

A night sky with a full moon and a tree silhouette. The sky is dark blue with many stars and the Milky Way galaxy visible. A large, full moon is in the center, partially obscured by the silhouette of a tree. The foreground shows the dark silhouette of a hillside with trees.

Observing Tool

2025. 3. 14. Seokho Lee (KASI)

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Joint Proposals/ VLBI and Phased-Array proposal

Quick Guidance



Install Observing Tool

Announcement & release observing tool: Middle (~20) of March on <https://almascience.nao.ac.jp>

Please use the current cycle Observing Tool.

But, we can test it by using the previous cycle's one. For example, we can prepare the spectral setup, which is constrained by the characteristics of ALMA.





Server Maintenance
ALMA Science Portal (EA) will be unavailable for a short period.
Period: **13th Mar. 9:00-13:00** JST
In the meantime, please use the Science Portal (NA) at almascience.nrao.edu.

- Cycle 11 Announcement Poster
- ALMA Cycle 11 Call for Proposals
- ALMA Proposal Review
- Proposing Guidance
- Cycle 11 Proposer's Guide**
- Cycle 11 Capabilities
- Observing Tool
- Sensitivity Calculator
- ALMA Primer
- Technical Handbook
- Proposal Template
- Duplicate Observations
- DDT proposals

Proposing

The ALMA Director solicits proposals for use of Director's facilities and associated software. Proposals for use of Director's facilities are available on a yearly basis. The solicitation is done through a Call for Proposals, announced in this Section of the ALMA Science Portal. Supporting documents and associated software are available here now pertains to the current Cycle 11 Call for Proposals.

Server Maintenance

ALMA Science Portal (EA) will be temporarily unavailable due to server maintenance work.

Period: 13th Mar. 9:00-13:00 JST

In the meantime, please use the Science Portal of our partner institutions, which are located at http://almascience.eso.org and http://almascience.nrao.edu. We regret the inconvenience it may cause.

Cycle 11

Previous Cycles

Cycle 11 Documents

Call for Proposals

Documentation supporting the current ALMA Call for Proposals – **Cycle 11**. Documents from previous Cycles are provided [here](#).

principles-review-process

Document	Description
ALMA Proposer's Guide	Contains all pertinent information regarding the ALMA Call for Proposals
ALMA Technical Handbook	A comprehensive description of the ALMA observatory and its components
ALMA Users' Policies	The long-term core policies for use of the ALMA and ALMA data by the science community
Observing With ALMA - A Primer	Introduction to interferometry and how to use ALMA
ALMA Proposal Template	Zip files containing the proposal templates in LaTeX format. Recommended but not mandatory
ALMA Proposal Review Process	A detailed description of the ALMA Proposal Review Process
Principles of the ALMA Proposal Review Process	The latest version of the Principles of the ALMA Proposal Review Process

Observing With ALMA - A Primer	Introduction to interferometry and how to use ALMA
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Phase 1 & 2

ALMA Phase 1 (observing proposal) and Phase 2 (telescope runfiles for accepted proposals) materials are submitted through the [ALMA Observing Tool \(OT\)](#). Below are documentation which will aid the created and submitted of Phase 1 and Phase 2 with the OT.

Document	Description
OT Quickstart	A Quick Start Guide for using the Observing Tool
OT User Manual	Describes how to use the Observing Tool for preparing ALMA proposals
OT Reference Manual	An in-depth description of the Observing Tool
Video Tutorials	Video how-to for the Observing Tool
Known OT issues	For those instances when OT problems are encountered
Phase 2 Quickstart Guide	A Quick Start Guide for approved ALMA observing proposals - the process of Phase 2
A User's Guide to ALMA Scheduling Blocks	(Cycle 4) Guide to understanding the structure and content of ALMA Scheduling Blocks (SBs) using the Observing Tool (OT)

Guides to the ALMA Regional Centers

The ALMA Regional Centers provide user support and host special activities related to their respective regions. Their functions are described in the 'Guide to'.

Installer Page



- [Mac OS Installer](#)
- [Linux Installer](#)
- [Windows Installer](#)

YOU (and your Co-Is) should be registered in ALMA site.

Click on one of the links next to the OT Logo to download the Cycle 11 OT Installer for your particular operating system. The Installer is an executable file which can be started by double-clicking in a file-manager window or started from a shell's command line. Once started, it will take you through a number of screens which, for example, allow you to change the default amount of memory available to the OT. In most cases you can just accept all the defaults using the 'Next' button and click 'Install' when you are happy.

After the Installer has finished, an executable file ('ALMA-OT.sh' on Linux and 'ALMA-OT.command' on Macs) should be found inside a directory named 'ALMAOT-C11-2024'. This can be run from the command line or by double-clicking in a file manager if this is configured in this way. We recommend that the name of this directory not be changed so that multiple versions of the OT (for use in different cycles) can be maintained on your computer. On Macs, a shortcut will be created on your Desktop with the name 'ALMAOT-C11-2024' - the OS will probably ask to control your Finder for this to happen. In the case of macOS, if the ALMA OT is started via clicking on the desktop icon, a separate terminal window opens which should not be shut down whilst the OT is running.

Additional Information

- The Mac download is a zip archive which must first be opened in order to extract the installer. This will often be done automatically for you or a suitable program will be suggested ('Archive Utility').
- On Linux, typing 'sh almaot-C11-2024.bin' is the recommended way of starting the installer - it should not be necessary to make it executable. However, if this does not work, please run "chmod u+x almaot-C11-2024.bin" and then "./almaot-C11-2024.bin".
- There may be various issues related to security when running the Installer. Mac users may need to give permission to run the tool by opening the 'Security & Privacy' menu of 'System Preferences' and this menu should also be set to allow the use of apps from 'identified developers'. Alternatively, running the installer by right-clicking and choosing 'Open' (maybe twice) might work. On Windows, we are aware of 'Defender SmartScreen' - this can be bypassed by clicking on 'More Info'.
- It also appears that the installer will not work on older versions of macOS. So far, we only know that this is the case for 10.10 Yosemite. Users of this OS will have to use the tarball version.
- In contrast to the previous 'automated' OT installation (Web Start), the OT will no longer update itself automatically if an update is released. However, the OT will inform you if an update is available after which a new version of the OT Installer should be downloaded and the install procedure repeated. Re-running the Installer will overwrite the previous installation.

WARNING: Oracle have reported a serious incompatibility between macOS Sonoma 14.4 and Java which may result in the OT terminating unexpectedly – there is no workaround. Users are advised to avoid using Sonoma 14.4 and the OT if possible. If this is not possible and a user encounters this problem, the OT does have a project auto-backup facility which can be used as a recovery mechanism. Please contact the helpdesk should more information be required.



Atacama Large Millimeter/submillimeter Array

In search of our Cosmic Origins



Account info

Demographics

Expertise

Conflicts of interest

Confirm

New Account Registration

→ Next

First name



Middle initials



Surname



Gender



E-mail



Re-type E-mail



Receive optional emails



Account name





Expertise

Please select the category/keyword pair/s that best match your scientific expertise. You may select keywords in more than one category.

If you are a reviewer for Distributed Peer Review (DPR) you will preferentially be assigned proposals that match your selected keywords.

- > Cosmology and the High Redshift Universe
- > Galaxies and Galactic Nuclei
- ▼ ISM, star formation and astrochemistry
 - Outflows, jets and ionized winds
 - High-mass star formation
 - Intermediate-mass star formation
 - Low-mass star formation
 - Pre-stellar cores, Infra-Red Dark Clouds (IRDC)
 - Astrochemistry
 - Inter-Stellar Medium (ISM)/Molecular clouds
 - Photon-Dominated Regions (PDR)/X-Ray Dominated Regions (XDR)
 - HII regions
 - Magellanic Clouds
- > Circumstellar disks, exoplanets and the solar system
- > Stellar Evolution and the Sun

↻ Reset



Conflicts of interest

← Previous

Next →

If you are a reviewer for Distributed Peer Review or the Panel Review, please provide a list of your conflicts of interest. Consult the [conflicts of interest criteria](#) for guidance on what is considered a conflict. You will not be assigned to review a proposal in which the PI, a coPI, or a col is in your list of conflicts of interest.

Reviewers only need to identify conflicts of interest that are registered ALMA users since all reviewers must be registered. If an investigator is not in the ALMA user registry below, they do not need to be listed.

Providing this information is optional. If you do not provide a list of conflicts and do not check the box below, the JAO will identify potential conflicts based on your past ALMA collaborations.

+ Add investigator

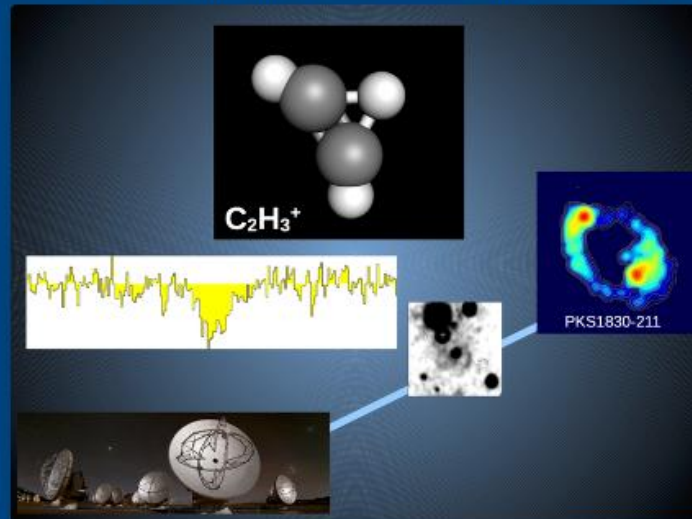
- Remove investigators



- Knowledgebase/FAQ
- Helpdesk
- EA ARC
- EU ARC
- NA ARC

Science Highlight

Protonated acetylene in the z=0.89 absorber toward PKS1830-211



View of the molecule (created with MolView) and an absorption spectrum observed with ALMA in the line of sight of the quasar PKS1830-211. The quasar (here observed with the MERLIN interferometer at radio wavelengths) is lensed by a foreground spiral galaxy at z=0.89 (optical image from HST).

The line of sight to the lensed blazar PKS1830-211 intercepts the disk of a foreground spiral galaxy at z=0.89 where absorption has been detected for more than 60 molecular species, mostly at mm wavelengths. In a paper accepted for publication in A&A. Dr. Sebastian Muller and colleagues report the detection of a new

Observatory News

ALMA Cycle 11 Call for Proposals is Now OPEN!
Mar 21, 2024

Planned Release of Data from Observatory Projects in Configurations 8 and 9
Mar 21, 2024

Over one third of all ALMA publications now make use of data from the ALMA Science Archive
Feb 16, 2024

Announcement for early proposal planning for Cycle 11
More...

NAOJ News

ALMA Cycle 11 - Workshop & Events
Jan 29, 2024

Dense GAS in the ALMA Science Archive
Jan 15, 2024

ALMAworkshop2023a on Ishigaki island - Workshop & Events
Jan 15, 2024

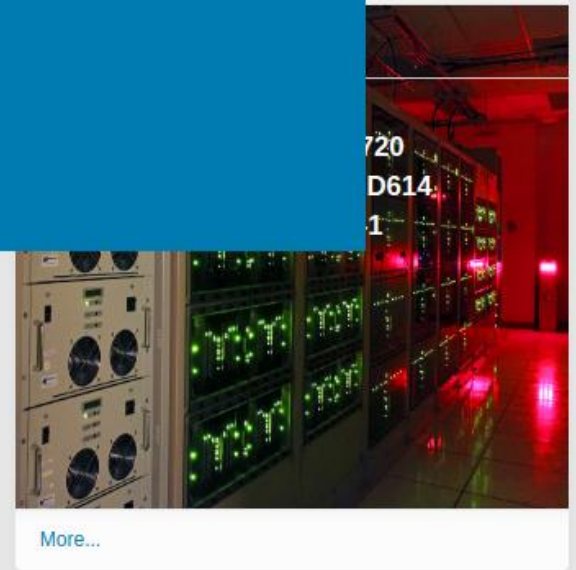
ALMA/45m/ASTE Users Meeting 2023 - Workshop & Events
Dec 13, 2023

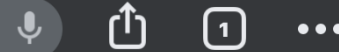
More...

The ALMA Science Portal is a one-stop source for information and tools aimed at the scientific community as a whole, including proposers, archive researchers, ALMA staff, journalists, and funding agencies.

Quick Links

ALMA Basics	Configuration Schedule
ALMA Science	SnooPI
ALMA Primer	DDT Proposals





Search bar with magnifying glass icon and placeholder text: "How can we help you today?"

Knowledgebase

Folder icon: General (1)

Folder icon: ALMA Observing Tool (OT) (44)

- What Cycle 9 proposal issues and clarifications should I ...
- What do the time estimates in the summary PDF page g...
- My project requires good uv-coverage. How should I set ...



Important Input Parameters

Important parameters I

- Expected Source properties (**Field Setup**)
 - Position, source velocity
 - **Peak Flux Density per beam**
 - SNR > 3-5
 - **Polarization**
 - linear > 0.1% (< 0.3 FOV)
 - circular > 1.8 % (< 0.1 FOV)
 - **Line widths**
 - > 3 x spectral resolution

You should describe how to derive/adopt these values in **Technical Justification**

ALMA Observing Tool (Cycle 11 (Phase 2 Patch 1)) - Project

Project Structure: Unsubmitted Proposal

Editors: Spectral, Spatial, Field Setup

Spatial Image: TW Hya

Source Name: TW Hya

Source Coordinates: RA 11:01:51.9053, Dec -34:42:17.033

Source Radial Velocity: 12.335 km/s

Expected Source Properties:

Peak Continuum Flux Density per Synthesized Beam	30.00000	mJy
Continuum Linear Polarization	0.0	per cent
Continuum Circular Polarization	0.0	per cent
Peak Line Flux Density per Synthesized Beam	45.00000	mJy
Line Width	0.30000	km/s
Line Linear Polarization	0.0	per cent
Line Circular Polarization	0.0	per cent

FOV Parameters: Representative Frequency (Sky) 330.574 GHz, Array Type 12m, Antenna Beamsize (HPBW) 17.615 arcsec

Image Query: Image Server Digitized Sky (Version II) at ESO, Image Size(arcmin) 10.0

Important parameters II

- Scales (Control and Performance; Field Setup)
 - Angular Resolution (beam size)
 - depends on the longest baseline and frequency
 - Maximum Recoverable Scale (MRS)
 - depends on the shortest baseline ($\sim 10 \times$ beam size)
 - When the scale is longer than MRS, the emission is resolve out
 - **Largest Angular Structure (LAS)** $< \text{MRS} \rightarrow$ single configuration
 - **LAS** $> \text{MRS} \rightarrow$ multiple configuration or ACA and TP are added.
 - Field of View (FOV)
 - FWHM of the 12m telescope primary beam
 - ~ 19 arcsec (33 arcsec) @ 300 GHz for 12m (7m)
 - Area of target is larger than $1/3$ FOV, **mosaic** is needed.

Schedule for C12 configurations

Cycle 12			
Start Date	Configuration	Longest baseline	LST: Best of
1-Oct-2025	C-8	8.5 km	22-10
20-Oct-2025	C-7	3.6 km	23-11
10-Nov-2025	C-6	2.5 km	1-13
1-Dec-2025	C-5	1.4 km	2-14
20-Dec-2025	C-4	0.78 km	4-15
10-Jan-2026	C-3	0.50 km	5-17
1-Feb-2026	No observations due to maintenance		
1-Mar-2026	C-1	0.16 km	8-21
26-Mar-2026	C-2	0.31 km	9-23
20-Apr-2026	C-3	0.50 km	11-0
10-May-2026	C-4	0.78 km	12-2
31-May-2026	C-5	1.4 km	13-4
23-Jun-2026	C-6	2.5 km	15-6
28-Jul-2026	C-5	1.4 km	17-7
18-Aug-2026	C-4	0.78 km	19-8
10-Sep-2026	C-3	0.50 km	20-9

AR and MRS for C12 configurations

Config.	Band	L _{max}	L _{min}	Freq. (GHz)	1	3	4	5	6	7	8	9	10
					40	100	150	185	230	345	460	650	870
7-m	45 m	θ_{res} (arcsec)			31.5	12.5	8.35	6.77	5.45	3.63	2.72	1.93	1.44
		θ_{MRS} (arcsec)			167	66.7	44.5	36.1	29.0	19.3	14.5	10.3	7.67
C-1	161 m	θ_{res} (arcsec)			8.45	3.38	2.25	1.83	1.47	0.98	0.74	0.52	0.39
		θ_{MRS} (arcsec)			71.2	28.5	19.0	15.4	12.4	8.25	6.19	4.38	3.27
C-2	314 m	θ_{res} (arcsec)			5.75	2.30	1.53	1.24	1.00	0.67	0.50	0.35	0.26
		θ_{MRS} (arcsec)			56.5	22.6	15.0	12.2	9.81	6.54	4.90	3.47	2.59
C-3	500 m	θ_{res} (arcsec)			3.55	1.42	0.94	0.77	0.62	0.41	0.31	0.22	0.16
		θ_{MRS} (arcsec)			40.5	16.2	10.8	8.73	7.02	4.68	3.51	2.48	1.86
C-4	784 m	θ_{res} (arcsec)			2.30	0.92	0.61	0.50	0.40	0.27	0.20	0.14	0.11
		θ_{MRS} (arcsec)			28.0	11.2	7.50	6.08	4.89	3.26	2.44	1.73	1.29
C-5	1.4 km	θ_{res} (arcsec)			1.38	0.55	0.36	0.30	0.24	0.16	0.12	0.084	0.063
		θ_{MRS} (arcsec)			16.8	6.70	4.47	3.62	2.91	1.94	1.46	1.03	0.77
C-6	2.5 km	θ_{res} (arcsec)			0.78	0.31	0.20	0.17	0.13	0.089	0.067	0.047	0.035
		θ_{MRS} (arcsec)			10.3	4.11	2.74	2.22	1.78	1.19	0.89	0.63	0.47
C-7	3.6 km	θ_{res} (arcsec)			0.53	0.21	0.14	0.11	0.092	0.061	0.046	0.033	0.024
		θ_{MRS} (arcsec)			6.45	2.58	1.72	1.40	1.12	0.75	0.56	0.40	0.30
C-8	8.5 km	θ_{res} (arcsec)			0.24	0.096	0.064	0.052	0.042	0.028	0.021	0.015	0.011
		θ_{MRS} (arcsec)			3.55	1.42	0.95	0.77	0.62	0.41	0.31	0.22	0.16

Band 5 (around 183GHz) and 7-10 are recommended within LST ranges (not Dec-March)

Important parameters II

- Scales (Control and Performance; Field Setup)
 - Angular Resolution (beam size)
 - depends on the longest baseline and frequency
 - Maximum Recoverable Scale (MRS)
 - depends on the shortest baseline (~ 10 x beam size)
 - When the scale is longer than MRS, the emission is resolve out
 - **Largest Angular Structure (LAS)** $< \text{MRS} \rightarrow$ single configuration
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 - Field of View (FOV)
 - FWHM of the 12m telescope primary beam
 - ~ 19 arcsec (33 arcsec) @ 300 GHz for 12m (7m)
 - Area of target is larger than $1/3$ FOV, **mosaic** is needed.

