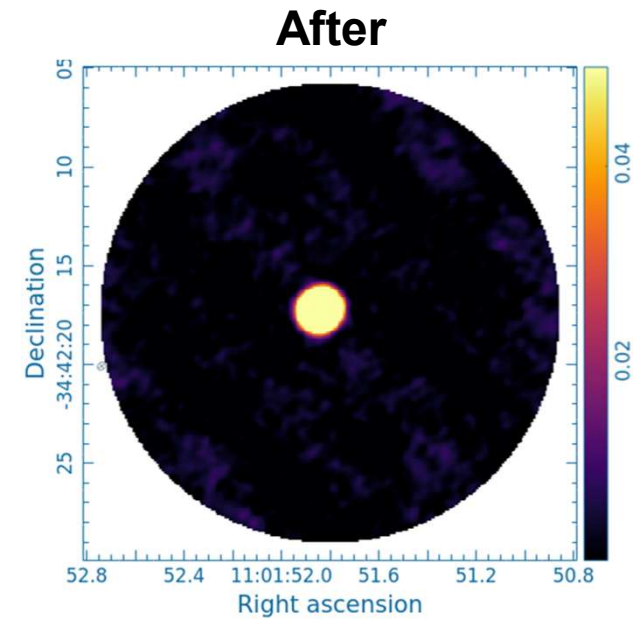
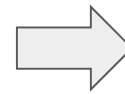
## Case 3: weighting “natural”



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SNR	36.6
beam	0.609''

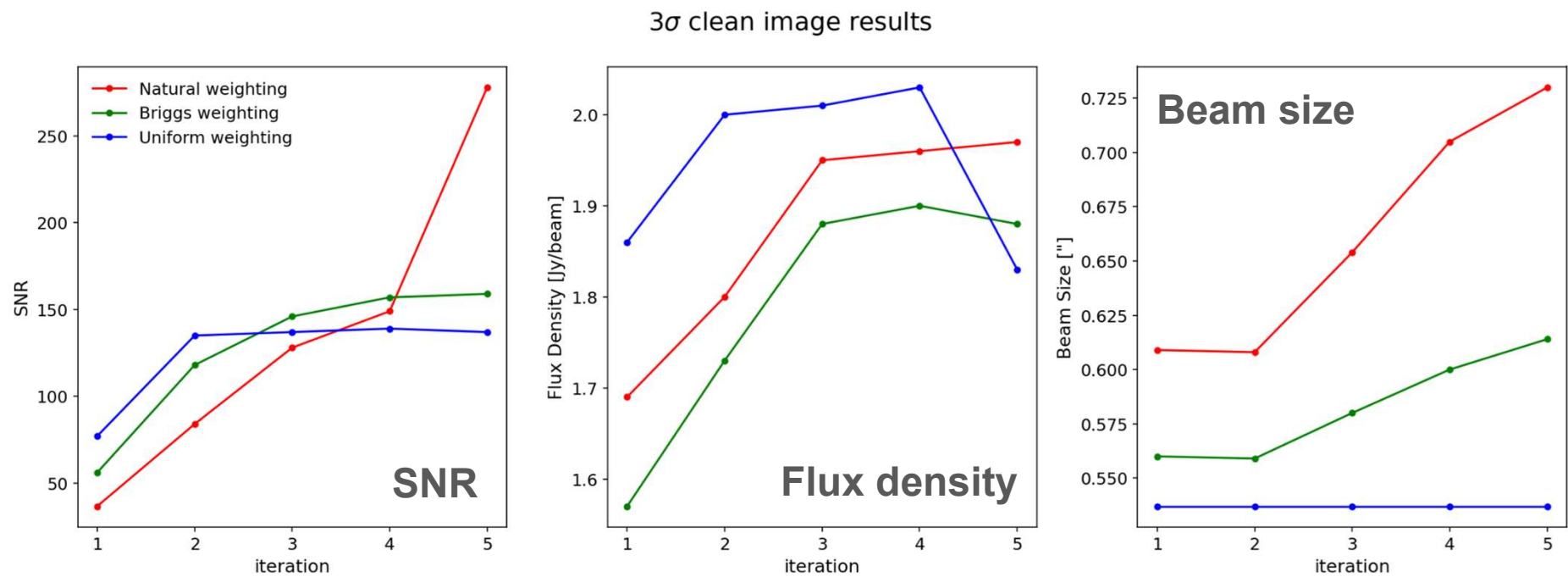


SNR	278
beam	0.730''