Most Extended configuration	Allowed Compact configuration pairings	Extended 12-m Array Multiplier	Multiplier if compact 12-m Array needed	Multiplier if 7-m Array needed	Multiplier if TP Array needed and allowed (with 7-m Array in 4x4-bit mode)	Multiplier if TP Array needed and allowed (with 7-m Array in 2x2-bit mode)
7-m Array	TP			1	1.7	1.4
C-1	7-m Array & TP	1		7.0	11.9	9.5
C-2	7-m Array & TP	1		4.7	7.9	6.3
C-3	7-m Array & TP	1		2.4	4.1	3.3
C-4	C-1 & 7-m Array & TP	1	0.34	2.4	4.0	3.2
C-5	C-2 & 7-m Array & TP	1	0.26	1.2	2.1	1.7
C-6	C-3 & 7-m Array & TP	1	0.25	0.6	1.0	0.8
C-7	C-4	1	0.23			
C-8	C-5	1	0.22			
C-9	C-6	1	0.21			
C-10	–	1				

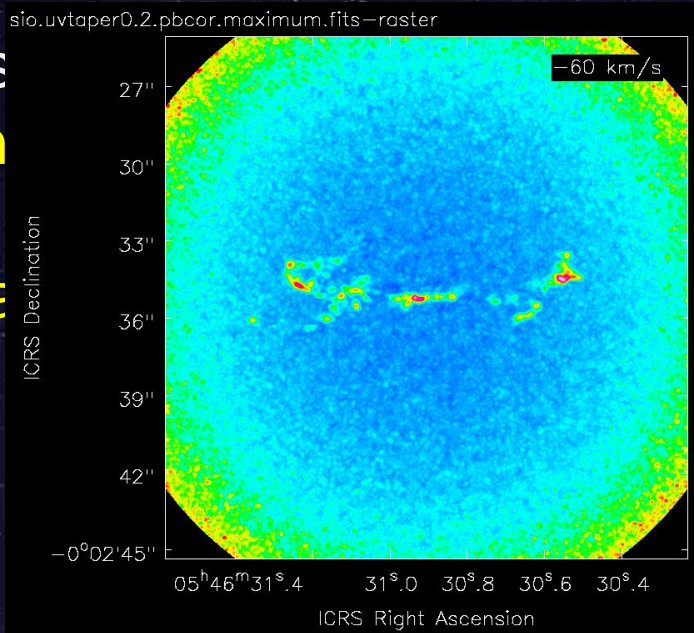
Table A-2: Allowed Array Combinations and Time Multipliers. See Chapter 7 of the [Technical Handbook](#) for relevant equations and detailed considerations. If the array configuration that meets the AR request according to Table A-1 has a MRS that is smaller than the LAS request, the OT checks if adding more compact array configurations, following the restrictions of this Table, fulfills the LAS request. If so, the final setup consists of the selected combination of arrays. Otherwise, the OT returns a validation error.

Important

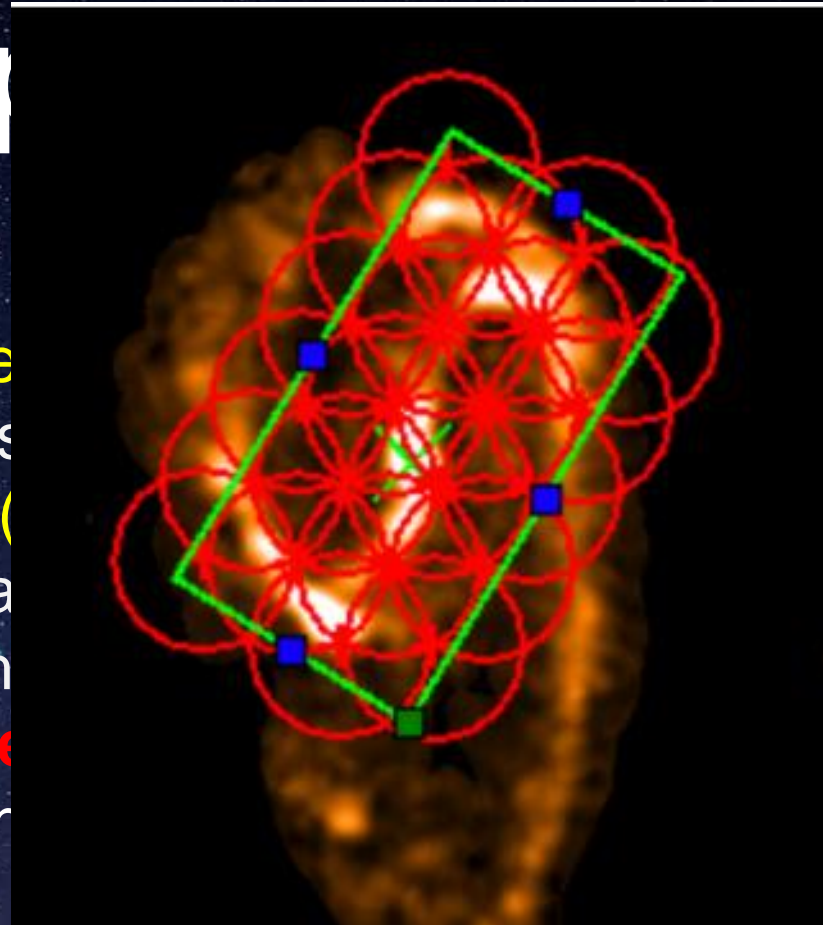
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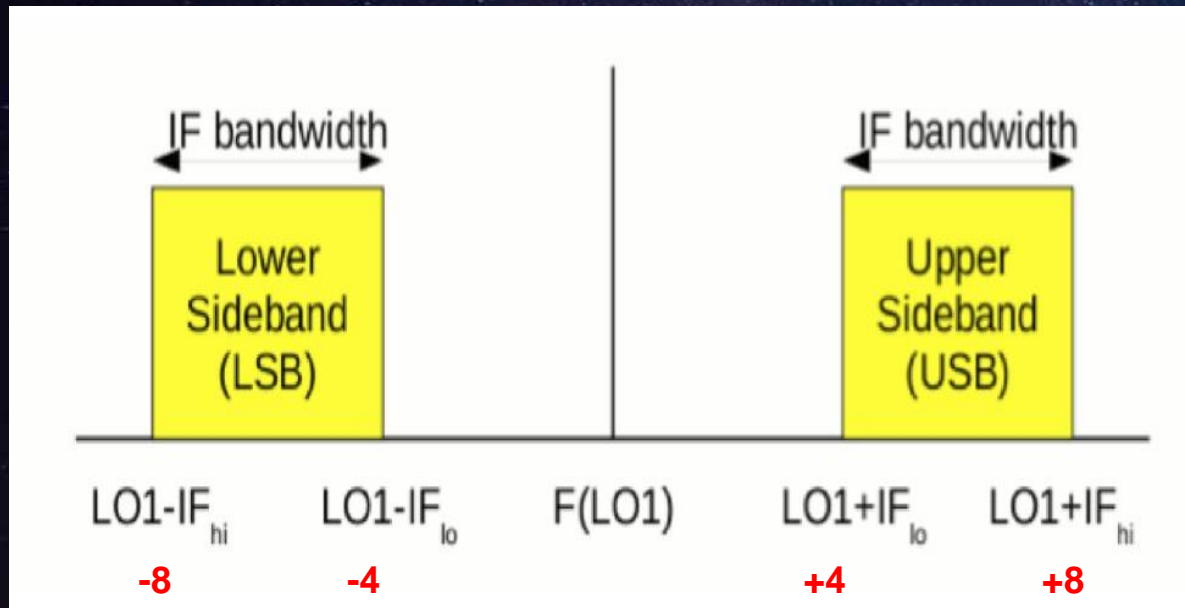
- Field of View (FOV)

- FWHM of the 12m telescope primary beam
 - ~19 arcsec (33 arcsec) @ 300 GHz for 12m (7m)
 - Area of target is larger than 1/3 FOV, **mosaic** is needed.

Important parameters III

- Spectral Setup
 - LSB and/or USB
 - 4 basebands (with 2GHz max. width)
 - 2 or 4 basebands in the one sideband

LSB/USB (<4GHz, ≤4 basebands)
 Baseband (<2GHz, ≤4spws)
 Spectral window(spw)



Band	Frequency range (GHz)	Wavelength range (mm)	IF range (GHz)	Type
1	35 – 50	8.5 – 6	4 – 12	SSB
3	84 – 116	3.6 – 2.6	4 – 8	2SB
4	125 – 163	2.4 – 1.8	4 – 8	2SB
5	158 – 211	1.9 – 1.4	4 – 8	2SB
6	211 – 275	1.4 – 1.1	4.5 – 10	2SB
7	275 – 373	1.1 – 0.8	4 – 8	2SB
8	385 – 500	0.78 – 0.60	4 – 8	2SB
9	602 – 720	0.50 – 0.42	4 – 12	DSB
10	787 – 950	0.38 – 0.32	4 – 12	DSB

Spws in a baseband

- one fraction 1
- two fraction $\frac{1}{2}$
- four fraction $\frac{1}{4}$
- one fraction $\frac{1}{2}$ + two fraction $\frac{1}{4}$

Spectral windows (SPW) should have the same resolution.

Bandwidth (MHz)	Channel spacing (MHz)	Spectral resolution (MHz)	Number of channels	Correlator mode	Bit Mode
1875	15.6	31.2	120	TDM	
938	0.976	1.952	1024	FDM	4x4 *
1875	0.488	0.976	3840	FDM	2x2
469	0.488	0.976	1024	FDM	4x4
938	0.244	0.488	3840	FDM	2x2
234	0.244	0.488	1024	FDM	4x4
469	0.122	0.244	3840	FDM	2x2
117	0.122	0.244	1024	FDM	4x4
234	0.061	0.122	3840	FDM	2x2
58.6	0.061	0.122	1024	FDM	4x4
117	0.0305	0.061	3840	FDM	2x2
58.6	0.0153	0.0305	3840	FDM	2x2

Table 5.1: Available spectral windows in multi-region mode (dual polarization). Each time the fraction is changed, the number of channels and bandwidth of a particular correlator mode is halved. Each row corresponds to a particular spectral resolution.

Spectral resolution \propto 1/ fraction for a given bandwidth

Fraction = 1		Fraction = 1/2		Fraction = 1/4	
Bandwidth (MHz)	# channels	Bandwidth (MHz)	# channels	Bandwidth (MHz)	# channels
1875	4096	937.5	2048	468.75	1024
937.5	4096	468.75	2048	234.375	1024
468.75	4096	234.375	2048	117.118	1024
234.375	4096	117.118	2048	58.594	1024
117.118	4096	58.594	2048	not available	
58.594	4096	not available		not available	



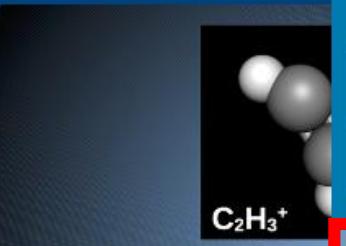
Procedures In OT



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- Help

Science Highlight

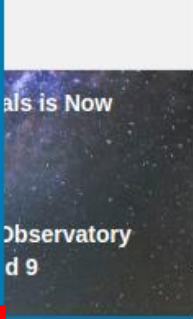
Protonated acetylene in the z=0.89 galaxy



View of the molecule (created from the absorption spectrum observed from the line of sight of the quasar PKS 1830-211 (here observed with the Mopra telescope at radio wavelengths) is lensed by a foreground galaxy at z=0.89 (optical image from HST).

The line of sight to the lensed blazar PKS1830-211 intercepts the disk of a foreground spiral galaxy at z=0.89 where absorption has been detected for more than 60 molecular species, mostly at mm wavelengths. In a paper accepted for publication in A&A, Dr. Sebastian Muller and colleagues report the detection of a new

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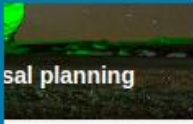
NAOJ News

ALMA Cycle 11 Proposal Preparation Meeting - Workshop & Events
Jan 29, 2024

Dense GAS in Nearby Galaxies - Workshop & Events
Jan 15, 2024

Troubleshooting

OT Video Tutorials



ALMA/45m/ASTE Users Meeting 2023 - Workshop & Events
Dec 13, 2023

More...

ALMA Status

Configuration Schedule

Refereed publications: 3720
Last observed source: XID614
Current configuration: C-1

More...

stop source for information and tools aimed at the scientific community as a whole, including proposers, archive researchers, ALMA facilities.

ALMA Basics	Configuration Schedule
ALMA Science	SnooPI
ALMA Primer	DDT Proposals



Observing Tool

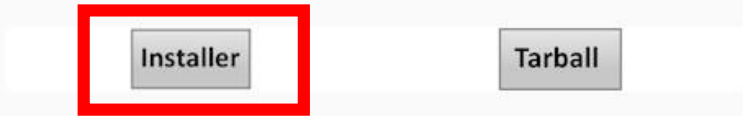
The ALMA Observing Tool (OT) is a Java desktop application used for the preparation and submission of ALMA Phase 1 proposals and, for those which are accepted, Phase 2 materials (Scheduling Blocks). It is also used for preparing and submitting Director's Discretionary Time (DDT) proposals and Supplemental Call (ACA stand-alone) proposals. The current *Cycle 11* release of the OT is configured for the present capabilities of ALMA as described in the Proposer's Guide. Note that in order to submit proposals you will have to register with the ALMA Science Portal beforehand.

Download & Installation

The OT should run on all common operating systems and depends on a version of Java being available. The Cycle 11 version of the OT will come with its own version of Java 17 and thus the users need no longer worry about their local Java installation. Unfortunately, as Java 17 does not include Web Start, this version of the OT is no longer available. The Cycle 11 OT can be installed in two different ways, either with a modern installer or manually with a tarball distribution.

It is recommended that the OT be installed using the ALMA **OT Installer**. This uses a modern graphical interface to report the progress of the installation and allows the user to change various settings from their defaults, including the amount of memory the OT may use. The installation will produce an executable file that can be used to start the OT. If problems are encountered with the installer, then the tarball must be used.

The **tarball** version must be installed manually and the instructions for doing this have not changed.



NOTE: For those who require the **Cycle 10 version of the OT**, it can be found [here](#).

Documentation

Extensive documentation is available to help you work with the OT and optimally prepare your proposal:

- If you are a novice OT user you should start with the [OT Quickstart Guide](#), which takes you through the basic steps of ALMA proposal preparation.



Installer Page



- [Mac OS Installer](#)
- [Linux Installer](#)
- [Windows Installer](#)

Click on one of the links next to the OT Logo to download the Cycle 11 OT Installer for your particular operating system. The Installer is an executable file which can be started by double-clicking in a file-manager window or started from a shell's command line. Once started, it will take you through a number of screens which, for example, allow you to change the default amount of memory available to the OT. In most cases you can just accept all the defaults using the 'Next' button and click 'Install' when you are happy.

After the Installer has finished, an executable file ('ALMA-OT.sh' on Linux and 'ALMA-OT.command' on Macs) should be found inside a directory named 'ALMAOT-C11-2024'. This can be run from the command line or by double-clicking in a file manager if this is configured in this way. We recommend that the name of this directory not be changed so that multiple versions of the OT (for use in different cycles) can be maintained on your computer. On Macs, a shortcut will be created on your Desktop with the name 'ALMAOT-C11-2024' - the OS will probably ask to control your Finder for this to happen. In the case of macOS, if the ALMA OT is started via clicking on the desktop icon, a separate terminal window opens which should not be shut down whilst the OT is running.

Additional Information

- The Mac download is a zip archive which must first be opened in order to extract the installer. This will often be done automatically for you or a suitable program will be suggested ('Archive Utility').
- On Linux, typing 'sh almaot-C11-2024.bin' is the recommended way of starting the installer - it should not be necessary to make it executable. However, if this does not work, please run "chmod u+x almaot-C11-2024.bin" and then "./almaot-C11-2024.bin".
- There may be various issues related to security when running the Installer. Mac users may need to give permission to run the tool by opening the 'Security & Privacy' menu of 'System Preferences' and this menu should also be set to allow the use of apps from 'identified developers'. Alternatively, running the installer by right-clicking and choosing 'Open' (maybe twice) might work. On Windows, we are aware of 'Defender SmartScreen' - this can be bypassed by clicking on 'More Info'.
- It also appears that the installer will not work on older versions of macOS. So far, we only know that this is the case for 10.10 Yosemite. Users of this OS will have to use the tarball version.
- In contrast to the previous 'automated' OT installation (Web Start), the OT will no longer update itself automatically if an update is released. However, the OT will inform you if an update is available after which a new version of the OT Installer should be downloaded and the install procedure repeated. Re-running the Installer will overwrite the previous installation.

WARNING: Oracle have reported a serious incompatibility between macOS Sonoma 14.4 and Java which may result in the OT terminating unexpectedly – there is no workaround. Users are advised to avoid using Sonoma 14.4 and the OT if possible. If this is not possible and a user encounters this problem, the OT does have a project auto-backup facility which can be used as a recovery mechanism. Please contact the helpdesk should more information be required.



Server Maintenance
ALMA Science Portal (EA) v...
Period: 13th Mar. 9:00-13:00
In the meantime, please use

- Cycle 11 Announcement Poster
- ALMA Cycle 11 Call for Proposals
- ALMA Proposal Review
- Proposing Guidance
- Cycle 11 Proposer's Guide
- Cycle 11 Capabilities
- Observing Tool**
- Sensitivity Calculator
- ALMA Primer
- Technical Handbook
- Proposal Template
- Duplicate Observations
- DDT proposals

Troubleshooting

On this page we provide help for users running an up-to-date version of the Observing Tool (OT). If you have any further questions or discover a new bug, please contact us via the ALMA Helpdesk. Please note that the following advice assumes that the user is running an up-to-date version of the Observing Tool (OT). If you are using an older version, please install the OT via the tarball.

I can't run the installer
There is an issue that the OT (Observing Tool) requires a copy of Java 17.

Do I need to install Java?
For the OT that will be used in Cycle 11, you need to have a copy of Java 17.

If so, which version of Java?
The Cycle 11 OT requires Java 17. You can find the Java 17 JRE (Java Runtime Environment) or the tarball (that does with its own version of Java. A tarball distribution without Java is available and if this is used, a version of Java 17 must be installed on your computer).

Should I install the tarball?

science.eso.org and <http://almascience.nrao.edu>. We regret the inconvenience it may cause.

On this page we provide help for users running an up-to-date version of the Observing Tool (OT). If you have any further questions or discover a new bug, please contact us via the ALMA Helpdesk. Please note that the following advice assumes that the user is running an up-to-date version of the Observing Tool (OT). If you are using an older version, please install the OT via the tarball.

Troubleshooting

OT Video Tutorials

For the OT that will be used in Cycle 11, you need to have a copy of Java 17.

The Cycle 11 OT requires Java 17. You can find the Java 17 JRE (Java Runtime Environment) or the tarball (that does with its own version of Java. A tarball distribution without Java is available and if this is used, a version of Java 17 must be installed on your computer).


```
shlee@nb-shlee: ~/ALMAOT-C11-2024_Prop
shlee@nb-shlee:~/ALMAOT-C11-2024_Prop$ ls
_ALMAOT-C11-2024_installation  ALMA-OT.sh  jre
shlee@nb-shlee:~/ALMAOT-C11-2024_Prop$ ./ALMA-OT.sh
Starting ALMA-OT using java from /home/shlee/ALMAOT-C11-2024/_ALMAOT-C11-2024_in
stallation/./jre/bin/java
```




```
shlee@nb-shlee: ~/ALMAOT-C11-2024_Prop
shlee@nb-shlee:~/ALMAOT-C11-2024_Prop$ ls
_ALMAOT-C11-2024_installation  ALMA-OT.sh  jre
shlee@nb-shlee:~/ALMAOT-C11-2024_Prop$ ./ALMA-OT.sh
Starting ALMA-OT using java from /home/shlee/ALMAOT-C11-2024/_ALMAOT-C11-2024_in
stallation/./jre/bin/java
```

Startup Options [X]

What would you like to do?

- Create a new proposal
- Create a new DDT proposal
- Open an existing project from disk
- Retrieve a project from the ALMA science archive

Do not show this message again

OK



Project Structure

New Science Goal

Editors

Proposal Program

Spectral Spatial Proposal

Unsubmitted Proposal

- Project
 - Proposal

Proposal Information

Proposal Title

Proposal Cycle

2024.1

Abstract
(max. 1200 characters)

Proposal Type

- Regular Target Of Opportunity VLBI
 Large Program Phased Array

Scientific Category

- Cosmology and the High Redshift Universe Galaxies and Galactic Nuclei ISM, star formation and astrochemistry
 Circumstellar disks, exoplanets and the solar system Stellar Evolution and the Sun

Feedback

Validation Validation History Log



Description

Suggestion

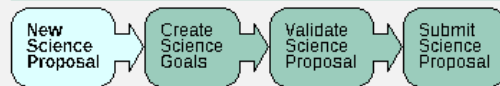
↑ ↓ ?

Overview

Contextual Help






- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
- Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help



- 1 New Proposal Ctrl+N
-  New DDT Proposal Ctrl+D
-  New Supplemental Call Proposal
- Open Project >
- Open Project as New Proposal >
-  Save Ctrl+S
- Save As
-  Show ALMA Template Library
- Use Project as Template >
- Validate Ctrl+L
-  Submit Project
- Preferences Ctrl+P
- Save Preferences
- Quit

Editors

< Spectral Spatial **Proposal** >

- Open Archive Proposal as a New Proposal
- Open Project from ALMA Archive, and edit as new Proposal
- Open Archive Proposal as a New ACA Supplemental Call Proposal

Science Case and Management Plan

Please ensure that your science case and management plan are properly anonymized following instructions on the Science Portal. The total length of this document is limited to 7 pages, where 6 pages should be allocated to the science case and references, and one page should be allocated to the management plan. Please refer to the Proposer's Guide for the details on the information that needs to be provided in the management plan.

Science Case + Management Plan (Mandatory, single PDF, 7 pages max),

Team Expertise

Team Expertise (PDF, 1 page max.)

Scheduling Feasibility

By checking this box, the PI confirms that the proposal is consistent with the configuration/LST scheduling constraints for Large Programs as described in the Proposer's Guide. Proposals that do not obey the scheduling constraints may be rejected on technical grounds.

Is the scheduling of this proposal feasible? Yes No

Duplicate observations

Briefly justify any new observations that duplicate archival data or accepted programs. Information regarding the ALMA Duplication Policy and how to search archival data and accepted programs can be found at:

<https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>

Observatory Use Only

Feedback



Project Structure

Editors

Proposal Program

Spectral Spatial Proposal

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup

Select PI

Add CoPI

Add Col

Remove Collaborator

Add fr

Find Projects In ALMA Archive

Search for Projects

 all my projects by PI ALMA ID by Project Code by Project Name

Search

Found Projects

Project Name	Project Code	PI ALMA ID	Creation Time	Submission Time	Project ID	Open
--------------	--------------	------------	---------------	-----------------	------------	------

Cancel

Observatory Use Only

Feedback

Validation

Validation History

Log



Project Structure

New Science Goal

Editors

Proposal Program

Spectral Spatial Proposal

Unsubmitted Proposal

- Project
 - Proposal

Proposal Information

Proposal Title

Proposal Cycle

2024.1

Abstract
(max. 1200 characters)

Proposal Type

- Regular Target Of Opportunity VLBI
 Large Program Phased Array

Scientific Category

- Cosmology and the High Redshift Universe Galaxies and Galactic Nuclei ISM, star formation and astrochemistry
 Circumstellar disks, exoplanets and the solar system Stellar Evolution and the Sun

Feedback

Validation Validation History Log



Description

Suggestion

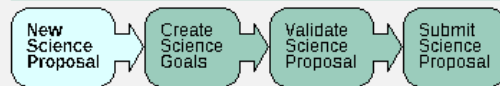
↑ ↓ ?

Overview

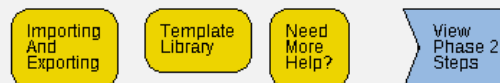
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#).
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
- Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help



File Edit View Tool Search Help

New Proposal Ctrl+N

New DDT Proposal Ctrl+D

New Supplemental Call Proposal

Open Project >

Open Project as New Proposal >

Save Ctrl+S

Save As

Show ALMA Template Library

Use Project as Template >

Validate Ctrl+L

Submit Project

Preferences Ctrl+P

Save Preferences

Quit

When you click a menu of 'File'

Editors

Spectral Spatial Proposal

Proposal Information

Proposal Title

Proposal Cycle

2024.1

Abstract
(max. 1200 characters)

Proposal Type

 Regular Target Of Opportunity VLBI Large Program Phased Array

Scientific Category

 Cosmology and the High
Redshift Universe Galaxies and Galactic
Nuclei ISM, star formation and
astrochemistry Circumstellar disks,
exoplanets and the solar
system Stellar Evolution and the
Sun

Feedback

Validation

Validation History



Log

Description

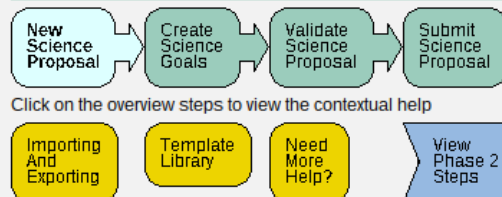
Suggestion

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
- Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal





Project Structure

Editors

Proposal Program

Spectral Spatial Project

Unsubmitted Proposal

- Project
 - Proposal

Principal Investigator

Select PI...

Main Project Information

Project

Assigned Priority

Project Code

None Assigned

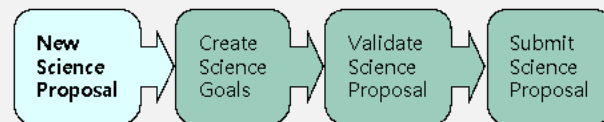
1. Open a Science Goal (SG)

Overview

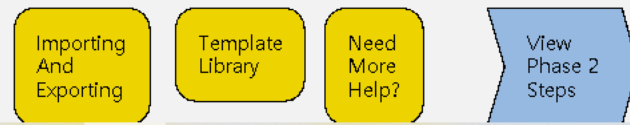
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the **1** icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help

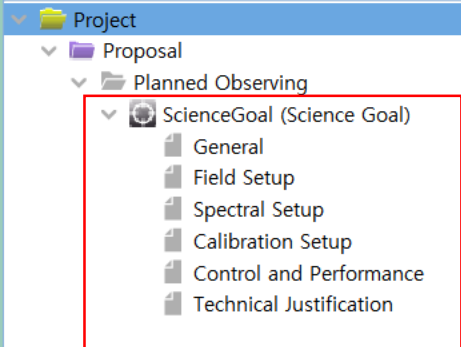




Project Structure

Proposal Program

Unsubmitted Proposal



Overview

Editors

Spectral Spatial Project

Principal Investigator ?



Main Project Information ?

Project

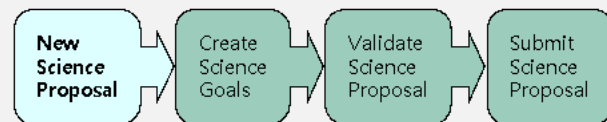
Assigned Priority

Project Code

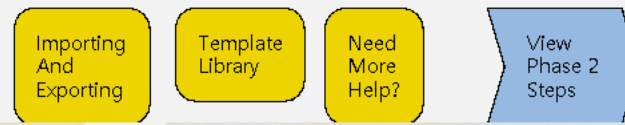
Contextual Help

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 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
- Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Proposal Information

Proposal Title

Proposal Cycle 2023.1

Abstract (max. 1200 characters)

Proposal Type

Regular
 Target Of Opportunity
 VLBI
 Large Program
 Phased Array

Scientific Category

Cosmology and the High Redshift Universe
 Galaxies and Galactic Nuclei
 ISM, star formation and astrochemistry
 Circumstellar disks, exoplanets and the solar system
 Stellar Evolution and the Sun

Please select one or two keywords

Student project



Feedback

Validation Validation History Log

Overview

Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)

Phase I: Science Proposal





Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Project
 - ▼ Proposal
 - ▼ Planned Observing
 - ScienceGoal (Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Proposal Information

Proposal Title

Proposal Cycle

ALMA Observing Tool Reference Manual

Contents Search Favourites

- ▶ Proposal Information
- ▶ Investigators
- ▶ Reviewer Information
- ▶ Science Case
- ▶ Management Plan
- ▶ Duplicate observations
- ▶ Observatory Use Only
- ▶ Phase 1 and Phase 2 Science Goals
 - ▶ General
 - ▶ Field Setup
 - ▶ Spectral Setup
 - ▶ Calibration Setup
 - ▶ Control and Performance
 - ▶ Technical Justification
- ▶ Phase 2 Program Scheduling Blocks

Next Up Previous Contents

Next: [Investigators](#) Up: [The Phase 1 Proposal](#) Previous: [Advanced Options](#) [Contents](#)

Proposal Information



- **Proposal Title:** The title of the project can be entered here and is limited to 120 characters.
- **Proposal Cycle:** This is formed from the proposal year and the submission period. If a non-submitted proposal created during a previous cycle is read into the OT the old proposal cycle will be shown, but this will be updated to the current cycle if the project is then submitted.
- **Abstract:** The abstract can be entered as plain text and is limited to 1200 characters.
- **Proposal Type:** Four projects types are currently available

Display help from reference manual

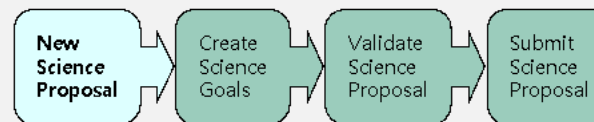


Overview

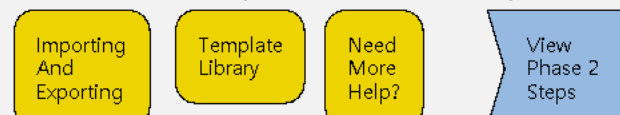
Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Proposal Information

Proposal Title

Proposal Cycle

Abstract (max. 1200 characters)

Proposal Type

Regular
 Target Of Opportunity
 VLBI
 Large Program
 Phased Array

Scientific Category

Cosmology and the High Redshift Universe
 Galaxies and Galactic Nuclei
 ISM, star formation and astrochemistry
 Circumstellar disks, exoplanets and the solar system
 Stellar Evolution and the Sun



Feedback

Validation Validation History Log

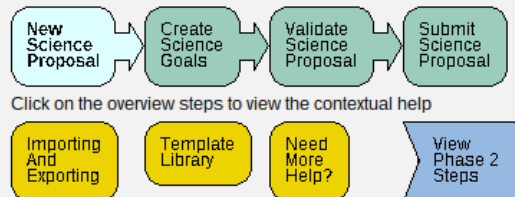
Description	Suggestion

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
- Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal





Project Structure

Editors

Proposal Program

Spectral Spatial Proposal

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)

Proposal Information

Proposal Title

Proposal Cycle

2024.1

ALMA Observing Tool Reference Manual

Contents Search Favourites

- ALMA Sensitivity Calculator
- LO Configuration Tool
- The Phase 1 Proposal: Basic Details
 - Project Node
 - Proposal Information
 - Joint Proposals
 - Investigators
 - Reviewer Information
 - Science Case
 - Team Expertise
 - Scheduling Feasibility
 - Duplicate observations
 - Observatory Use Only
- Phase 1 and Phase 2 Science Goals
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification
- Phase 2 Program Scheduling Blocks
 - Observing Unit Set
 - The Scheduling Block
 - Observing Groups
 - Targets
 - Field Sources
 - Instrument Setup
 - Observing Parameters
- The Visual Editors
 - The Visual Spatial Editor
 - The Visual Spectral Editor
- Acronym List

characters.

- **Proposal Cycle:** This is formed from the proposal year and the submission period. If a non-submitted proposal created during a previous cycle is read into the OT the old proposal cycle will be shown, but this will be updated to the current cycle if the project is then submitted.

- **Abstract:** The abstract can be entered as plain text and is limited to 1200 characters.

- **Proposal Type:** Four projects types are currently available

- Regular (Letter Code = S): This type of project is essentially the opposite of Large i.e. the total time required must be less than the criteria for Large projects.
- Large (Letter Code = L): If > 50 hours of 12-m Array or > 150 hours of 7-m

Stand-alone ACA time is requested, this Proposal Type must be selected. Projects must receive an A grade in order to receive time and are allowed a longer Scientific Justification. This proposal type is not available for ACA Supplemental Call proposals.

- Target of Opportunity (Letter Code = T): These projects will need to be triggered during the observing period. It is not necessary to enter the source coordinates as these may not be known at proposal submission time. This proposal type is not available for ACA Supplemental Call proposals.
- VLBI (Letter Code = V): Selecting this will significantly change the information that must be entered into the Science Goal. A proposal must also be submitted to the other telescopes that will form the VLBI array. This proposal type is not available for ACA Supplemental Call proposals.
- Phased Array Mode (Letter Code = P): This is very similar to the VLBI mode and requests use of the phased-up array that is used for VLBI observations. However, in this mode the ALMA antennas operate as a single instrument and are not combined with data from other telescopes.
- Once a proposal type has been selected and the proposal submitted, the proposal type cannot be changed anymore. If the proposing team wishes to submit the proposal with a different proposal type, they need to retract the current proposal via a helpdesk ticket and re-open their project as a new proposal where they then can change the proposal type.

- **Scientific Category:** A set of radio buttons for selecting the proposal's science category:

- Cosmology and the High Redshift Universe
- Galaxies and Galactic Nuclei
- ISM, star formation and astrochemistry
- Circumstellar disks, exoplanets and the solar system
- Stellar Evolution and the Sun.

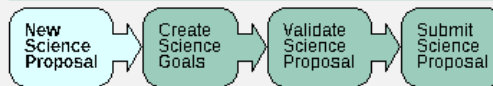
- **Keywords:** Assign at least one and preferably two keywords to the project. The second keyword is added by holding down the Shift or Ctrl key in the way that you're probably used to.

- Target Of Opportunity
- VLBI
- Phased Array

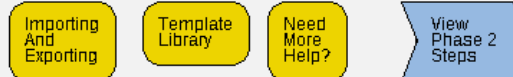
- Galaxies and Galactic Nuclei
- ISM, star formation and astrochemistry
- Stellar Evolution and the Sun

Suggestion

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Proposal Cycle: 2024.1

Abstract (max. 1200 characters)

Proposal Type:
 Regular
 Target Of Opportunity
 VLBI
 Large Program
 Phased Array

Scientific Category:
 Cosmology and the High Redshift Universe
 Galaxies and Galactic Nuclei
 ISM, star formation and astrochemistry
 Circumstellar disks, exoplanets and the solar system
 Stellar Evolution and the Sun

Please select one or two keywords:

- Low-mass star formation
- Pre-stellar cores, Infra-Red Dark Clouds (IRDC)
- Astrochemistry
- Inter-Stellar Medium (ISM)/Molecular clouds
- Photon-Dominated Regions (PDR)

Select the appropriate science category keywords, maximum 2 keywords

Student project

Joint Proposals

Is this a Joint Proposal? Yes No

Feedback

Validation Validation History Log

Description	Suggestion



Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Project
 - ▼ Proposal
 - ▼ Planned Observing
 - ▼ ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Circumstellar disks, exoplanets and the solar system
 Stellar Evolution and the Sun

Please select one or two keywords

Student project

Joint Proposals

Is this a Joint Proposal? Yes No

Investigators

Type	Full name	Email	Affiliation	ALMA ID	Executive	Reviewer
PI	Not set	Not set	Not set	Not set	Non-ALMA	<input checked="" type="checkbox"/>

Select PI

Add CoPI

Add Col

Remove Collaborator

Add from Proposal

Feedback

Validation Validation History Log

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Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)

Phase I: Science Proposal





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Spectral Spatial Proposal

Circumstellar disks, exoplanets and the solar system
 Stellar Evolution and the Sun

Please select one or two keywords

Student project

Joint Proposals

Is this a Joint Proposal? Yes No

Investigators

Type	Full name	Email	Affiliation	ALMA ID	Executive	Reviewer
PI	Not set	Not set	Not set	Not set	Non-ALMA	<input checked="" type="checkbox"/>

Investigator search constraints

Name contains Seokho Lee

Attribute to match

Find Investigators

Full name	Email	Affiliation	ALMA ID
Seokho Lee	seokholee@kasi.re.kr	Radio Astronomy Division, Kore...	shlee

Select PI Cancel



Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)

Phase I: Science Proposal





Project Structure

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Spectral Spatial Proposal

Select PI
Add CoPI
Add Col
Remove Collaborator
Add from Proposal

Reviewer Information

Please designate a reviewer who will participate in the distributed review process. The reviewer may be the PI of the proposal or one of the other investigators. A student (without a PhD) may serve as the reviewer only if they are the PI of the proposal and a mentor (with a PhD) is identified. The mentor does not need to be an investigator on the proposal.

Reviewers are requested to:

- Abide by the maximum number of Proposal Sets that are to be assigned for review to any individual (refer to the Proposer's Guide for more information).
- Update their user profiles with combinations of scientific categories and keywords which describe their area(s) of expertise using the new 'Expertise' tab in the link below. Available expertise information will be used in the distribution of proposal assignments.

<https://asa.alma.ci/UserRegistration/secure/updateAccount.jsp>

Reviewer has a PhD? No Yes

Select Mentor

Mentor name

Mentor has a PhD? No Yes

Science Case

Please ensure that your science case is properly anonymized following instructions on the Science Portal

Science Case (Mandatory, PDF, 4 pages max.)

Attach

Detach

View

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Description

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Q

Overview

Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:

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 - Calibration Setup
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Spectral Spatial Proposal

Investigator search constraints

Name contains

Find Investigators

Full name	Email	Affiliation	ALMA ID
-----------	-------	-------------	---------

Select Mentor Cancel



Overview

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1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:

Phase I: Science Proposal





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Editors

Spectral Spatial Proposal

Select PI
Add CoPI
Add Col
Remove Collaborator
Add from Proposal

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<https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>

Reviewer has a PhD? No Yes

Science Case

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Science Case (Mandatory, PDF, 4 pages max.)

Attach

Detach

View

Duplicate observations

Briefly justify any new observations that duplicate archival data or accepted programs. Information regarding the ALMA Duplication Policy and how to search archival data and accepted programs can be found at:

<https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>

Feedback

Validation Validation History Log

Description	Suggestion

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Spectral Spatial Proposal

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<https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>

Reviewer has a PhD? No Yes

Select Mentor

Mentor name

Mentor has a PhD? No Yes

12 point (<15%)
< 20 MB

Science Case ?

Please ensure that your science case is properly anonymized following instructions on the Science Portal

Science Case (Mandatory, PDF, 4 pages max.) Attach Detach View

Duplicate observations ?

Briefly justify any new observations that duplicate archival data or accepted programs.
Information regarding the ALMA Duplication Policy and how to search archival data and accepted programs can be found at:

<https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>

Observatory Use Only ? +

Feedback

Validation Validation History Log



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Editors

Spectral Spatial Proposal

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<https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>

Reviewer has a PhD? No Yes

Select Mentor

Mentor name

Mentor has a PhD? No Yes

Science Case

Please ensure that your science case is properly anonymized following instructions on the Science Portal

Science Case (Mandatory, PDF, 4 pages max.)

Attach Detach View

Duplicate observations

Briefly justify any new observations that duplicate archival data or accepted programs.
Information regarding the ALMA Duplication Policy and how to search archival data and accepted programs can be found at:

<https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>

Empty text input field for duplicate observations.

Observatory Use Only

Feedback

Validation Validation History Log

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- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)

Phase I: Science Proposal



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Spectral S

Science Case

Please ensure

Science Case (

Duplicate obs

Observatory b

Feedback

Validation

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Appendix A Definition of a Duplicate Observation

A proposed observation is considered a duplicate of another observation if *all* of the following conditions are met:

Target field location

- For single-field interferometry, the proposed position coincides within the half-power beam width of the other observation. Moving objects (e.g., Solar System objects) will be identified by name.
- For mosaic observations, more than 50% of the proposed pointings are within the half power beam width area covered by the other observation.

Angular Resolution

- The proposed angular resolution differs by a factor of ≤ 2 from the other observation.

Spectral windows

- Continuum: The requested sensitivity (rms) for the aggregate bandwidth is better by a factor of ≤ 2 from the other observation and the requested frequency is within a factor of 1.3.
- or –
- Spectral line: If the central frequency in any requested correlator window observed in Frequency Division Mode (FDM) mode is encompassed by the other observation observed in FDM mode and the sensitivity per spectral channel, after smoothing to the same spectral resolution, is better by a factor of ≤ 2 .

To be considered a “continuum” observation, the proposed correlator setup must contain 2 or more windows with a bandwidth > 1.8 GHz.

Solar observations will not be checked for duplications.



Project Structure

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Editors



Spectral Spatial Planned Observing

Summary table of Science Goals. Double-clicking on a row will take you to that Science Goal

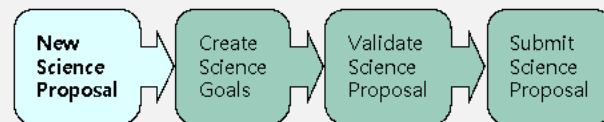
Science Goal	No. Sources	Band	Spec. Type	No. Spec. Win.	Pol.	Calibration Setup	Ang. Res.	Largest Scale	Rep. Freq.	Sens.
Science Goal	1	undefined	Spectral line	0	Dual	System	0.0 arcsec	Undefined	Undefined	Undefined

Overview

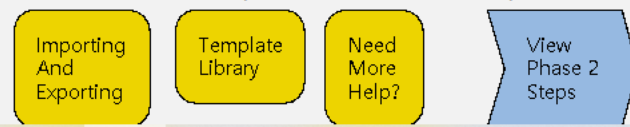
Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





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Editors

Spectral Spatial ScienceGoal (Science Goal)

General (Optional)

Science Goal Name Science Goal

Description

SinglePoint

Source

Source Name

Resolve

Choose a Solar System Object?