**+659%**

**+19.9%**

# Comparison of three cases





# Adjusting Solution Interval

## gaincal

```
gaincal(vis, caltable="", field="", spw="", intent="", selectdata=True, timerange="", uvrange="", antenna="", scan="", observation="", msselect="",  
solint='inf', combine="", preavg=-1.0, refant="", refantmode='flex', minblperant=4, minsnr=3.0, solnorm=False, normtype='mean', gaintype='G',  
smodel="", calmode='ap', solmode="", rmsthresh="", corrdepflags=False, append=False, splinetime=3600.0, npointaver=3, phasewrap=180.0,  
docallib=False, callib="", gaintable="", gainfield="", interp="", spwmap="", parang=False) \[source\]
```

vis, caltable, field, spw, intent, selectdata, solint, combine,  
preavg, refant, refantmode, minblperant antenna, minsnr,  
solnorm, gaintype ...

# Adjusting Solution Interval

## gaincal

```
gaincal(vis, caltable="", field="", spw="", intent="", selectdata=True, timerange="", uvrange="", antenna="", scan="", observation="", msselect="",  
solint='inf', combine="" preavg=-1.0, refant="", refantmode='flex', minblperant=4, minsnr=3.0, solnorm=False, normtype='mean', gaintype='G',  
smodel="", calmode='ap', solmode="", rmsthresh="", corrdepflags=False, append=False, splinetime=3600.0, npointaver=3, phasewrap=180.0,  
docallib=False, callib="", gaintable="", gainfield="", interp="", spwmap="", parang=False) \[source\]
```

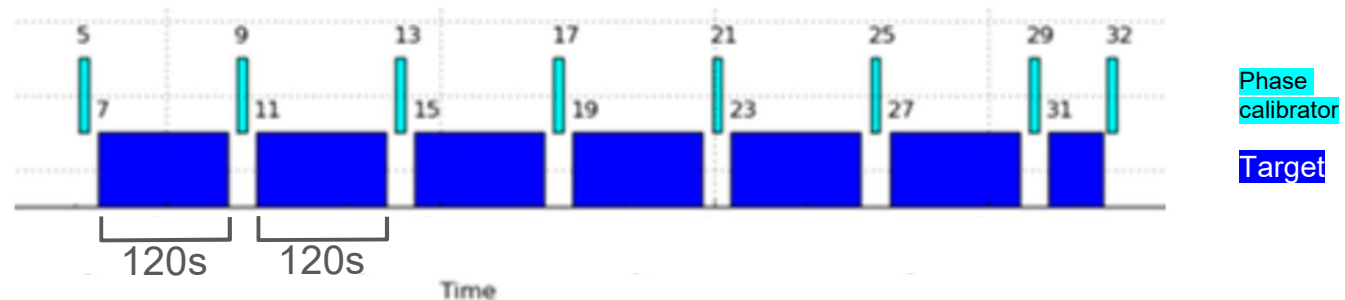
vis, caltable, field, spw, intent, selectdata, **solint**, **combine**,  
preavg, refant, refantmode, minblperant antenna, minsnr,  
solnorm, gaintype ...

# Adjusting Solution Interval

On-source

**solint** defines **how frequently** gain solutions (phase or amplitude) are calculated in time.

**combine** controls **which axes of the data are grouped together** when calculating solutions.



e.g.

Combine	Solint	
=''	='30s'	
=''	='inf'	
='scan'	='240s'	
='scan'	='inf'	

solution X



# Adjusting Solution Interval

**solint** defines **how frequently** gain solutions (phase or amplitude) are calculated in time.

**combine** controls **which axes of the data are grouped together** when calculating solutions.

## Solution Interval Trade-off

- **Short solint** → Enables precise tracking of temporal variations but reduces SNR, potentially leading to unstable solutions
- **Long solint** → Improves SNR but may fail to capture temporal variations accurately

# Goal!

To understand how self-calibration performance  
**depends on solution time interval**

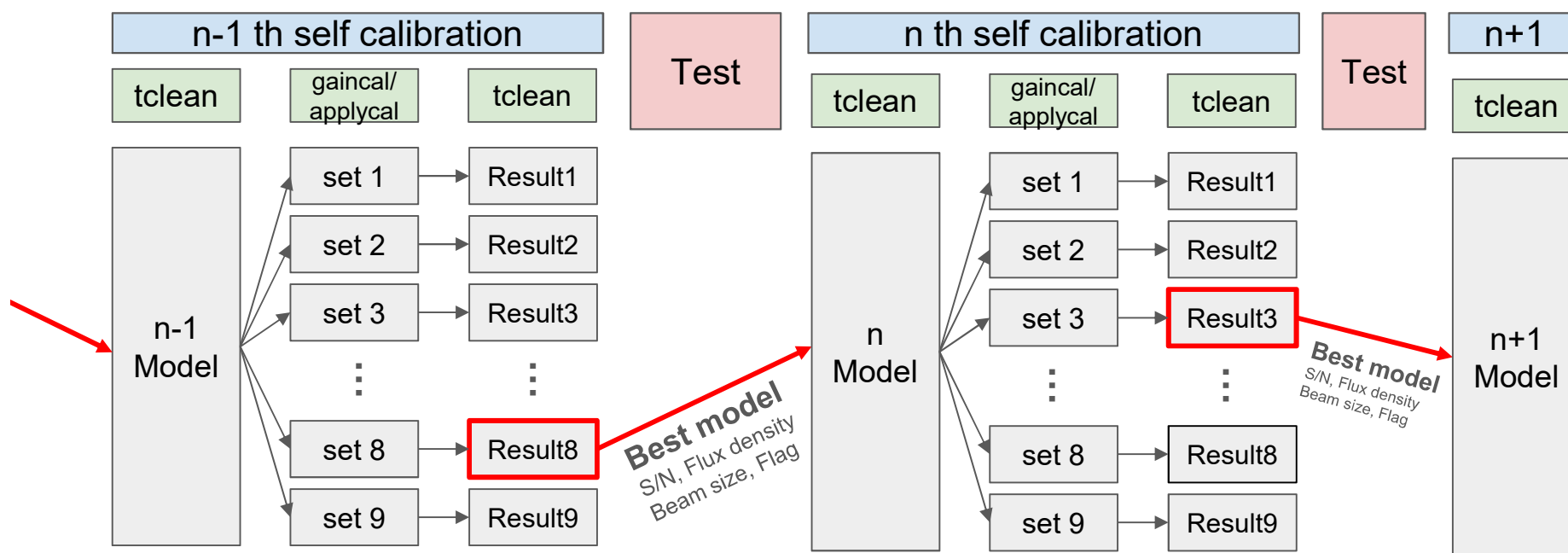
# Strategy

obs = 5 scan

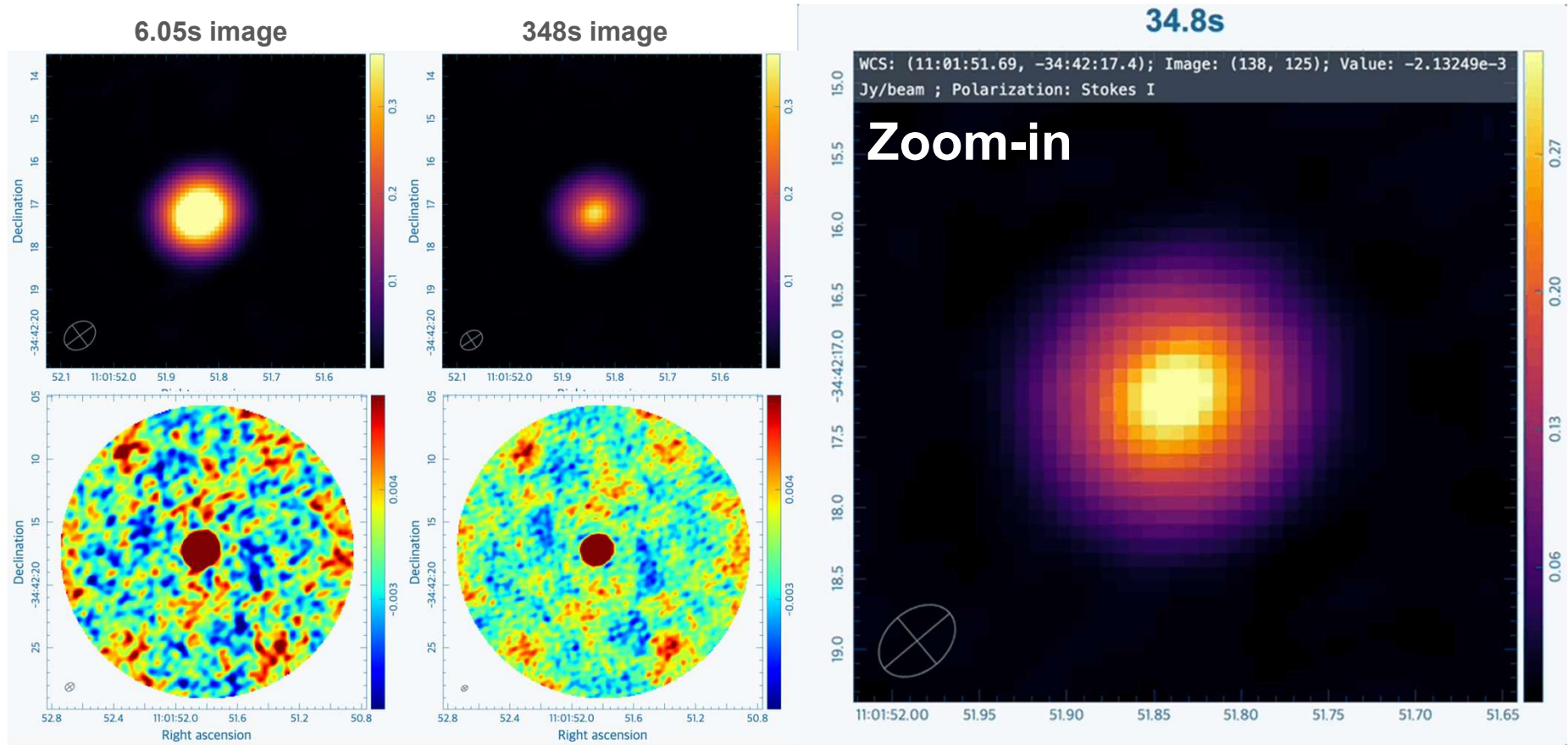
Combine	Solint	Combine	Solint	Combine	Solint
-	int (6.05s)	-	34.8s (0.1 scan)	-	87s ( $\frac{1}{4}$ scan)
-	116s ( $\frac{1}{3}$ scan)	-	174s ( $\frac{1}{2}$ scan)	-	inf (1 scan)
scan	509s ( $\frac{1}{3}$ obs)	scan	763.5s ( $\frac{1}{2}$ obs)	scan	inf (1 obs)