Name of object Unspecified

System ICRS

Sexagesimal display?

Parallax 0.00000 mas

Source Coordinates

RA 00:00:00.0000



PM RA 0.00000 mas/yr

Dec 00:00:00.000

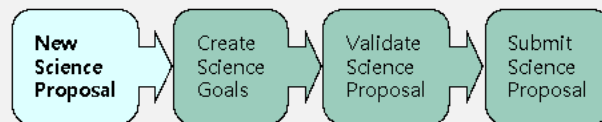
PM Dec 0.00000 mas/yr

Overview

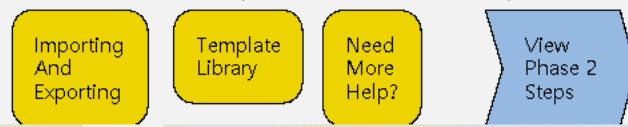
Contextual Help

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 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the  [proposal](#) tree node and complete the relevant fields.

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Proposal Program

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- Test
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- Spectral Set
- Calibration
- Control and
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Right button

- Cut Ctrl-X
- Copy Ctrl-C
- Paste Ctrl-V
- Clone node
- Show Printable Summary of ScienceGoal
- Generate a PDF of Whole Proposal
- Display Project Time Summary
- Rename
- Expand all Ctrl-Z
- Collapse all
- Find previous Alt-Up
- Find next Alt-Down
- Delete Delete

Editors

Spectral Spatial ScienceGoal (Test)

General (Optional)

Science Goal Name Test

Description

TW_Hya

Resolve

System Object?

Name of object Unspecified

System ICRS

Sexagesimal display?

Parallax 16.62890 mas

PM RA -68.30900 mas/yr

PM Dec -13.90000 mas/yr

Source Coordinates

RA 11:01:52.0913

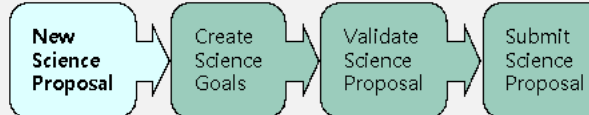
Dec -34:42:15.750

Overview

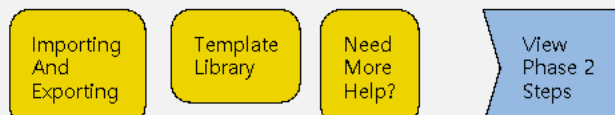
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

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Project Structure

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- General
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- Spectr
- Calibra
- Contro
- Techni

Editors

Spectral Spatial Planned Observing

Summary table of Science Goals. Double-clicking on a row will take you to that Science Goal

Science Goal	No. Sources	Band	Spec. Type	No. Spec. Win.	Pol.	Calibration Setup	Ang. Res.	Largest Scale	Rep. Freq.	Sens.
Test	1	7	Spectral line	3	Dual	System	0.015 arcsec	1.0 arcsec	345.795990 GHz	10 mJy

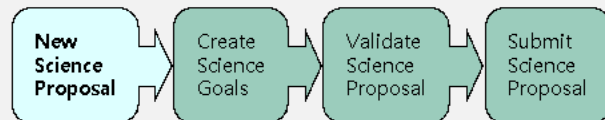
- Cut Ctrl-X
- Copy Ctrl-C
- Paste Ctrl-V
- New Science Goal
- Clone node
- Show Printable Summary of all Science Goals
- Generate a PDF of Whole Proposal
- Display Project Time Summary
- Expand all Ctrl-Z
- Collapse all
- Find previous Alt-Up
- Find next Alt-Down
- Delete Delete

Overview

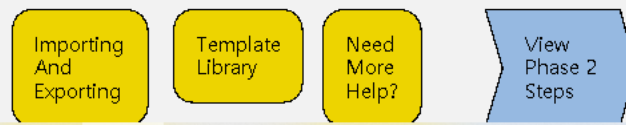
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the **1** icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

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 - Technical Justification

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Spectral Spatial Planned Observing

Summary table of Science Goals. Double-clicking on a row will take you to that Science Goal

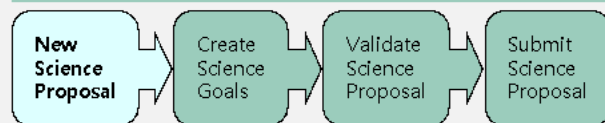
Science Goal	No. Sources	Band	Spec. Type	No. Spec. Win.	Pol.	Calibration Setup	Ang. Res.	Largest Scale	Rep. Freq.	Sens.
Test	1	7	Spectral line	3	Dual	System	0.015 arcsec	1.0 arcsec	345.795990 GHz	10 mJy

Overview

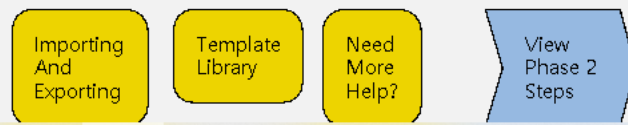
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
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 - Clicking on the **1** icon in the toolbar
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Phase I: Science Proposal



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 - Spectral Setup
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Editors

Spectral Spatial Field Setup

Spatial Image

Image Filename

TW Hya

Source

Source Name

TW Hya

Resolve

Choose a Solar System Object?

Name of object Unspecified

Get source details from SIMBAD or NED

Source Coordinates

System ICRS

Sexagesimal display?

Parallax 0.00000

mas

RA 00:00:00.0000

PM RA 0.00000

mas/yr

Dec 00:00:00.0000

PM Dec 0.00000

mas/yr

Source Radial Velocity

0.000

km/s

Isrk

z 0.000000000

Doppler Type

RADIO

Target Type

 Individual Pointing(s)
 1 Rectangular Field

Expected Source Properties

Peak Continuum Flux Density per Synthesized Beam

0.00000

Jy

Continuum Linear Polarization

0.0

per cent

Continuum Circular Polarization

0.0

per cent

Peak Line Flux Density per Synthesized Beam

0.00000

Jy

Line Width

0.00000

km/s

Line Linear Polarization

0.0

per cent

Line Circular Polarization

0.0

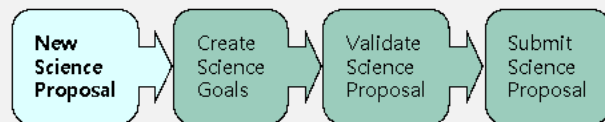
per cent

Overview

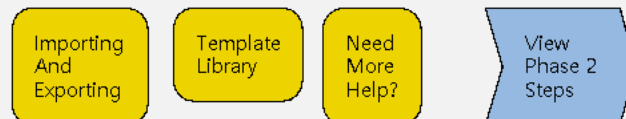
Contextual Help

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3. Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



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Project Structure

Proposal Program

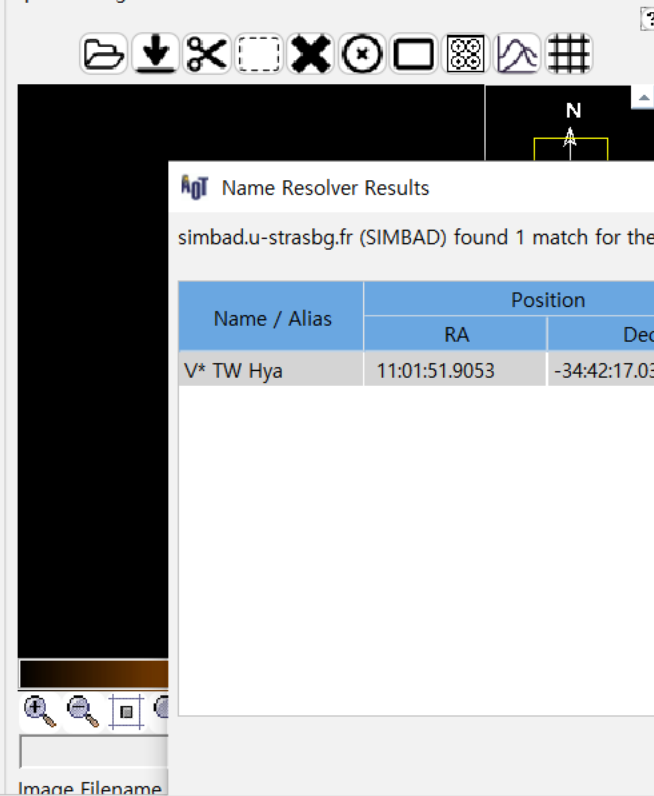
Unsubmitted Proposal

- Project
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 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Field Setup

Spatial Image



TW Hya

Source

Source Name

TW Hya

Choose a Solar System Object?

Name of object

Unspecified

Resolve

Name Resolver Results

simbad.u-strasbg.fr (SIMBAD) found 1 match for the object 'TW Hya'.

Name / Alias	Position		Proper Motion		Velocity
	RA	Dec	RA	Dec	
V* TW Hya	11:01:51.9053	-34:42:17.033	-68.309 mas/yr	-13.9 mas/yr	12.335 km/s



Cancel

Select

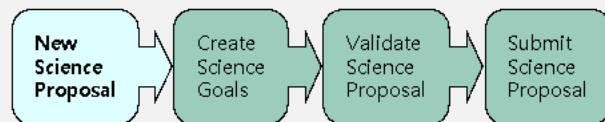
Image Filename

Overview

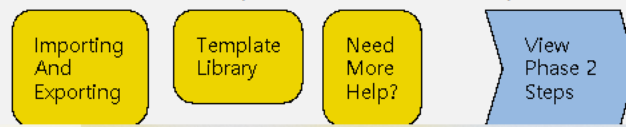
Contextual Help

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 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help



You should check that the "Resolved information" is correct!!!

Project Structure

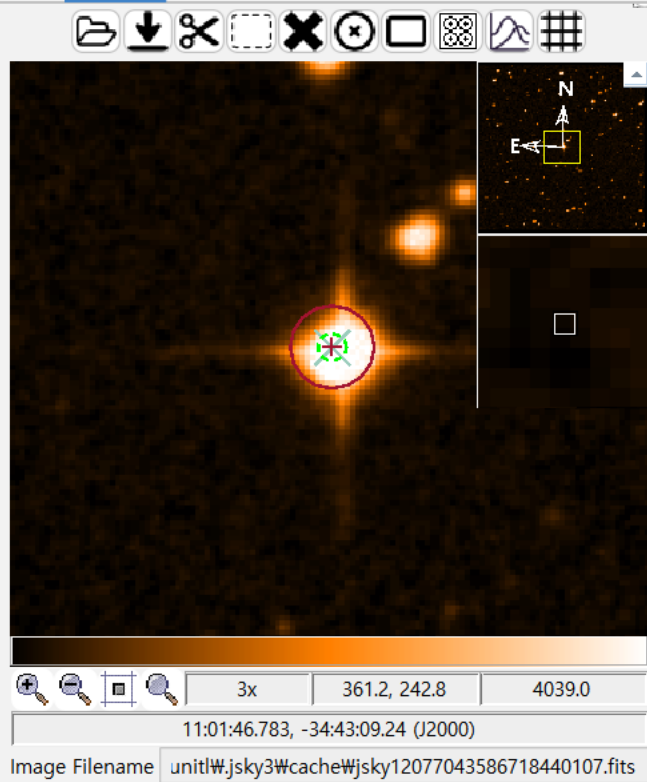
Proposal Program

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 - Spectral Setup
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Spectral Spatial Field Setup



FOV Parameters

Source

Source Name Resolve

Choose a Solar System Object? Name of object

System Sexagesimal display?

Source Coordinates

RA	<input type="text" value="11:01:52.0913"/>	Parallax	<input type="text" value="16.62890"/>	<input type="text" value="mas"/>
Dec	<input type="text" value="-34:42:15.750"/>	PM RA	<input type="text" value="-68.30900"/>	<input type="text" value="mas/yr"/>
		PM Dec	<input type="text" value="-13.90000"/>	<input type="text" value="mas/yr"/>

Source Radial Velocity z Doppler Type

Target Type Individual Pointing(s) 1 Rectangular Field

Expected Source Properties

Peak Continuum Flux Density per Synthesized Beam	<input type="text" value="50.00000"/>	<input type="text" value="mJy"/>
Continuum Linear Polarization	<input type="text" value="0.0"/>	<input type="text" value="per cent"/>
Continuum Circular Polarization	<input type="text" value="0.0"/>	<input type="text" value="per cent"/>
Peak Line Flux Density per Synthesized Beam	<input type="text" value="10.00000"/>	<input type="text" value="mJy"/>
Line Width	<input type="text" value="1.00000"/>	<input type="text" value="km/s"/>
Line Linear Polarization	<input type="text" value="0.0"/>	<input type="text" value="per cent"/>
Line Circular Polarization	<input type="text" value="0.0"/>	<input type="text" value="per cent"/>

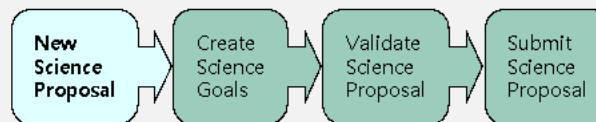
Field Centre Coordinates

Overview

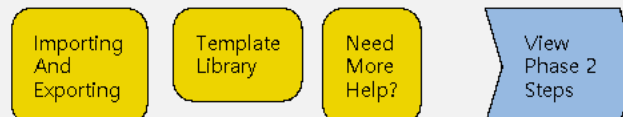
Contextual Help

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Phase I: Science Proposal



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 - Calibration Setup
 - Control and Performance
 - Technical Justification
 - ScienceGoal (Copy of Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

3x 361.2, 242.8 4039.0

11:01:46.783, -34:43:09.24 (J2000)

Image Filename unit\W\jsky3\wcache\Wjsky12077043586718440107.fits

Source

Source Name TW Hya

Source Name dropdown: ICRS (selected), ICRS, FK5 J2000, galactic, ecliptic, horizon, azel

System ICRS

RA 11:01:52.0913

Dec -34:42:15.750

Velocity: 12.335 km/s, hel

Velocity dropdown: lsrk (selected), bar, lsrk, topo, hel

Parallax 16.62890 mas

PM RA -68.30900 mas/yr

PM Dec -13.90000 mas/yr

Doppler Type RELATIVISTIC

Continuum Flux Density per Synthesized Beam 5

Continuum Linear Polarization 0.0 per cent

Continuum Circular Polarization 0.0 per cent

Peak Line Flux Density per Synthesized Beam 10.00000 mJy

Line Width 1.00000 km/s

Line Linear Polarization 0.0 per cent

Line Circular Polarization 0.0 per cent

ALMA OT - Information

Changed coordinate system type from "ICRS" to "J2000", but no coordinate conversion was performed

Do not show this message again

OK

ALMA OT - Information

Changed velocity reference system type from "lsrk" to "hel", but no velocity conversion was performed

Do not show this message again

OK

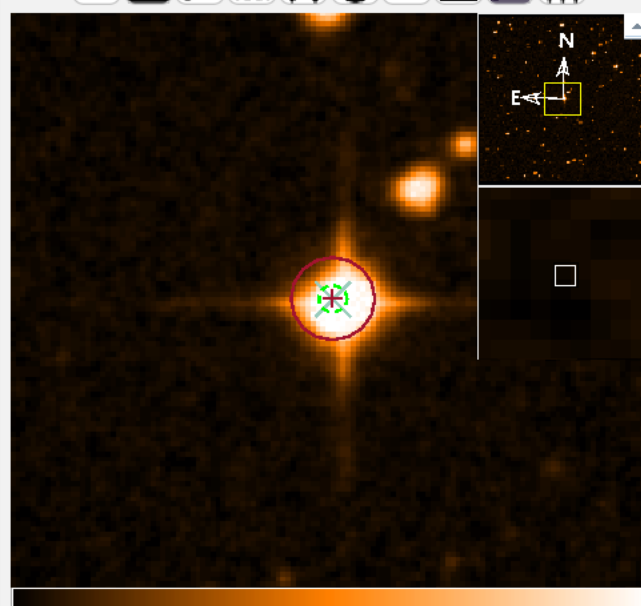
File Edit View Tool Search Help

Project Structure

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification
 - ScienceGoal (Copy of Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification
 - ScienceGoal (Copy of Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Field Setup



Source

Source Name: TW Hya



Source Coordinates: RA: 11:01:52.0913, Dec: -34:42:15.750

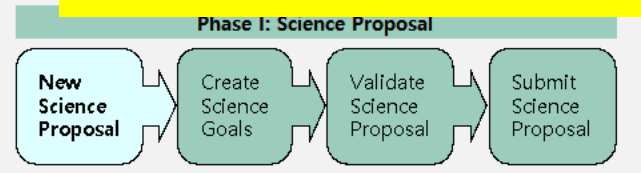
Source Radial Velocity: 12.335 km/s

Expected Source Properties:

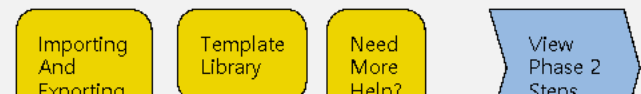
Peak Continuum Flux Density per Synthesized Beam	50.00000	mJy
Continuum Linear Polarization	0.0	per cent
Continuum Circular Polarization	0.0	per cent
Peak Line Flux Density per Synthesized Beam	10.00000	mJy
Line Width	1.00000	km/s
Line Linear Polarization	0.0	per cent
Line Circular Polarization	0.0	per cent

You should describe how to get these values in T.J.

- Contextual Help**
- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
 - Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
 - Click on the  [proposal](#) tree node and complete the relevant fields.



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

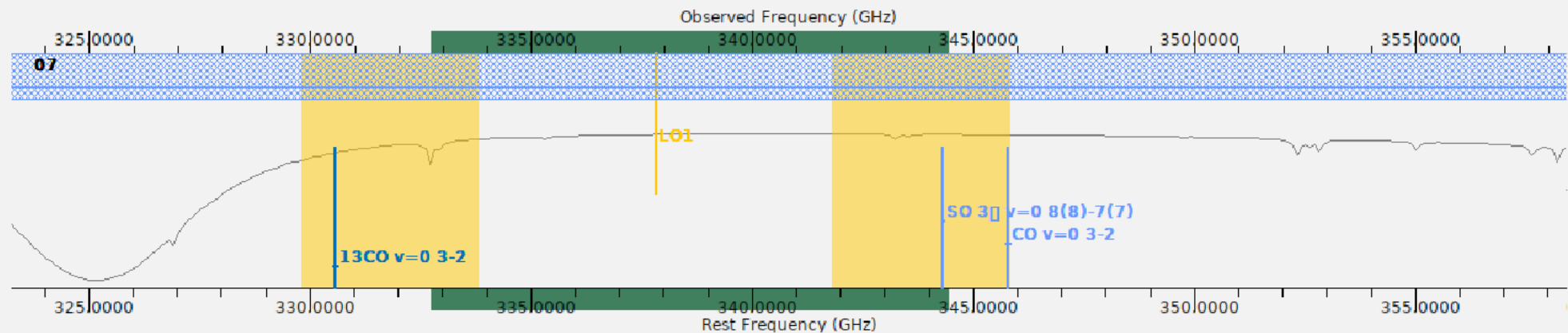
Spectral Spatial Spectral Setup

Visualisation

In the table below, it is possible to define up to 16 spectral windows, 4 per baseband as long as the total Fraction per baseband is no more than 1. Each baseband is 2GHz wide and can be separately configured i.e. each spectral window can have a different bandwidth and resolution. Note that for bands 3 to 8, it is not possible to put 3 basebands in one sideband and the fourth one in the other.