3 Phase self-cal + 1 Amplitude self-cal =  $4 * 9 =$  **36 test sets**

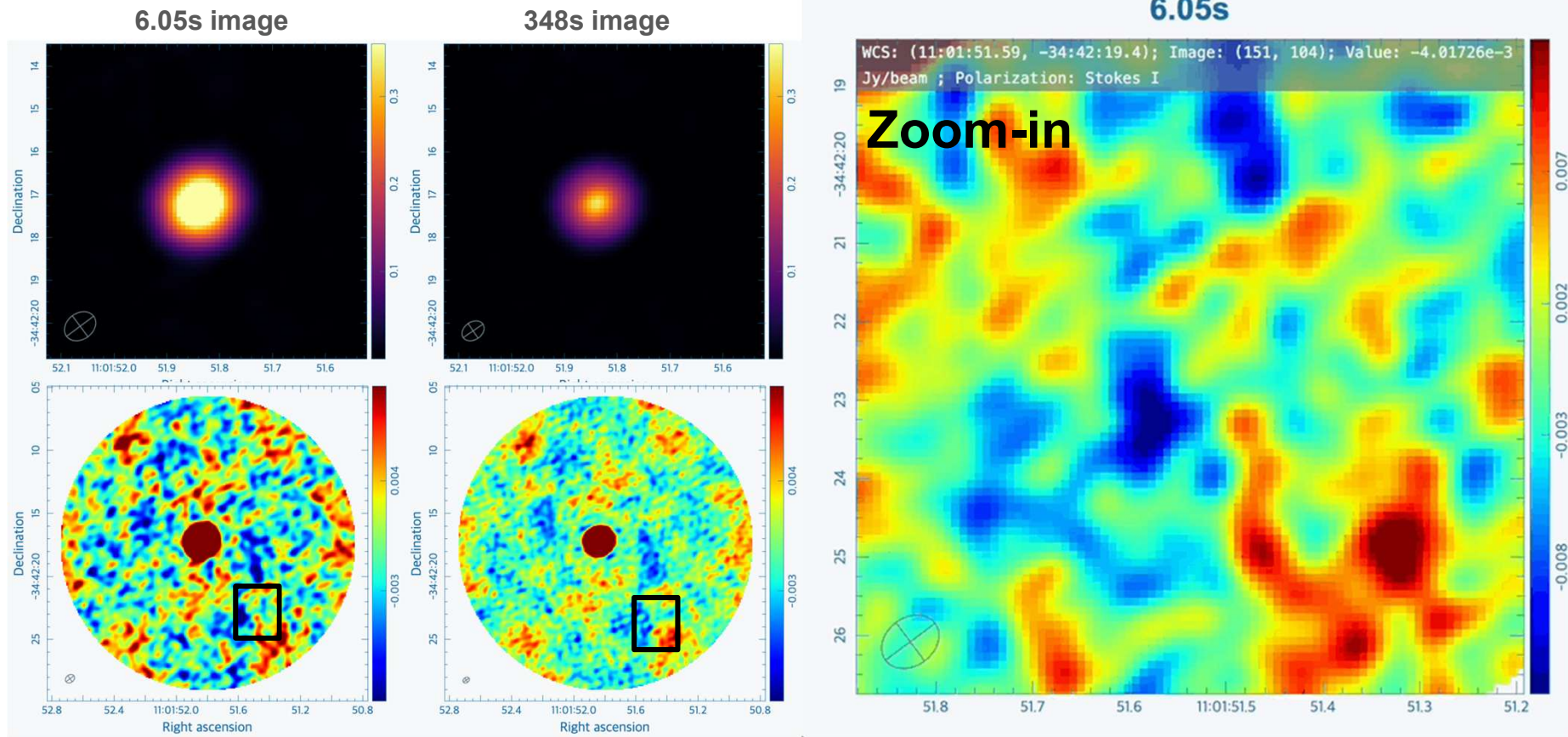
# Method



## 3rd Phase calibration image

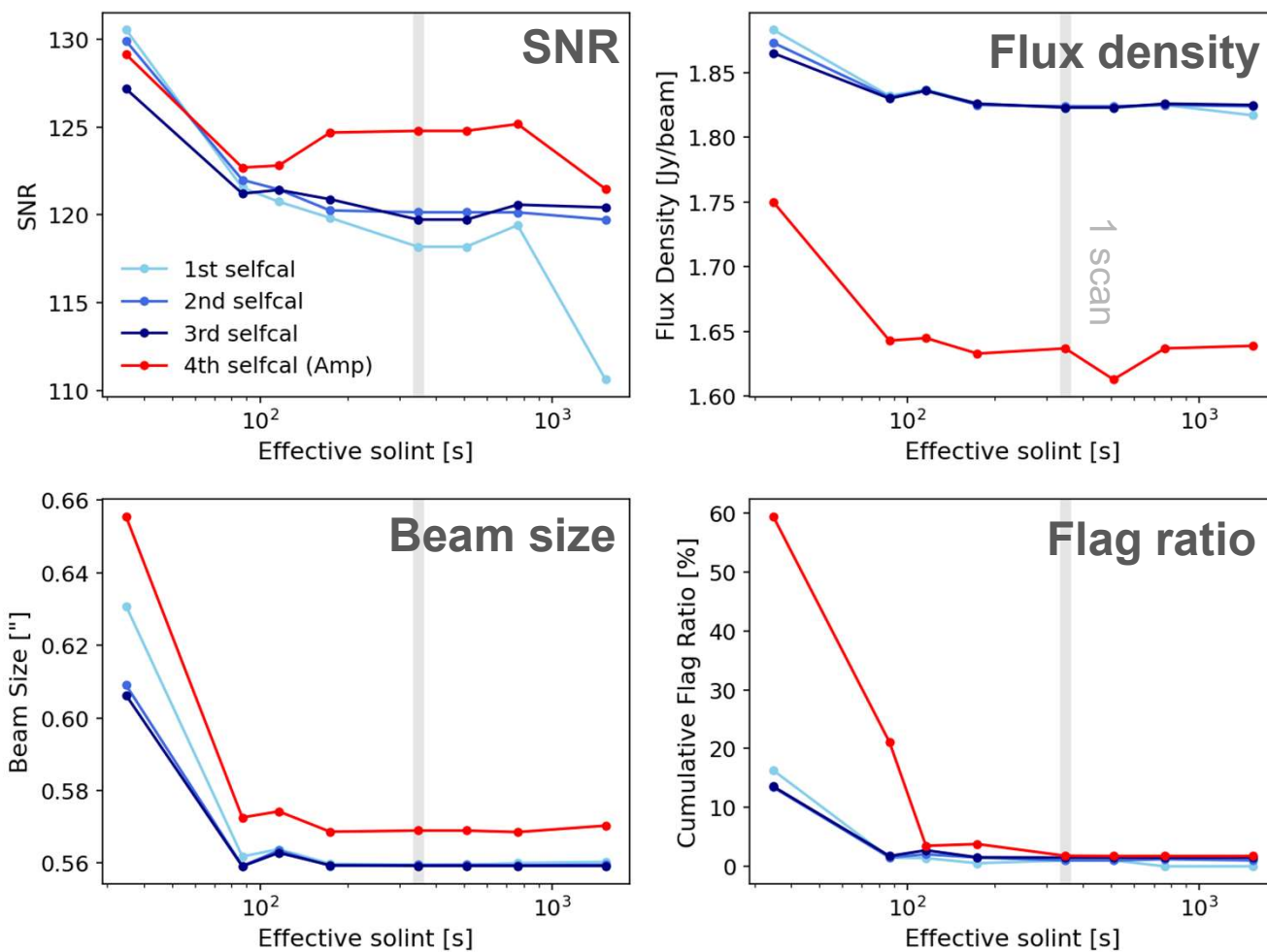


## 3rd Phase calibration image



# Result

$3\sigma$  clean image results



# Result $P = \text{Phase self-cal} / A = \text{Amplitude self-cal}$

**Flagging limit (<15%)** → Failed solution

**P:** 87s ( $\frac{1}{4}$  scan); **A:** 116s ( $\frac{1}{3}$  scan)

## Optimal solint

**P:** 1st →  $\frac{1}{2}$  obs; 2nd, 3rd →  $\frac{1}{2}$  scan

Iteration increase → More refined solutions at finer solint.