Left/right click to zoom in/out, grab sliding bar to pan

Note: Moving LO1 here is for experimentation only - actual setup determined by the windows



Overlays: Receiver Bands Transmission DSB Image Spectral Lines Select Lines to Overlay

Water Vapour Column Density: Automatic Ch Toggle receiver bands overlay (4th Octile)

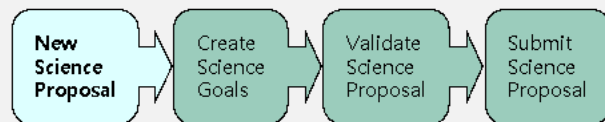
Viewport: Pan to Spectral Window Zoom to Band Reset

Overview

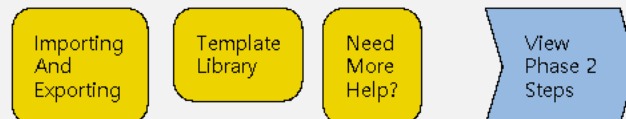
Contextual Help

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2. Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Accretion Instabilities in EC 53
 - ▼ Proposal
 - ▼ Planned Observing
 - > ScienceGoal (B6 C180 + SO + SiO)
 - ▼ ScienceGoal (Copy of B6 C180 + S
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Spectral Setup

Water Vapour Column Density: Automatic Choice Manual Choice 0.913mm (3rd Octile) ▾

Viewport:

Pan to Spectral Window

Zoom to Band

Reset

Spectral Type

Spectral Type

-
- Spectral Line
-
-
- Single Continuum
-
-
- Spectral Scan

1. Continuum Obs.

Produce image sidebands (Bands 9 and 10 only)

Polarization products desired

-
- XX
-
- DUAL
-
- FULL

Spectral Setup Errors

Single Continuum

Receiver Band 7 [275.0-373.0 GHz] ▾

Reset to Standard Frequency

Sky Frequency 343.50000 GHz ▾

Rest Frequency 343.500720 GHz

-
- Low spectral resolution (TDM)
-
-
- High spectral resolution (FDM)

2 Band (Freq.)

3. TDM vs. FDM

Baseband-1

Fraction	Centre Freq (rest,topo)	Centre Freq (sky,topo)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1(Full)	336.50954 GHz	336.50000 GHz	Single Continuum	1875.000 MHz(1670 km/s), 1.129 MHz(1.006 km/s) (2-bit)	2	<input type="radio"/>

Feedback

Validation

Validation History

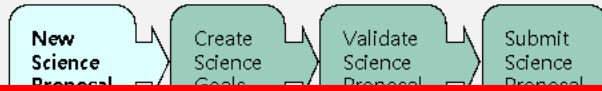
Log

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal

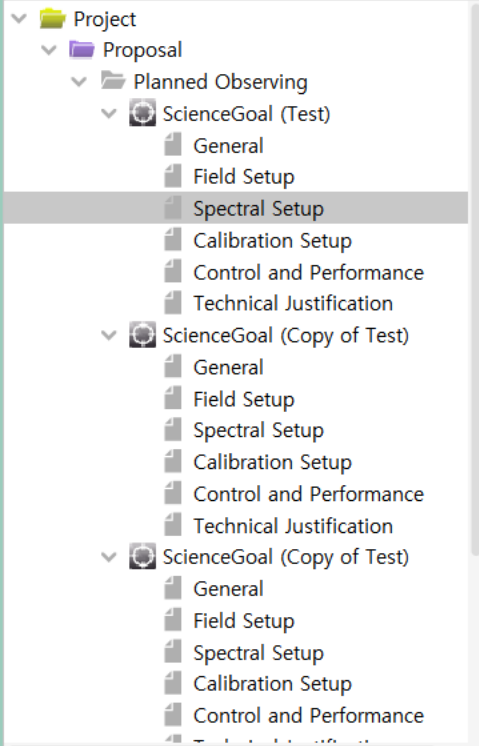




Project Structure

Proposal Program

Unsubmitted Proposal



Overview

Editors

Spectral Spatial Spectral Setup

Line Obs.

Spectral Type

Spectral Type

- Spectral Line
 Single Continuum
 Spectral Scan

Line survey at a point

Highest spectral resolution

Produce image sidebands (Bands 9 and 10 only)

Polarization products desired

- XX DUAL FULL

polarization

Spectral Setup Errors

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1/2	345.79599 GHz	345.78176 GHz	CO v=0 3-2	234.375 MHz(203 km/s), 244.141 kHz(0.212 km/s)	1	<input checked="" type="radio"/>
1/4	344.31061 GHz	344.29645 GHz	SO 3Σ v=0 8(8)-7(7)	117.188 MHz(102 km/s), 282.227 kHz(0.246 km/s)	2	<input type="radio"/>

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

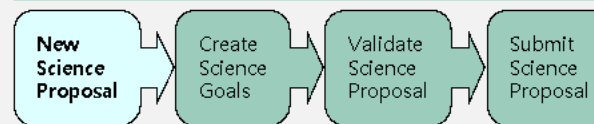
Baseband-2

1(Full)	330.58797 GHz	330.57436 GHz	13CO v=0 3-2	234.375 MHz(213 km/s), 141.113 kHz(0.128 km/s)	2	<input type="radio"/>
---------	---------------	---------------	--------------	--	---	-----------------------

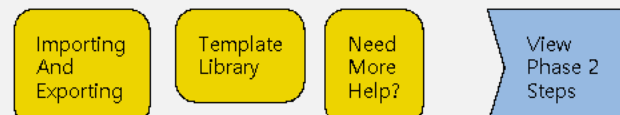
Contextual Help

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Phase I: Science Proposal



Click on the overview steps to view the contextual help



Create spectral windows centred on spectral lines

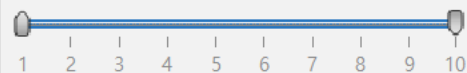
Transition Filter

*
e.g. CO*2-1* or *oxide*

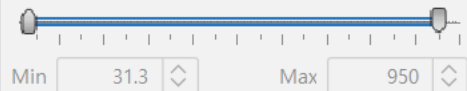
Include description

Frequency Filters

ALMA Band



Sky Frequency (GHz)



Receiver/Back End Configuration

- All lines
- Potentially selectable lines
- Lines in defined spws
- Filtering unobservable lines

Upper-state Energy (K)

Min 0 Max 0

Molecule Filter / Environment

- Show all atoms and molecules
- all atoms and molecules
 - most common molecules
 - hot cores
 - dark clouds
 - diffuse clouds
 - AGBs, PPNs and PNs
 - comets
 - planets
 - extra-Galactic

Transitions matching your filter settings:

(double-click column header for primary sort, single-click subsequent columns for secondary sorting. Single clicks will reverse sort order of already selected columns.)

Transition ▲	Description	Rest Frequency ▲	Sky Frequency	Upper-state Energy	Lovas Intensity	Sij μ ²	Catalog
CH3CN v=0 18(4)-17(4), F=17-16	Methyl Cyanide	330.969808 GHz	330.956190 GHz	265.219 K	1.38	496.315 D ²	Offline
CH3CN v=0 18(4)-17(4), F=17-18	Methyl Cyanide	330.969812 GHz	330.956195 GHz	265.219 K	1.38	0.001 D ²	Offline
CH3CN v=0 18(4)-17(4), F=19-18	Methyl Cyanide	330.969815 GHz	330.956198 GHz	265.219 K	1.38	554.827 D ²	Offline
CH3CN v8=1 J =18-17, K = -12 --12	Methyl Cyanide	330.977817 GHz	330.964199 GHz	1861.315 K		286.041 D ²	Offline
CH3CN v8=1 J =38-38, K =10-8	Methyl Cyanide	330.988159 GHz	330.974541 GHz	1758.834 K		0.017 D ²	Offline
CH3CN v8=1 J =18-17, K =14-14	Methyl Cyanide	331.009015 GHz	330.995396 GHz	1881.454 K		203.422 D ²	Offline
CH3CN v=0 18(3)-17(3), F=18-17	Methyl Cyanide	331.014296 GHz	331.000677 GHz	215.24 K	1.38	1073.219 D ²	Offline
CH3CN v=0 18(3)-17(3), F=17-16	Methyl Cyanide	331.014306 GHz	331.000687 GHz	215.24 K	1.38	1015.053 D ²	Offline
CH3CN v=0 18(3)-17(3), F=19-18	Methyl Cyanide	331.014315 GHz	331.000695 GHz	215.24 K	1.38	1134.981 D ²	Offline
CH3CN v=0 18(3)-17(3), F=17-18	Methyl Cyanide	331.014315 GHz	331.000695 GHz	215.24 K	1.38	0.003 D ²	Offline
CH3CN v=0 18(2)-17(2), F=18-17	Methyl Cyanide	331.046102 GHz	331.032481 GHz	179.533 K	1.6	545.146 D ²	Offline
CH3CN v=0 18(2)-17(2), F=17-16	Methyl Cyanide	331.046104 GHz	331.032483 GHz	179.533 K	1.6	515.6 D ²	Offline
CH3CN v=0 18(2)-17(2), F=19-18	Methyl Cyanide	331.046113 GHz	331.032492 GHz	179.533 K	1.6	576.385 D ²	Offline
CH3CN v=0 18(2)-17(2), F=17-18	Methyl Cyanide	331.046115 GHz	331.032494 GHz	179.533 K	1.6	0.001 D ²	Offline
CH3CN v=0 18(1)-17(1), F=17-16	Methyl Cyanide	331.065188 GHz	331.051566 GHz	158.106 K	1.64	520.407 D ²	Offline
CH3CN v=0 18(1)-17(1), F=18-17	Methyl Cyanide	331.065191 GHz	331.051570 GHz	158.106 K	1.64	550.355 D ²	Offline
CH3CN v=0 18(1)-17(1), F=19-18	Methyl Cyanide	331.065197 GHz	331.051576 GHz	158.106 K	1.64	581.893 D ²	Offline
CH3CN v=0 18(1)-17(1), F=17-18	Methyl Cyanide	331.065201 GHz	331.051579 GHz	158.106 K	1.64	0.001 D ²	Offline
CH3CN v=0 18(0)-17(0), F=17-18	Methyl Cyanide	331.071563 GHz	331.057942 GHz	150.963 K	1.77	0.001 D ²	Offline

Add to spectral window list

Spectral windows in this baseband (maximum of four)

Transition ▲	Description	Rest Frequency ▲	Sky Frequency

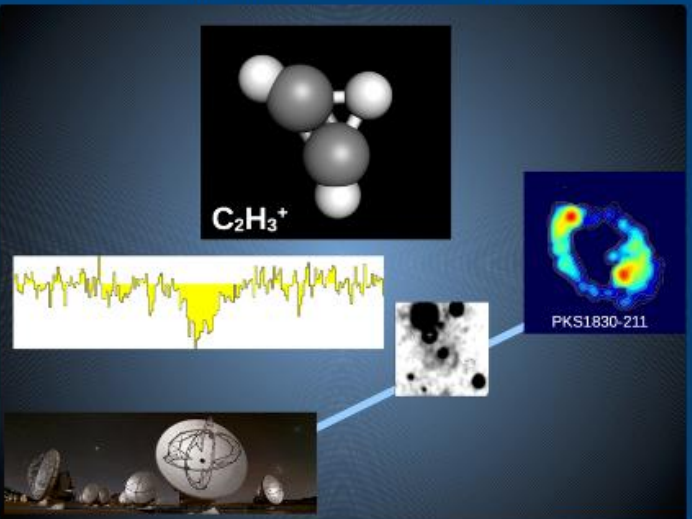
Remove spectral window(s)

Cancel Ok



Science Highlight

Protonated acetylene in the z=0.89 absorber toward PKS1830-211



View of the molecule (created with MolView) and an absorption spectrum observed with ALMA in the line of sight of the quasar PKS1830-211. The quasar (here observed with the MERLIN interferometer at radio wavelengths) is lensed by a foreground spiral galaxy at z=0.89 (optical image from HST).

The line of sight to the lensed blazar PKS1830-211 intercepts the disk of a foreground spiral galaxy at z=0.89 where absorption has been detected for more than 60 molecular species, mostly at mm wavelengths. In a paper accepted for publication in A&A. Dr. Sebastian Muller and colleagues report the detection of a new

Observatory News

ALMA Cycle 11 Call for Proposals OPEN!
Mar 21, 2024

Planned Release of Data from Projects in Configurations 8 and 9
Mar 21, 2024

Over one third of all ALMA proposals make use of data from the ALMA Archive
Feb 16, 2024

Announcement for early proposal for Cycle 11

[More...](#)

The ALMA Science Portal is a one-stop resource for ALMA staff, journalists, and funding agencies.

Quick Links

- [ALMA Basics](#)
- [ALMA Science](#)
- [ALMA Primer](#)

- Observing Tool
- Sensitivity Calculator
- CASA Simulator
- ALMA Primer Instructional Videos
- Observation Support Tool
- Splatalogue**
- NRAO Science Ready Data Products
- Toyama Microwave Atlas
- Community-Developed
- EU ARC network
- Staff Tools
- Japanese Virtual Obs.
- Solar Ephemeris

ALMA Status

Configuration Schedule

Refereed publications: 3720
Last observed source: XID614
Current configuration: C-1

[More...](#)

...whole, including proposers, archive researchers, ALMA

- [SnooPI](#)
- [DDT Proposals](#)

Splatalogue

Database for Astronomical Spectroscopy
Giving you the right frequency one line at a time.

[Basic Version](#)

[Advanced Version](#)

[FAQs](#)



Photo Credit: NRAO/AUI/NSF

Search Species

e.g. ammonia, carbon, methanol, CO

Search Parameters

Quick Picker

CO v=0	<input checked="" type="checkbox"/>	¹³ CO v=0	<input type="checkbox"/>
C ¹⁷ O	<input type="checkbox"/>	C ¹⁸ O	<input type="checkbox"/>
CH ₃ OH v ₁ =0	<input type="checkbox"/>	H ₂ CO	<input type="checkbox"/>
HCN v=0	<input type="checkbox"/>	HNC v=0	<input type="checkbox"/>
H ¹³ CN v=0	<input type="checkbox"/>	HC ¹⁵ N v=0	<input type="checkbox"/>
DCN v=0	<input type="checkbox"/>	HCO ⁺ v=0	<input type="checkbox"/>
CS	<input type="checkbox"/>	H ¹³ CO ⁺	<input type="checkbox"/>
NH ₃	<input type="checkbox"/>	C I	<input type="checkbox"/>
C II	<input type="checkbox"/>	O I	<input type="checkbox"/>
O III	<input type="checkbox"/>	N II	<input type="checkbox"/>
H ₂ O v=0	<input type="checkbox"/>	HDO	<input type="checkbox"/>
SiO v=0	<input type="checkbox"/>	More molecules	

Save and Restore Settings

Settings Name

Query Filters

Telescope Bands

ALMA Band 5 (183-211 GHz)

ALMA Band 6 (211-275 GHz)

ALMA Band 7 (275-373 GHz)

ALMA Band 8 (385-500 GHz)

Energy Range

From To

E_L (cm⁻¹) E_L (K)

Specify a Frequency Range

From To

Frequency Unit

Load Freqs From listobs

No file chosen

Load Freqs From ObsProposal.xml

No file chosen

Redshift

Astronomical Filters

None

Top 20 list

Planetary Atmosphere

Hot Cores

Dark Clouds

Diffuse Clouds

Comets

AGB/PPNP/N

Extragalactic

Export Data

Export Fields

Export current fields

Export current fields without Resolved QNs

Export CASA fields

Field Separator

Tab

Colon

Range

All Records

Current Page

Search Results

Transition Filter

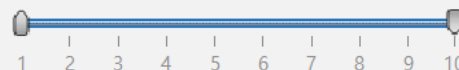
co v=0*

e.g. CO*2-1* or *oxide*

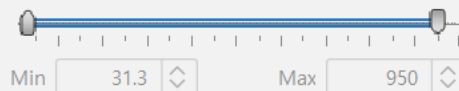
 Include description

Frequency Filters

ALMA Band



Sky Frequency (GHz)



Receiver/Back End Configuration

 All lines Potentially selectable lines Lines in defined spws Filtering unobservable lines

Upper-state Energy (K)

Min Max

Molecule Filter / Environment

Show

Can't find the transition you're looking for in the offline pool? Find more in the online Splatalogue.

Transitions matching your filter settings:

(double-click column header for primary sort, single-click subsequent columns for secondary sorting. Single clicks will reverse sort order of already selected columns.)

Transition ▲	Description	Rest Freque... ▲	Sky Frequency	Upper-state Ene...	Lovas Inten...	Sij μ^2	Catalo...
CO v=0 3-2	Carbon Monoxide	345.795990 GHz	345.781762 GHz	33.192 K	70	0.036 D ²	Offline

Spectral windows in this baseband (maximum of four)

Transition ▲	Description	Rest Frequency ▲	Sky Frequency
CO v=0 3-2	Carbon Monoxide	345.795990 GHz	345.781762 GHz

3. Click on the [proposal](#) tree node and complete the relevant fields.Importing
And
ExportingTemplate
LibraryNeed
More
Help?View
Phase 2
Steps



Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Project
 - ▼ Proposal
 - ▼ Planned Observing
 - ▼ ScienceGoal (Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Spectral Setup

Spectral type Single Continuum
 Spectral Scan

Produce image sidebands (Bands 9 and 10 only)

Polarization products desired XX DUAL FULL

Spectral Setup Errors

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1(Full) ▼	345.79599 GHz	345.78176 GHz	CO v=0 3-2	117.188 MHz(102 km/s), 70.557 kHz(0.061 km/s)	2	<input checked="" type="radio"/>
1(Full)						<input type="radio"/>
1/2						<input type="radio"/>
1/4						<input type="radio"/>

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-2

Add spectral window centred on a spectral line

Add spectral window manually

Delete

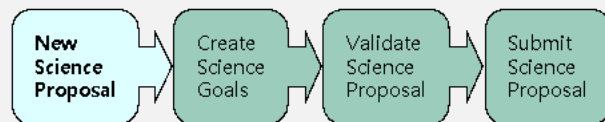
 Show image spectral windows

Overview

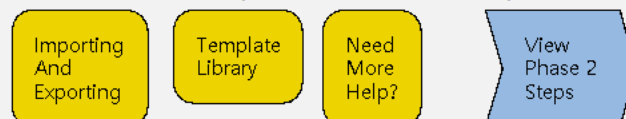
Contextual Help

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 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Fraction 1(Full)

Editors

Spectral Spatial Spectral Setup

Spectral Type

- Spectral Type
- Spectral Line
 Single Continuum
 Spectral Scan
- Produce image sidebands (Bands 9 and 10 only)
- Polarization products desired XX DUAL FULL

Spectral Setup Errors

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1(Full)	230.53800 GHz	230.52851 GHz	CO v=0 2-1	58.594 MHz (76 km/s), 35.278 kHz (0.046 km/s) (2-bit)	2	<input checked="" type="radio"/>
				58.594 MHz (76 km/s), 141.113 kHz (0.184 km/s) (4-bit)		
				117.188 MHz (152 km/s), 70.557 kHz (0.092 km/s) (2-bit)		
				117.188 MHz (152 km/s), 282.227 kHz (0.367 km/s) (4-bit)		
				234.375 MHz (305 km/s), 141.113 kHz (0.184 km/s) (2-bit)		
				234.375 MHz (305 km/s), 564.453 kHz (0.734 km/s) (4-bit)		
				468.750 MHz (610 km/s), 282.227 kHz (0.367 km/s) (2-bit)		
				468.750 MHz (610 km/s), 1.129 MHz (1.468 km/s) (4-bit)		
				937.500 MHz (1219 km/s), 564.453 kHz (0.734 km/s) (2-bit)		
				937.500 MHz (1219 km/s), 2.258 MHz (2.936 km/s) (4-bit)		
				1875.000 MHz (2438 km/s), 1.129 MHz (1.468 km/s) (2-bit)		
				1875.000 MHz (2438 km/s), 36.125 MHz (46.979 km/s) (2-bit)		

Add spectral window centred on a spectral line Add spectral window manually Delete Show image spectral windows

Baseband-2

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window

Add spectral window centred on a spectral line Add spectral window manually Delete Show image spectral windows

Feedback

Validation Validation History Log

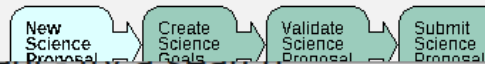
Q

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Fraction 1/2

Editors

Spectral Spatial Spectral Setup

Spectral Type

- Spectral Type
- Spectral Line
 Single Continuum
 Spectral Scan
- Produce image sidebands (Bands 9 and 10 only)
- Polarization products desired XX DUAL FULL

Spectral Setup Errors

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1/2	230.53800 GHz	230.52851 GHz	CO v=0 2-1	58.594 MHz (76 km/s), 70.557 kHz (0.092 km/s) (2-bit)	2	<input checked="" type="radio"/>
				58.594 MHz (76 km/s), 70.557 kHz (0.092 km/s) (2-bit)		
				117.188 MHz (152 km/s), 141.113 kHz (0.184 km/s) (2-bit)		
				234.375 MHz (305 km/s), 282.227 kHz (0.367 km/s) (2-bit)		
				468.750 MHz (610 km/s), 564.453 kHz (0.734 km/s) (2-bit)		
				937.500 MHz (1219 km/s), 1.129 MHz (1.468 km/s) (2-bit)		

Baseband-2

Add spectral window centred on a spectral line
 Add spectral window manually

 Show image spectral windows

Baseband-3

Add spectral window centred on a spectral line
 Add spectral window manually

 Show image spectral windows

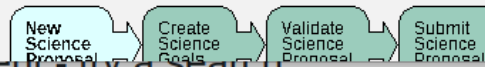
Feedback

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal





Project Structure

Proposal Program

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Fraction 1/4

Editors

Spectral Spatial Spectral Setup

Spectral Type

Spectral Type

Spectral Line
 Single Continuum
 Spectral Scan

Produce image sidebands (Bands 9 and 10 only) Polarization products desired XX DUAL FULL

Spectral Setup Errors

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1/4	230.53800 GHz	230.52851 GHz	CO v=0 2-1	58.594 MHz (76 km/s), 141.113 kHz (0.184 km/s) (2-bit)	2	<input checked="" type="radio"/>
				58.594 MHz (76 km/s), 141.113 kHz (0.184 km/s) (2-bit)		
				117.188 MHz (152 km/s), 282.227 kHz (0.367 km/s) (2-bit)		
				234.375 MHz (305 km/s), 564.453 kHz (0.734 km/s) (2-bit)		
				468.750 MHz (610 km/s), 1.129 MHz (1.468 km/s) (2-bit)		

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-2

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-3

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Feedback

Validation Validation History Log

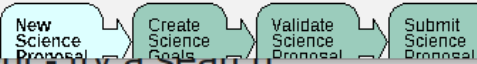
Q

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal



Create spectral windows centred on spectral lines

Transition Filter

Proposal: *

e.g. CO*2-1* or *oxide*

Include description

Frequency Filters

ALMA Band:

Sky Frequency (GHz):

Min: 31.3 Max: 950

Receiver/Back End Configuration

All lines

Potentially selectable lines

Lines in defined spws

Filtering unobservable lines

Upper-state Energy (K)

Min: 0 Max: 0

Molecule Filter / Environment

Show: all atoms and molecules

Can't find the transition you're looking for in the offline pool? Find more in the online Splatalogue.

Transitions matching your filter settings:

(double-click column header for primary sort, single-click subsequent columns for secondary sorting. Single clicks will reverse sort order of already selected columns.)

Transition ▲	Description	Rest Frequency ▼	Sky Frequency ▲	Upper-state Energy	Lovas Intensity ▼	Sij μ ²	Catalog
13CH3OH v t=0 5 (...)	Methanol	349.222565 GHz	349.208196 GHz	56.339 K		0 D ²	Offline
CH3CN v=0 19(6)-1...	Methyl Cyanide	349.212338 GHz	349.197970 GHz	424.7 K	0.71	1106.34 D ²	Offline
CH3CN v=0 19(6)-1...	Methyl Cyanide	349.212332 GHz	349.197964 GHz	424.7 K	0.71	995.379 D ²	Offline
CH3CN v=0 19(6)-1...	Methyl Cyanide	349.212321 GHz	349.197953 GHz	424.7 K	0.71	0.002 D ²	Offline
CH3CN v=0 19(6)-1...	Methyl Cyanide	349.212285 GHz	349.197917 GHz	424.7 K	0.71	1049.515 D ²	Offline
CH3CN v8=1 J =19...	Methyl Cyanide	349.208926 GHz	349.194558 GHz	2089.502 K		204.763 D ²	Offline
CH3CN v8=1 J =19...	Methyl Cyanide	349.174395 GHz	349.160028 GHz	2068.707 K		289.09 D ²	Offline
CH3CN v=0 19(7)-1...	Methyl Cyanide	349.125319 GHz	349.110955 GHz	517.407 K	0.5	531.025 D ²	Offline
CH3CN v=0 19(7)-1...	Methyl Cyanide	349.125315 GHz	349.110950 GHz	517.407 K	0.5	477.876 D ²	Offline
CH3CN v=0 19(7)-1...	Methyl Cyanide	349.125298 GHz	349.110934 GHz	517.407 K	0.5	0.001 D ²	Offline
CH3CN v=0 19(7)-1...	Methyl Cyanide	349.125249 GHz	349.110884 GHz	517.407 K	0.5	503.75 D ²	Offline
CH3OH v t=0 14(1,...)	Methanol	349.107020 GHz	349.092656 GHz	260.203 K	3.52	25.799 D ²	Offline
13CH3OH v t=0 4 (...)	Methanol	349.097921 GHz	349.083558 GHz	45.01 K		0 D ²	Offline
13CH3OH v t=0 3 (...)	Methanol	349.034424 GHz	349.020063 GHz	35.947 K		0 D ²	Offline
CH3CN v=0 19(8)-1...	Methyl Cyanide	349.025009 GHz	349.010648 GHz	624.32 K	1.03	505.523 D ²	Offline
CH3CN v=0 19(8)-1...	Methyl Cyanide	349.025006 GHz	349.010646 GHz	624.32 K	1.03	454.821 D ²	Offline
CH3CN v=0 19(8)-1...	Methyl Cyanide	349.024983 GHz	349.010623 GHz	624.32 K	1.03	0.001 D ²	Offline
CH3CN v=0 19(8)-1...	Methyl Cyanide	349.024918 GHz	349.010558 GHz	624.32 K	1.03	479.558 D ²	Offline
CH3CN v8=1 J =19...	Methyl Cyanide	349.016636 GHz	349.002276 GHz	2294.548 K		316.133 D ²	Offline

Add to spectral window list

Spectral windows in this baseband (maximum of four)

Transition ▲	Description	Rest Frequency ▲	Sky Frequency
SO 3Σ v=0 8(8)-7(7)	Sulfur Monoxide	344.310612 GHz	344.296446 GHz
CO v=0 3-2		345.795990 GHz	345.781762 GHz
SiO v=0 8-7	Silicon Monoxide	347.330579 GHz	347.316288 GHz

Remove spectral window(s)

Cancel Ok

Click on the [proposal](#) tree node and complete the relevant fields.

Importing And Exporting | Template Library | Need More Help? | View Phase 2 Steps



Project Structure

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Editors

Spectral Spatial Spectral Setup

Produce image sidebands (Bands 9 and 10 only) Polarization products desired XX DUAL FULL

Spectral Setup Errors

The spectral window range exceeds the baseband width : 3.051092345699999 GHz

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1/2	345.79599 GHz	345.78176 GHz	CO v=0 3-2	58.594 MHz, 70.557 kHz	2	<input checked="" type="radio"/>
1/4	344.31061 GHz	344.29645 GHz	SO 3Σ v=0 8(8)-7(7)	58.594 MHz, 141.113 kHz	2	<input type="radio"/>
1/4	347.33058 GHz	347.31628793178334 GHz	SiO v=0 8-7	Please select a correlator mode	1	<input type="radio"/>

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-2

Error : over the width of the baseband

Add spectral window centred on a spectral line

Add spectral window manually

Delete

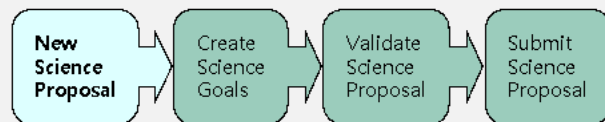
 Show image spectral windows

Overview

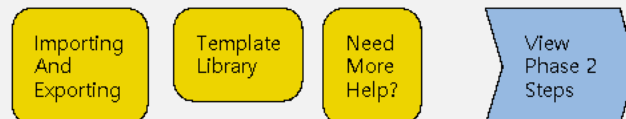
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





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Editors

Spectral Spatial Spectral Setup

Produce image sidebands (Bands 9 and 10 only) Polarization products desired XX DUAL FULL

Spectral Setup Errors

Baseband-1 : Spectral window resolution mismatch in spectral set-up. All windows must be allocated the same resolution.

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1/2	345.79599 GHz	345.78176 GHz	CO v=0 3-2	58.594 MHz(51 km/s), 70.557 kHz(0.061 km/s)	2	<input checked="" type="radio"/>
1/4	344.31061 GHz	344.29645 GHz	SO 3Σ v=0 8(8)-7(7)	58.594 MHz(51 km/s), 141.113 kHz(0.123 km/s)	2	<input type="radio"/>

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-2

Error : different spectral resolution within the baseband

Add spectral window centred on a spectral line



Add spectral window manually

Delete

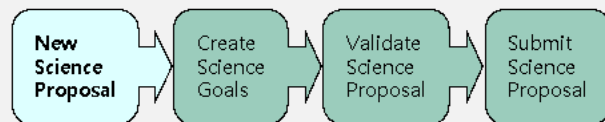
 Show image spectral windows

Overview

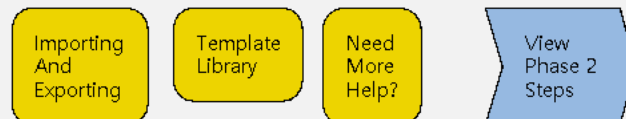
Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



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Editors

Spectral Spatial Spectral Setup

Reduce image subbands (bands 0 and 1 only)

Polarization products desired XX DUAL FULL

Spectral Setup Errors

Can not configure LOs for these spectral windows

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1/2	345.79599 GHz	345.78176 GHz	CO v=0 3-2	117.188 MHz, 141.113 kHz	2	<input checked="" type="radio"/>
1/4	344.31061 GHz	344.29645 GHz	SO 3Σ v=0 8(8)-7(7)	58.594 MHz, 141.113 kHz	2	<input type="radio"/>

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-2

1/2	329.33055 GHz	329.31700 GHz	C18O 3-2	Please select a correlator mode	1	<input type="radio"/>
-----	---------------	---------------	----------	---------------------------------	---	-----------------------

Error : Outside of IF bandwidth

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

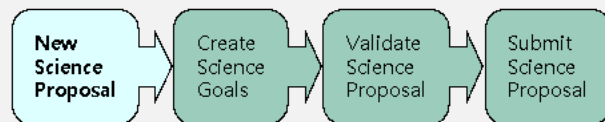
Baseband-3

Overview

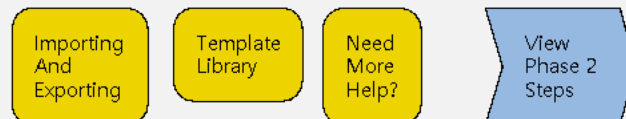
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





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Spectral Spatial Spectral Setup

Reduce image subbands (bands 0 and 1 only)

Polarization products desired

 XX DUAL FULL

Spectral Setup Errors

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1/2	345.79599 GHz	345.78176 GHz	CO v=0 3-2	117.188 MHz(102 km/s), 141.113 kHz(0.122 km/s)	2	<input checked="" type="radio"/>
1/4	344.31061 GHz	344.29645 GHz	SO 3Σ v=0 8(8)-7(7)	58.594 MHz(51 km/s), 141.113 kHz(0.123 km/s)	2	<input type="radio"/>

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-2

1(Full)	330.58797 GHz	330.57436 GHz	13CO v=0 3-2	117.188 MHz(106 km/s), 70.557 kHz(0.064 km/s)	2	<input type="radio"/>
---------	---------------	---------------	--------------	---	---	-----------------------

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-3

Overview

Contextual Help

Phase I: Science Proposal

1. Please ensure you and your co-Is are registered with the ALMA Science Portal.
2. Create a new proposal
 - Selecting a template
 - Clicking on the 'Create Proposal' button
 - Or clicking on the 'New Proposal' button
3. Click on the [proposal](#) tree node and complete the relevant fields.

Representative window (frequency) is used to estimate beam size and sensitivity

Importing And Exporting

Template Library

Need More Help?

View Phase 2 Steps

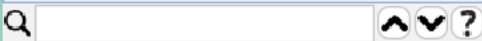


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Spectral Spatial Spectral Setup

Representative Frequency

The representative frequency is used in conjunction with the sensitivity entered on the 'Control and Performance' page to estimate the required observing time and to set the size of the antenna beam shown in the 'Spatial Visual' editor. If the transition you are most interested in does not fall in the centre of the chosen spectral window, its frequency can be changed here. The sky equivalents of the representative frequency are shown in the targets table below.

 GHz

Rest Frequencies

Please set the rest frequencies of spectral lines that will be observed. These will be used during data reduction to set the velocity scale and will enhance the ALMA Science Archive. We recommend that this be done once the spectral setup is fully defined.

Sources

Source Name	Velocity	Frame	Representative Frequency (Observed)
TW Hya	12.335 km... hel		345.7818 GHz

Contextual Help

Phase I: Science Proposal

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#).
2. Create a new proposal
 - Selecting 'New Proposal' from the 'Tools' menu.
 - Clicking on the 'New Proposal' button in the 'Tools' menu.
 - Or clicking on the 'New Proposal' button in the 'Tools' menu.
3. Click on the [proposal](#) tree node and complete the relevant fields.

Representative window (frequency) is used to estimate beam size and sensitivity



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Spectral Spatial Spectral Setup

Spectral Line

Baseband-1

Fraction	Centre Freq (rest, hel)	Centre Freq (sky, hel)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1(Full)	230.53800 GHz	230.52851 GHz	CO v=0 2-1	117.188 MHz(152 km/s), 282.227 kHz(0.367 km/s) (4-bit)	2	<input type="radio"/>

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-2

1(Full)	231.32183 GHz	231.31231 GHz	N2D+ J=3-2	58.594 MHz(76 km/s), 141.113 kHz(0.183 km/s) (4-bit)	2	<input type="radio"/>
---------	---------------	---------------	------------	--	---	-----------------------

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-3

1(Full)	219.56036 GHz	219.55132 GHz	C180 2-1	58.594 MHz(80 km/s), 35.278 kHz(0.048 km/s) (2-bit)	2	<input checked="" type="radio"/>
---------	---------------	---------------	----------	---	---	----------------------------------

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Baseband-4

1(Full)	218.00000 GHz	217.99103 GHz	cont	1875.000 MHz(2579 km/s), 1.129 MHz(1.553 km/s) (2-bit)	2	<input type="radio"/>
---------	---------------	---------------	------	--	---	-----------------------

Add spectral window centred on a spectral line

Add spectral window manually

Delete

 Show image spectral windows

Feedback

Validation Validation History Log

Q

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal



TIP:

When all targeted lines have narrow bandwidths, put at least one baseband for the continuum (broad bandwidth) for (self)calibration.

Total bandwidth > 1~2 GHz



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 - ▼ ScienceGoal (Copy of Test)
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 - Control and Performance



Overview

Editors

Spectral Spatial Calibration Setup

Select calibration strategy.

Goal Calibrators

By default, calibrators will be selected automatically at runtime and a single observation will be used to calibrate the bandpass and flux scale.

- System-defined calibration (recommended)
- System-defined calibration (force separate amplitude calibration using solar-system object)
- User-defined calibration



Astrometry

If you wish positional accuracy that is better than that provided by default (see the Proposer's Guide for more information) then select enhanced accuracy.

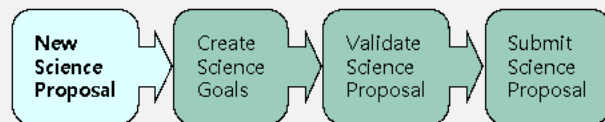
- Standard positional accuracy (default)
- Enhanced positional accuracy

DGC Override (observatory-use only)

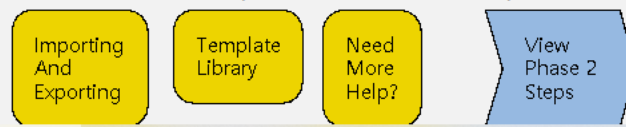
Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





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Editors

Spectral Spatial **Control and Performance**

These parameters are used to control various aspects of the observations, including the required antenna configurations and integration times.

Configuration Information

Antenna Beamsize ($1.13 * \lambda / D$) 12m 7m

Number of Antennas 12m 7m TP

ACA 7m configuration Most compact 12m configuration Most extended 12m configuration

Longest baseline

Synthesized beamsize

Shortest baseline

Maximum recoverable scale

Desired Performance

Desired Angular Resolution (Synthesized Beam) Single Range Any Standalone ACA

arcsec to arcsec

Largest Angular Structure in source arcsec

Desired sensitivity per pointing mJy equivalent to @ 0.150 "

Feedback

Validation Validation History Log

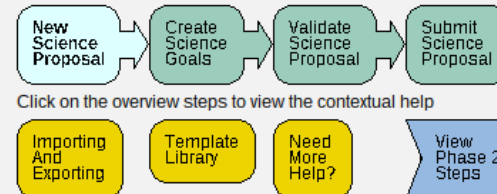
Description	Suggestion

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal





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Editors

Spectral Spatial **Control and Performance**

Shortest baseline	<input type="text" value="0.009 km"/>	<input type="text" value="0.015 km"/>	<input type="text" value="0.256 km"/>
Maximum recoverable scale	<input type="text" value="19.388 arcsec"/>	<input type="text" value="8.391 arcsec"/>	<input type="text" value="0.145 arcsec"/>

Desired Performance

Desired Angular Resolution (Synthesized Beam) Single Range Any Standalone ACA

arcsec to arcsec

Largest Angular Structure in source arcsec

Desired sensitivity per pointing mJy equivalent to @ 0.150 " and @ 0.0500 "

Bandwidth used for Sensitivity Frequency Width

Override OT's sensitivity-based time estimate (must be justified) Yes No

Science Goal Breakdown: time estimate, clustering, beam and configurations

Simultaneous 12-m and ACA observations Yes No

Are the observations time-constrained? Yes No

Range is recommended
(single ~ < 20%)

Flux unit is used although
K can be chosen.

$$\left(\frac{T}{1 \text{ K}}\right) = \left(\frac{S_\nu}{1 \text{ Jy}}\right) \left[13.6 \left(\frac{300 \text{ GHz}}{\nu}\right)^2 \left(\frac{1''}{\theta_{max}}\right) \left(\frac{1''}{\theta_{min}}\right)\right]$$

Feedback

Validation Validation History Log

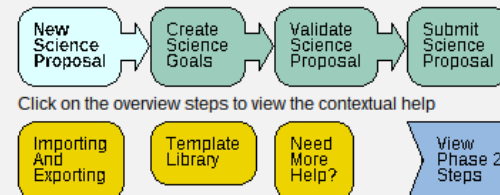
Description	Suggestion

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal





Project Structure

Proposal Program

Unsubmitted Proposal

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 - Technical Justification

Editors

Spectral Spatial Control and Performance

Shortest baseline	0.009 km	0.015 km	0.256 km
Maximum recoverable scale	19.388 arcsec	8.391 arcsec	0.145 arcsec

Desired Performance

Desired Angular Resolution (Synthesized Beam) Single Range Any Standalone ACA

0.05 arcsec to 0.15 arcsec

Largest Angular Structure in source: 1.00000 arcsec

Desired sensitivity per pointing: 3.3 mJy equivalent to 1.4998 K @ 0.150 "

and 13.499 K @ 0.0500 "

Bandwidth used for Sensitivity: RepWindowEffectiveChannelWidth Frequency Width: 0.195313 MHz

Override OT's sensitivity-based time estimate (must be justified)

Science Goal Breakdown: time estimate, clustering, beam and configurations

Simultaneous 12-m and ACA observations

Are the observations time-constrained? Yes No

- RepWindowEffectiveChannelWidth
- RepresentativeWindowBandWidth
- RepWindowEffectiveChannelWidth
- AggregateBandWidth
- LargestWindowBandWidth
- FinestEffectiveChannelWidth
- User

Line
Continuum

Feedback

Validation Validation History Log

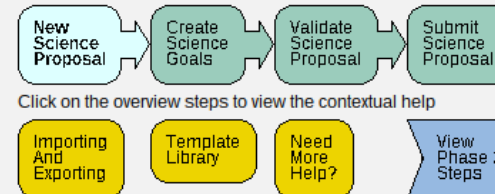
Description	Suggestion

Overview

Contextual Help

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Phase I: Science Proposal





Project Structure

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Editors

Spectral Spatial Control and Performance

Maximum recoverable scale

Desired Performance

Desired Angular Resolution (Synthesized Beam) Single Range Any Standalone ACA

arcsec to

Largest Angular Structure in source arcsec

Desired sensitivity per pointing mJy equivalent to and

Bandwidth used for Sensitivity Frequency Width

Override OT's sensitivity-based time estimate (must be justified) Yes No

Science Goal Breakdown: time estimate, clustering, beam and configurations

Simultaneous 12-m and ACA observations Yes No

Are the observations time-constrained? Yes No

Feedback

Validation Validation History Log

Description

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
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 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

New Science Proposal

Click on the overview s

Importing And Exporting

Phas

Cre Sci Go

Ten Libr

Planning and Time Estimate

Note: The time in brackets is that required to reach the sensitivity. Operational requirements often mean that the actual observed time is longer, especially for mosaics. Please see the User Manual for more details.