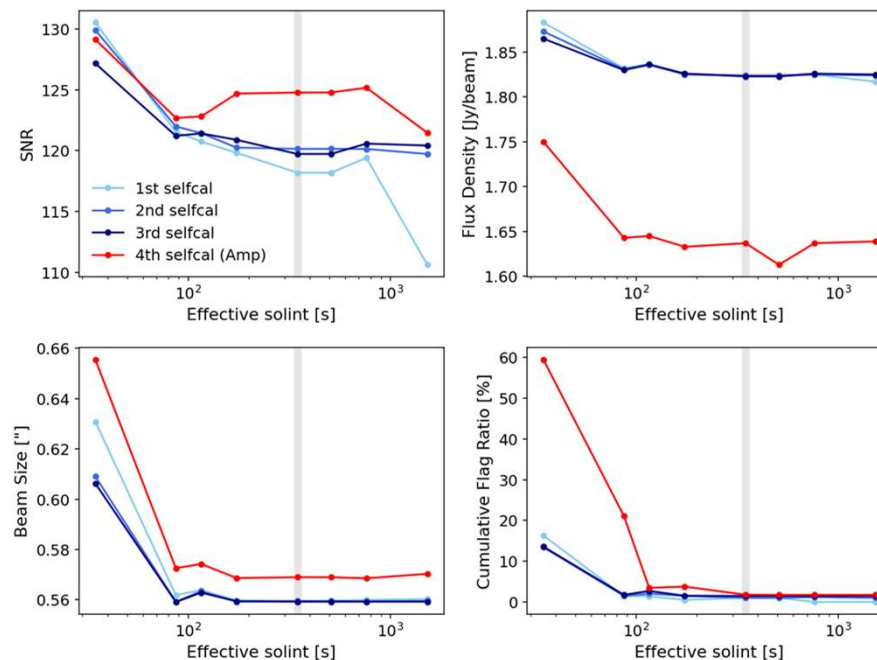
**A:**  $\frac{1}{2}$  obs, More accurate solutions at longer solint

∴ Require higher SNR data

## Solint='inf' + combine='scan'?

SNR degradation in self-cal solutions  
significantly varying between scans?

3 $\sigma$  clean image results





**Thank You**

## Case 1: weighting “briggs” (robust = 0.5)

clean image	1st image	2nd image	3rd image	4th image	final image
peak intensity (mJy/beam)	314	335	367	387	396
rms (mJy/beam)	5.64	2.83	2.51	2.46	2.49
SNR	56	118	146	157	159
flux (Jy)	1.57	1.73	1.88	1.90	1.88
beam size (arcsec)	0.560 * 0.416	0.559 * 0.415	0.580 * 0.439	0.600 * 0.445	0.614 * 0.457

## Case 2: weighting “uniform”

<b>clean image</b>	<b>1st image</b>	<b>2nd image</b>	<b>3rd image</b>	<b>4th image</b>	<b>final image</b>
peak intensity (mJy/beam)	298	319	320	321	310
rms (mJy/beam)	3.88	2.36	2.33	2.31	2.26
SNR	77	135	137	139	137
flux (Jy)	1.86	2.00	2.01	2.03	1.83
beam size (arcsec)	0.537*0.390	0.537*0.390	0.537*0.390	0.537*0.390	0.537*0.390

## Case 3: weighting “natural”

<b>clean image</b>	<b>1st image</b>	<b>2nd image</b>	<b>3rd image</b>	<b>4th image</b>	<b>final image</b>
peak intensity (mJy/beam)	377	397	456	497	528
rms (mJy/beam)	10.3	4.72	3.56	3.33	2.97
SNR	36.6	84	128	149	278
flux (Jy)	1.69	1.8	1.95	1.96	1.97
beam size (arcsec)	0.609 * 0.470	0.608 * 0.470	0.654 * 0.519	0.705 * 0.539	0.730 * 0.565