Input Parameters

Requested sensitivity 3.300 mJy
 Bandwidth used for sensitivity 0.195 MHz
 Representative frequency (sky, first source) 345.782 GHz

Estimated Total time for Science Goal

8.93 h

Cluster 1

Source Name	RA	Dec	Velocity
TW Hya	11:01:51.9053	-34:42:17.033	12.335 km/s

Possible Configuration Combinations

12-m (1)	12-m (2)	7-m	TP
C-6	None	No	No

Input Parameters

Precipitable water vapour (all sources) 1.262mm (4th Octile)

Time required for 12m (1) [C-6]

Time on source per pointing (first source) 2.82 h [2.77 h]
 Total number of pointings (all sources) 1
 Number of tunings 1
 Total time on source 2.82 h [2.77 h]
 Total calibration time 5.80 h
 Other overheads 18.40 min
 Total time for 1 SB execution 1.49 h
 Number of SB executions 6
 Total time to complete SB 8.93 h

Calibration Breakdown per SB execution

4 x Pointing 8.00 min
 1 x Amplitude/bandpass 10.00 min
 19 x Phase 5.70 min
 4 x CheckSource 2.00 min
 2 x DGC 8.00 min
 10 x Atmospheric 6.67 min
 1 x DGC Bandpass 5.00 min

Close



Project Structure

Proposal Program

Unsubmitted Proposal

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 - ▼ ScienceGoal (Copy of Test)
 - General
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 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Technical Justification

Enter a Technical Justification for this Science Goal, paying special attention to the parameters reproduced below.

Sensitivity

Requested RMS over is

For a peak flux density of , the S/N is

Achieved RMS over the total bandwidth is

For a continuum flux density of , the achieved S/N is

For a peak line flux of , the achieved S/N over 1/3 of the source line width (/ 3 =) is

Line width / bandwidth used for sensitivity (/) =

Spectral Dynamic Range (continuum flux / line rms):

Justify your requested RMS and resulting S/N for the spectral line and/or continuum observations.



For line observations also justify the bandwidth used for the sensitivity calculation.

aa

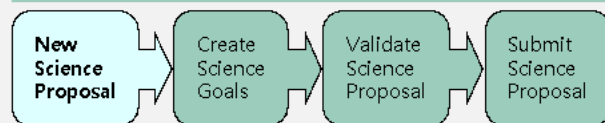
Imaging

Overview

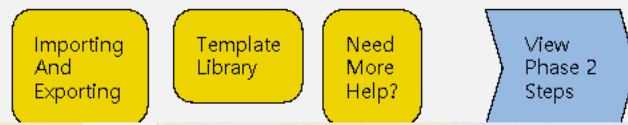
Contextual Help

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2. Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the  icon in the toolbar
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3. Click on the  [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help



Sensitivity

Requested RMS over 88.563 m/s is 7.07 mJy

A For a peak flux density of 45.00 mJy , the S/N is 6.4

Achieved RMS over the total 4.102 GHz bandwidth is 34.29 uJy

For a continuum flux density of 0.00 mJy , the achieved S/N is 0.0

For a peak line flux of 45.00 mJy **B** the achieved S/N over 1/3 of the source line width (300.00 m/s / 3 = 100.00 m/s) is 6.8

Line width / bandwidth used for sensitivity (300.00 m/s / 88.56 m/s) = 3.39

C Integrated Intensity

Justify your requested RMS and resulting S/N for the spectral line and/or continuum observations.

For line observations also justify the bandwidth used for the sensitivity calculation.

1. How the expected intensity is calculated
2. Which SNR is used among A, B, and C.
3. Why is the SNR enough to achieve the Science Goal?

- the fine spectral resolution can be chosen because of the kinematics, but analysis will be done with poor spectral resolution (B)

- Although it is a line observation, describe the information and arguments for the continuum

Tip: You can refer the figures in S.J. and/or cite references (XX et al. ApJ, ###, ###)

Requested angular resolution Requested Largest Angular Scale

Justify the chosen angular resolution and largest angular scale for the source(s) in this Science Goal

1. Why the requested beam size is needed?
 1. To resolve the emitting region
 2. You need several points for analysis
2. LAS
 1. Guarantees No flux missing, i.e., if the total flux is important in your science, you should set LAS as the largest emitting area
 2. If not, LAS can be smaller than the latter, but you describe that you are interested in just morphonology,

The expected emitting area is $\sim 0.3''$, thus the requested angular resolution **can resolve** the emitting area by a factor of two, which can ...

The disk is extended up to $0.3''$, and we set LAS of $0.3''$.

Correlator configuration



line width / representative spectral window resolution: / =

Representative spectral window width :

Justify your correlator set-up with particular reference to the number of spectral resolution elements per line width.
You may want to consider spectral averaging to lower the data rate

The spectral resolution is based on

- flux measurement (3 points within Line width)
 - Kinematics (more fine resolution is needed)
- ex) resolution of 0.065 km/s is enough to resolve the line width of 0.3.

Goals of each targeted lines and why the transitions are selected

CO and ^{13}CO lines are chosen for ...

Furthermore, the lines in the Band 7 provide the highest SNR for a given observing time.

SO and HCO^+ are set for bonus lines, which can provide ...

Tips) At least One baseband is set to continuum for the (self) calibration

File Edit View Tool Search Help



Project Structure

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Editors

Spectral Spatial Technical Justification

Enter a Technical Justification for this Science Goal, paying special attention to the parameters reproduced below.

Sensitivity

Requested RMS over is For a peak flux density of , the S/N is

Achieved RMS over the total bandwidth is For a continuum flux density of , the achieved S/N is

For a peak line flux of , the achieved S/N over 1/3 of the source line width (/ 3 =) is

Line width / bandwidth used for sensitivity (/) =

Note that the bandwidth used for sensitivity is larger than 1/3 of the linewidth.
The S/N achieved for a resolution element that allows the line to be resolved will be lower than that reported.

Spectral Dynamic Range (continuum flux / line rms):

Justify your requested RMS and resulting S/N for the spectral line and/or continuum observations.

For line observations also justify the bandwidth used for the sensitivity calculation.

Imaging

Requested angular resolution Requested Largest Angular Scale

Justify the chosen angular resolution and largest angular scale for the source(s) in this Science Goal

Correlator configuration

line width / representative spectral window resolution: / =

Feedback

Validation Validation History Log

Q ^ v ?

Overview

If the blue text appears, you **should** describe why that parameter is requested, Or change the parameter.

ALMA Observing Tool Reference Manual

Contents Search Favourites

Next Up Previous Contents

Next: [Phase 2 Program Scheduling](#) Up: [Technical Justification](#) Previous: [Correlator configuration](#) [Contents](#)

Choices to be justified

If selected by the user, the following user settings will also need to be justified. A free-format box appears for each.

- Non-Nyquist mosaic sampling: Non-Nyquist values are usually used when the scientific goal is to cover a large survey area and large-scale structures are not being observed. Polarization mosaicing may benefit from sampling at higher than the Nyquist frequency.
- Single polarization: Dual polarization is more usually selected and single polarization only used when the highest spectral resolution is required.
- Low max elevation: Sources with low declinations will suffer large atmospheric attenuation and be difficult to schedule because of limited time above the horizon.
- User-defined calibration: As the observatory guarantees appropriate calibration without user input, this must be rigorously justified. For VLBI and Phased Array projects, the choice of phasor should be explained here.
- Override of OT's sensitivity-based time estimate: This may be necessary if you want to monitor a source over a certain time span or if your source has a complicated structure but the sensitivity-based time estimate does not allow for sufficient uv -coverage. You must give a detailed justification for the time override, and explain how the new time estimate was calculated. The time override must include calibrations and overheads.
- Time-constrained observing: These imply significant constraints on the scheduling of all ALMA projects and must therefore be fully justified.
- Astrometry: If enhanced positional accuracy was selected, please explain why this is necessary i.e. state the positional requirements with reference to what ALMA provides by default (consult the Proposer's Guide and Technical Handbook for more information).
- Solar Technical Justification: All technical details of solar observations should be discussed in this text box. See the Proposer's Guide for more information.
- VLBI Technical Justification: All technical details of VLBI observations should be discussed in this text box. If passive phasing has been selected then the choice of calibrator must be discussed in the box reserved for that purpose. See the Proposer's Guide for more information.
- Phased Array Technical Justification: All technical details of phased-array observations should be discussed in this text box and there should also be discussion of the post-processing steps. See the Proposer's Guide for more information.
- Passive Phasing Technical Justification: Describe here the source selected to be used for passive phasing i.e. why it is suitable to be used for this purpose. Its brightness and proximity should be discussed.
- High Imaging and/or Spectral Dynamic Range Technical Justification: Describe here why high dynamic range is required and how this can be achieved.

Editors

Spectral Spatial **Technical Justification**

Correlator configuration

line width / representative spectral window resolution: / =

Representative spectral window width :

Justify your correlator set-up with particular reference to the number of spectral resolution elements per line width. You may want to consider spectral averaging to lower the data rate

Choices to be justified

Justify and provide additional details on time constraints.

Feedback

Validation Validation History Log

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reviewers are requested to update their user profiles with combinations of scientific categories and keywords which describe their area(s) of expertise using the new 'Expertise' tab in <https://asa.alma.cl/UserRegistration/secure/updateAccount.jsp>. Available expertise information will be used in the distribution of proposal assignments.

Reviewer has a PhD? No Yes

Select Mentor

Mentor name

Mentor has a PhD? No Yes

Science Case

Please ensure that your science case is properly anonymized following instructions on the Science Portal

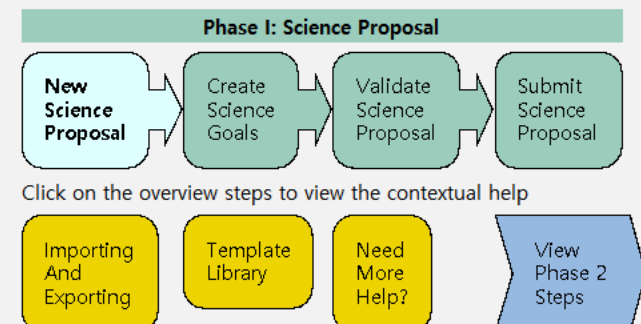
... justify any new observations that duplicate archival data or accepted programs. Information regarding the ALMA Duplication Policy and how to search archival data and accepted programs can be found at: ...

1. Save
2. Validate check
3. No errors can submit the proposal
4. Warnings can be ignored.

10 errors, 3 warnings : double-click on each row to be taken to the problem

Description	Suggestion
No document found - you must add a Science Case to your proposal	Select the proposal node in the Proposal tab and add your document
No mentor has been defined	Please select a mentor (must be a registered ALMA user)
Neither the reviewer or mentor have a PhD	Please select a reviewer or mentor with a PhD

- ### Contextual Help
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Project Structure

Submit Project to ALMA

Editors

Spectral Spatial Control and Performance

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Synthesized beamsize	5.721 arcsec	1.477 arcsec	0.025 arcsec
Shortest baseline	0.009 km	0.015 km	0.256 km
Maximum recoverable scale	30.534 arcsec	13.215 arcsec	0.228 arcsec

Desired Performance

Desired Angular Resolution (Synthesized Beam) Single Range Any Standalone ACA

0.00000 arcsec

Largest Angular Structure in source Undefined arcsec

Desired sensitivity per pointing 0.00000 Jy equivalent to Infinity K

RepWindowEffectiveChannelWidth Frequency Width 0.048828 MHz

 Yes No

Planning and Time Estimate

 Yes No Yes No

Time Windows

 None Single Visit Multiple Visits

Number of time windows specified : 1

Start Date/Time (UTC)	End Date/Time (UTC)
2025-03-13T07:25:26Z	2025-03-27T07:25:26Z

Please specify one or more suitable time windows for your observation

Your observation will be scheduled once during **ONE** of these intervals.

Add Edit Delete Delete All Import

Feedback

Validation Validation History Log

Q

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal



1. Submit

2. You can re-submit it until the deadline.

The server is very busy around the deadline, thus, submit it a few days ago, and resubmit it.

ALMA LO Configuration Tool...
Sensitivity Calculator...

Project Structure

Proposal Program

Unsubmitted Proposals

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Generate SBs from the Selected Goal Ctrl-B
Display Project Time Summary
Generate Phase 2 SBs from all the Science Goals Ctrl-B
Generate a PDF of Whole Proposal
Disable Edit Proposal
Generate a PDF of Whole Proposal



Technical Justification

Requested angular resolution 50.00 mas - 10.00 mas

Requested Largest Angular Scale 1.00 arcsec

Justify the chosen angular resolution and largest angular scale for the source(s) in this Science Goal

aaaaaaaaaaaaaaaaaaaaaaaaaaaa

Make a PDF file for circulating the proposal to co-I



Feedback

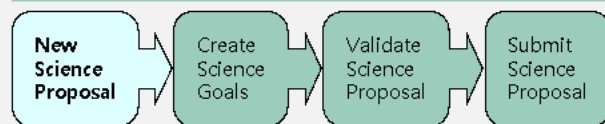
Validation Validation History Log

7 errors, 3 warnings : double-click on each row to be taken to the problem

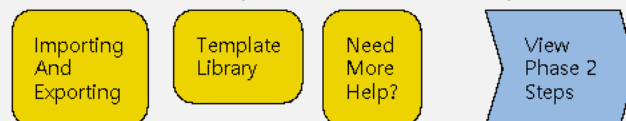
Description	Suggestion
⚠ Neither the reviewer or mentor have a PhD	Please select a reviewer or mentor with a PhD
⚠ Spectral setup has low aggregate bandwidth	Calibration will use bandwidth switching. Increase the bandwidth if possible
⚠ Desired angular resolution is outside the range allowed by the available arrays and	Select the Control Parameters in the Science Goal and enter a valid value
⚠ Part of spw CO v=0 3-2 in BB 2 lies within 30 MHz of the baseband edge. This could result in	Move the spw away from the baseband edge to avoid this problem.

Overview**Contextual Help**

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Phase I: Science Proposal

Click on the overview steps to view the contextual help





**Multiple targets
&
Mosaics**



Project Structure

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Editors

Spectral Spatial Field Setup

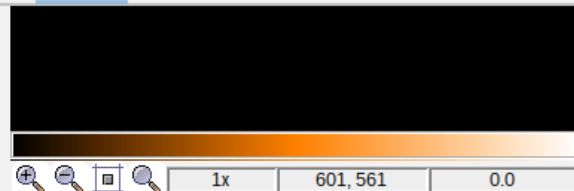


Image Filename

1x 601, 561 0.0

FOV Parameters

Representative Frequency (Sky)

Array Type 12m

Antenna Beamsize (HPBW)

Show Antenna Beamsize

Image Query

Image Server

Image Size(arcmin)

Peak Line Flux Density per Synthesized Beam

Line Width

Line Linear Polarization

Line Circular Polarization

Field Centre Coordinates

Coord Type Relative AbsoluteArray Type 12mOffset Unit #Pointings 12m Array

RA [arcsec]	Dec [arcsec]
0.00000	0.00000

Multiple sources within S.G.

Feedback

Validation Validation History Log

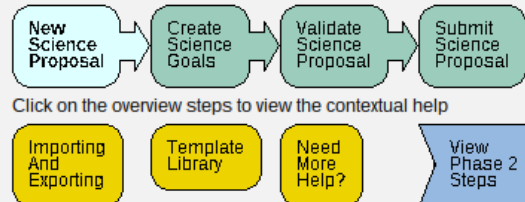
Description	Suggestion

Overview

Contextual Help

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Phase I: Science Proposal





Project Structure

Proposal Program

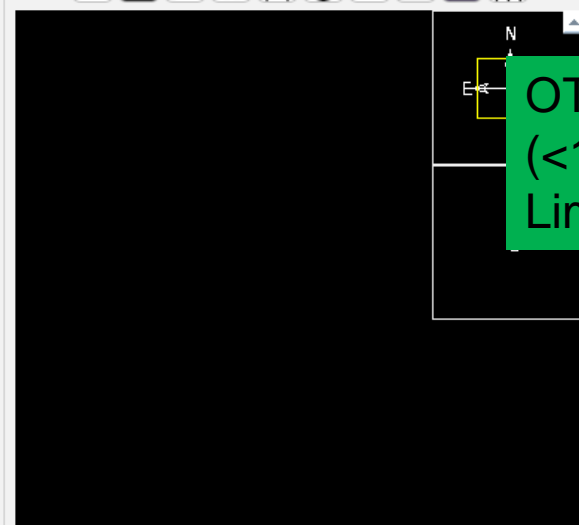
Unsubmitted Proposal

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Spectral Spatial Field Setup

Spatial Image



TW_Hya TW_Hya_1 TW_Hya_2 TW_Hya_5 TW_Hya_4 TW_Hya_5

Source

Source Name TW_Hya_3

Choose a Solar System Object?

Name of object Unspecified

Resolve

OT automatically split sources into different "clusters (<5)"
 (<10 degrees or < 1 degree for C7-C10)
 Limit of sources : ≤ 150

Expected Source Properties

Peak Continuum Flux Density per Synthesized Beam	5.00000	Jy
Continuum Linear Polarization	0.0	per cent
Continuum Circular Polarization	0.0	per cent
Peak Line Flux Density per Synthesized Beam	2.00000	Jy
Line Width	3.00000	km/s
Line Linear Polarization	0.0	per cent

Feedback

Validation Validation History Log

Description

Suggestion

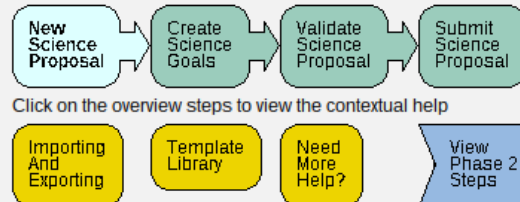


Overview

Contextual Help

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Phase I: Science Proposal



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Mosaic or several positions within a target

- set Representative Frequency before Field setup
- Upload or Query the image

3x 359.5, 251.2 4124.0

11:01:46.941, -34:43:00.91 (J2000)

Image Filename unitWjsky3WcacheWjsky12077043586718440107.fits

FOV Parameters

Representative Frequency (Sky) 345.782 GHz

Array Type 12m

Antenna Beamsize (HPBW) 16.840 arcsec

Show Antenna Beamsize

Image Query

Image Server Digitized Sky (Version II) at ESO

Image Size(arcmin) 10.0 Query

Peak Line Flux Density per Synthesized Beam 10.00000 mJy

Line Width 1.00000 km/s

Line Linear Polarization 0.0 per cent

Line Circular Polarization 0.0 per cent

Rectangle

Coords Type Relative Absolute

Field Centre Coordinates Offset(Longitude) 0 arcsec

Offset(Latitude) 0 arcsec

p length 50.0 arcsec

q length 40.0 arcsec



Position Angle 0.0 deg

Spacing 0.51093 fraction of antenna beamsize Reset to Nyquist

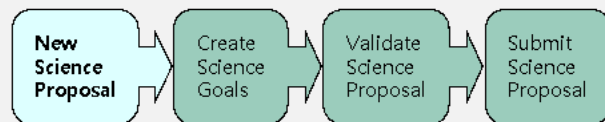
#Pointings 12m Array 33 Export

Add Source Load from File Export to File Clone Source Delete Source Delete All Sources

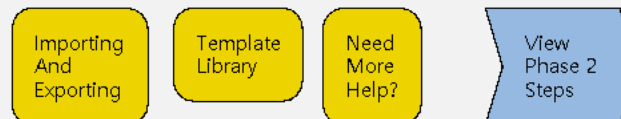
Contextual Help

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Phase I: Science Proposal



Click on the overview steps to view the contextual help





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Spectral Spatial Field Setup

Spatial Image

Add a FOV

11:01:51.412, -34:40:48.86 (J2000)

Image Filename shlee/.jsky3/cache/jsky1015745759644008315.fits

FOV Parameters

Representative Frequency (Sky) 350.500 GHz

Array Type 12m

Antenna Beamsize (HPBW) 16.613 arcsec

Show Antenna Beamsize

Image Query

Image Server Digitized Sky (Version II) at ESO

TW Hya

Source

Source Name TW Hya Resolve

Choose a Solar System Object? Name of object Unspecified

System ICRS Sexagesimal display? Parallax 16.62890 mas

Source Coordinates RA 11:01:53.6609 PM RA -68.30900 mas/yr

Dec -34:42:04.778 PM Dec -13.90000 mas/yr

Source Radial Velocity 12.335 km/s Isrk z 0.000041146 Doppler Type RELATIVISTIC

Target Type Individual Pointing(s) 1 Rectangular Field

Expected Source Properties

Peak Continuum Flux Density per Synthesized Beam 0.00000 Jy

Continuum Linear Polarization 0.0 per cent

Continuum Circular Polarization 0.0 per cent

Peak Line Flux Density per Synthesized Beam 0.00000 Jy

Line Width 0.00000 km/s

Line Linear Polarization 0.0 per cent

Line Circular Polarization 0.0 per cent

Field Centre Coordinates

Coord Type Relative Absolute

Array Type 12m

Offset Unit arcsec

#Pointings 12m Array 4

RA [arcsec]	Dec [arcsec]
-21.93508	-12.73871
-36.96023	12.58119
30.00000	0.00000
-15.95692	49.35603

Add Delete Reset Import Export

1. Click Add a FOV
2. Double click at the targeted position within the panel

Or

1. Click 'Add' button
2. Put RA and Dec.

Feedback

Validation Validation History Log

Q

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal

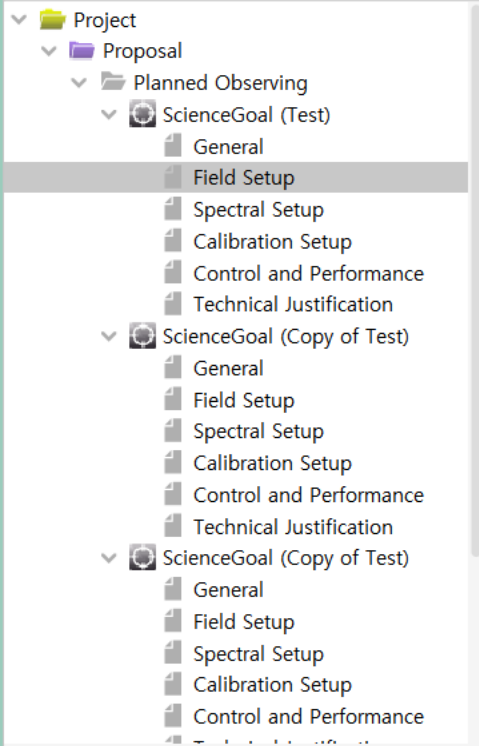




Project Structure

Proposal Program

Unsubmitted Proposal



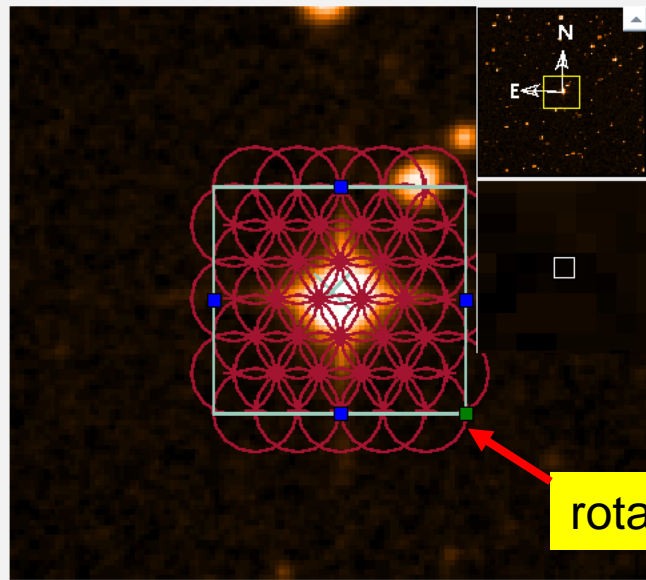
Editors

Spectral Spatial Field Setup

Draw box

Show mosaic

Spatial Image



rotation

3x 359.8, 265.8 3842.0
11:01:46.951, -34:42:46.14 (J2000)
Image Filename unitWiskv3WcracheWiskv12077043586718440107.fits

TW Hya

Source

Source Name TW Hya Resolve
Choose a Solar System Object? Name of object Unspecified
System ICRS Sexagesimal display?
Source Coordinates RA 11:01:52.0913 Parallax 16.62890 mas
Dec -34:42:15.750 PM RA -68.30900 mas/yr
Source Radial Velocity 12.335 km/s hel z 0.000041146 Doppler Type RELATIVISTIC
Target Type Individual Pointing(s) 1 Rectangular Field

Expected Source Properties

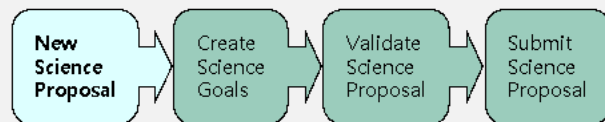
Peak Continuum Flux Density per Synthesized Beam 50.00000 mJy
Continuum Linear Polarization 0.0 per cent
Continuum Circular Polarization 0.0 per cent
Peak Line Flux Density per Synthesized Beam 10.00000 mJy
Line Width 1.00000 km/s
Line Linear Polarization 0.0 per cent
Line Circular Polarization 0.0 per cent

Overview

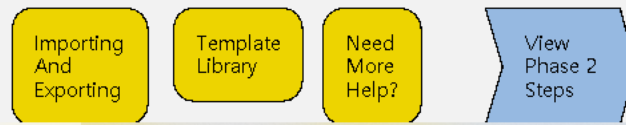
Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
3. Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Test)
 - General
 - Field Setup**
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification
 - ScienceGoal (Copy of Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification
 - ScienceGoal (Copy of Test)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Field Setup



3x 359.5, 251.2 4124.0
 11:01:46.941, -34:43:00.91 (J2000)
 Image Filename unitWjsky3WcacheWjsky12077043586718440107.fits

FOV Parameters

Representative Frequency (Sky) 345.782 GHz
 Array Type 12m
 Antenna Beamsize (HPBW) 16.840 arcsec
 Show Antenna Beamsize

Image Query

Image Server Digitized Sky (Version II) at ESO
 Image Size(arcmin) 10.0 Query

Continuum Linear Polarization 0.0 per cent
 Continuum Circular Polarization 0.0 per cent
 Peak Line Flux Density per Synthesized Beam 10.00000 mJy
 Line Width 1.00000 km/s
 Line Linear Polarization 0.0 per cent
 Line Circular Polarization 0.0 per cent

Rectangle

Coords Type Relative Absolute

Field Centre Offset(Longitude) 0 arcsec

Coordinates Offset(Latitude) 0 arcsec

p length 50.0 arcsec

Length of the 'p' side of the rectangle

q length 40.0 arcsec

Position Angle 0.0 deg

Spacing 0.51093 fraction of antenna beamsize Reset to Nyquist

#Pointings 12m Array 33 Export

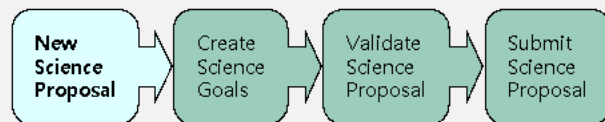
Add Source Load from File Export to File Clone Source Delete Source Delete All Sources

Overview

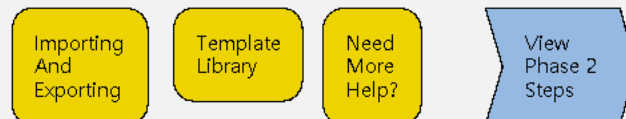
Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
2. Create a new proposal by either:
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 - Clicking on the icon in the toolbar
 - Or clicking on this [link](#)
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Phase I: Science Proposal



Click on the overview steps to view the contextual help





Time Constrain Observation



Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Control and Performance

Synthesized beamsize	5.721 arcsec	1.477 arcsec	0.025 arcsec
Shortest baseline	0.009 km	0.015 km	0.256 km
Maximum recoverable scale	30.534 arcsec	13.215 arcsec	0.228 arcsec

Desired Performance

Desired Angular Resolution (Synthesized Beam) Single Range Any Standalone ACA

0.00000 arcsec

Largest Angular Structure in source

Undefined arcsec

Desired sensitivity per pointing

0.00000 Jy equivalent to Infinity K

Bandwidth used for Sensitivity

RepWindowEffectiveChannelWidth Frequency Width 0.048828 MHz

Override OT's sensitivity-based time estimate (must be justified)

Yes No

Science Goal Breakdown: time estimate, clustering, beam and configurations

Planning and Time Estimate

Simultaneous 12-m and ACA observations

Yes No

Are the observations time-constrained?

Yes No

Time Windows

None Single Visit Multiple Visits

Number of time windows specified : 1

Start Date/Time (UTC)	End Date/Time (UTC)
2025-03-13T07:25:26Z	2025-03-27T07:25:26Z

Please specify one or more suitable time windows for your observation

Your observation will be scheduled once during ONE of these intervals.

Add

Edit

Delete

Delete All

Import

Feedback

Validation Validation History Log

Q

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal





Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Control and Performance

Synthesized beamsize	5.721 arcsec	1.477 arcsec	0.025 arcsec
Shortest baseline	0.009 km	0.015 km	0.256 km
Maximum recoverable scale	30.534 arcsec	13.215 arcsec	0.228 arcsec

Desired Performance

Desired Angular Resolution (Synthesized Beam) Single Range Any Standalone ACA

Largest Angular Structure in source

Desired sensitivity per pointing equivalent to

Bandwidth used for Sensitivity Frequency Width

Override OT's sensitivity-based time estimate (must be justified) Yes No

Science Goal Breakdown: time estimate, clustering, beam and configurations

Simultaneous 12-m and ACA observations Yes No

Are the observations time-constrained? Yes No

Time Windows

None Single Visit Multiple Visits

Visits specified : 2

Visit Constraints (UTC)
Visit 1 : Arbitrary start
Visit 2 : To be scheduled 14.0 d after visit 1 with a margin of ± 7.0 d

Please specify the arrangement of visits for your observation.

Visits can either be for a specific date or relative to a previous visit.

The first visit can be defined as having an arbitrary start date/time

Feedback

Validation Validation History Log



Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal





Joint Proposal



Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Project
 - ▼ Proposal
 - ▼ Planned Observing
 - ▼ ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Joint Proposals

Is this a Joint Proposal? Yes NoType of Joint Proposal Main Partner

Observatory	Project Code	Requested Time

Add Partner Observatory

Remove Partner Observatory

Investigators

Type	Full name	Email	Affiliation	ALMA ID	Executive	Reviewer

Feedback

Validation Validation History Log

Description

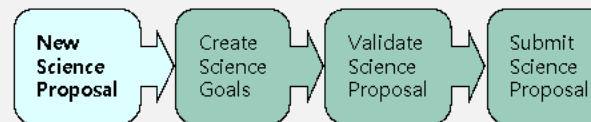
Suggestion

Overview

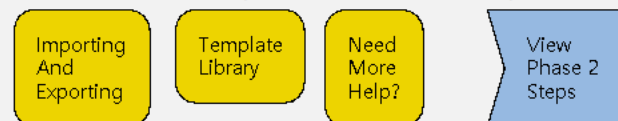
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the **1** icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help



Comparison of Observing time
 Main : ALMA > Partner
 Partner : ALMA < Partner



Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Project
 - ▼ Proposal
 - ▼ Planned Observing
 - ▼ ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Joint Proposals

Is this a Joint Proposal? Yes NoType of Joint Proposal Main Partner

Observatory	Project Code	Requested Time
JWST	N/A	0.00 h
JWST		
VLA		
VLT		

Add Partner Observatory

Remove Partner Observatory

Please provide the technical justification for the time requested on JWST as a joint proposal

Feedback

Validation Validation History Log

Description

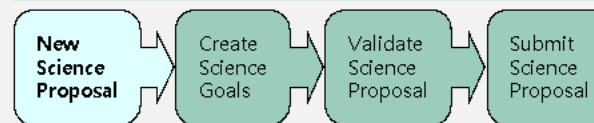
Suggestion

Overview

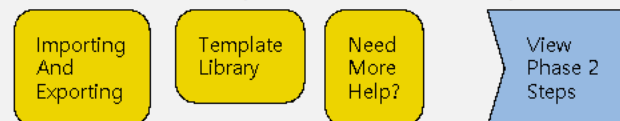
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the **1** icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Project
 - ▼ Proposal
 - ▼ Planned Observing
 - ▼ ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Joint Proposals

Is this a Joint Proposal? Yes NoType of Joint Proposal Main Partner

Observatory	Project Code	Requested Time
JWST	N/A	1.0 h

Add Partner Observatory

Remove Partner Observatory

Please provide the technical justification for the time requested on JWST as a joint proposal



Feedback

Validation Validation History Log

Description

Suggestion

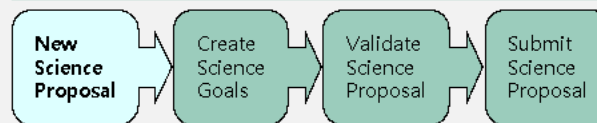


Overview

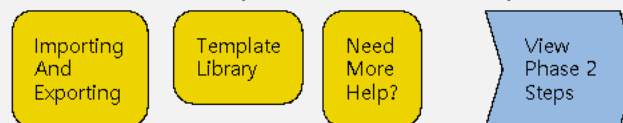
Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:
 - Selecting *File > New Proposal*
 - Clicking on the **1** icon in the toolbar
 - Or clicking on this [link](#)
- Click on the [proposal](#) tree node and complete the relevant fields.

Phase I: Science Proposal



Click on the overview steps to view the contextual help





Project Structure

Proposal Program

Unsubmitted Proposal

- ▼ Project
 - ▼ Proposal
 - ▼ Planned Observing
 - ▼ ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Student project

Joint Proposals

Is this a Joint Proposal? Yes NoType of Joint Proposal Main Partner

Observatory	Project Code	Requested Time
JWST	▼ N/A	1.00 h

Add Partner Observatory

Remove Partner Observatory

Please provide the technical justification for the time requested on JWST as a joint proposal

over 50 characters!

Investigators

Feedback

Validation Validation History Log

Description

Suggestion

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)

Phase I: Science Proposal

New Science

Create Science

Validate Science

Submit Science



Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Proposal

Is this a Joint Proposal? Yes NoType of Joint Proposal Main Partner

Observatory	Project Code	Requested Time
JWST	N/A	1.00 h
VLA	N/A	1.00 h



Add Partner Observatory

Remove Partner Observatory

Please provide the technical justification for the time requested on JWST as a joint proposal

Please provide the technical justification for the time requested on VLA as a joint proposal

Investigators

Type	Full name	Email	Affiliation	ALMA ID	Executive	Reviewer
...						

Feedback

Validation Validation History Log

Overview

Contextual Help

1. Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)

Phase I: Science Proposal

New Science

Create Science

Validate Science

Submit Science

A background image of a starry night sky with a faint, colorful nebula or galaxy structure visible in the lower half. The stars are of various colors, including white, blue, and orange.

Proposal type of VLBI & Phased Array

File Edit View Tool Search Help

Project Structure

Editors

Proposal Program

Spectral Spatial Field S

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup

Proposal Information

Proposal Title

Proposal Cycle

Abstract (max. 1200 characters)

Proposal Type

Regular Target Of Opportunity VLBI

Large Program Phased Array

ALMA Observing Tool Reference Manual

Contents Search Favourites

- Expected Source Properties
- Field Center Coordinates
- Rectangle
- TP Regional Mapping
- Spectral Setup
- Calibration Setup
- Control and Performance
- Technical Justification
- Phase 2 Program Scheduling Blocks
- Observing Unit Set
- The Scheduling Block
- Observing Groups
- Targets
- Field Sources
- Instrument Setup
- Observing Parameters

Next: [Field Center Coordinates](#) Up: [Field Setup](#) Previous: [Source](#) [Contents](#)

Expected Source Properties

One of the reasons for providing this information is to aid in technical assessment. All relevant items should be filled in or the project will not validate. This means that a continuum flux must be entered for a continuum project and a line flux and width must be filled in for a spectral line project. If full polarization is selected, at least one of the line or continuum polarization percentages must also be given.

- Passive phasing is required:** This is only shown for VLBI and Phased Array projects and is used to signify that the total flux of the source being observed is not sufficient to allow phasing of the array using that source and thus that an additional calibrator must be used. The continuum flux limits are 0.35 Jy for Band 1, 0.5 Jy for Band 3 and Band 6, and 0.7 Jy for Band 7. For spectral line observing the flux thresholds are calculated using the formula reported in the Proposer's Guide. If this is selected, a specific ("fixed") phase calibrator must be defined in Calibrator Setup.
- Peak Continuum Flux Density per Synthesized Beam:** The expected flux density per

Expected Source Properties

Passive phasing is required (science target < 0.7 Jy)

Field Center Coordinates

Coord Type Relative Absolute

Array Type 12m

Offset Unit

#Pointings 12m Array

RA [arcsec]	Dec [arcsec]
0.00000	0.00000

Add Delete Reset Import Export

Add Source Load from File Export to File Clone Source Delete Source Delete All Sources

Representative Frequency (Sky)

Array Type 12m

Antenna Beamsize (HPBW)

Show Antenna Beamsize

Image Query

Image Server

Feedback

Validation Validation History Log

Q

^ v ?

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal

New Science Proposal → Create Science Goals → Validate Science Proposal → Submit Science Proposal

Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Proposal Information

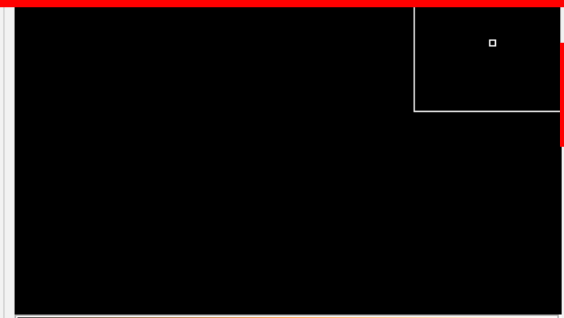
Proposal Title

Proposal Cycle

Abstract (max. 1200 characters)

Proposal Type

Regular
 Target Of Opportunity
 VLBI
 Large Program
 Phased Array



1x | 268, 447 | 0.0

Image Filename

FOV Parameters

Representative Frequency (Sky)

Array Type 12m

Antenna Beamsize (HPBW)

Show Antenna Beamsize

Image Query

Image Server

Name of object Resolve

Sexagesimal display?

Parallax

PM RA

PM Dec

z Doppler Type

Target Type Individual Pointing(s) 1 Rectangular Field

Expected Source Properties

Passive phasing is required (science target < 0.7 Jy)

Field Centre Coordinates

Coord Type Relative Absolute

Array Type 12m

Offset Unit

#Pointings 12m Array

RA [arcsec]	Dec [arcsec]
0.00000	0.00000

Add Delete Reset Import Export

Add Source Load from File Export to File Clone Source Delete Source Delete All Sources

Feedback

Validation Validation History Log

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal





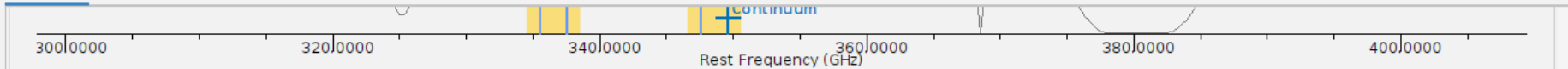
Project Structure

Proposal Program

- Unsubmitted Proposal
 - Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Science Goal)
 - General
 - Field Setup
 - Spectral Setup**
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Editors

Spectral Spatial Spectral Setup



Overlays: Receiver Bands Transmission DSB Image Spectral Lines Select Lines to Overlay

Water Vapour Column Density: Automatic Choice Manual Choice 0.913mm (3rd Octile) ▾

Viewport: Pan to Spectral Window Zoom to Band Reset

Spectral Type

Spectral Type ? ▾

Spectral Line

Single Continuum

Spectral Scan

Produce image sidebands (Bands 9 and 10 only)

Polarization products desired XX DUAL FULL

Spectral Setup Errors

Single Continuum ? ▾

Receiver Band 7 [275.0-373.0 GHz] ▾

1 [35.0-50.0 GHz]

3 [84.0-116.0 GHz]

Sky Frequency 6 [211.0-275.0 GHz]

7 [275.0-373.0 GHz]

342.000000 GHz

Low spectral resolution (TDM)

High spectral resolution (FDM)

Baseband-1

Fraction	Centre Freq (rest,topo)	Centre Freq (sky,topo)	Transition	Bandwidth, Resolution (smoothed)	Spec. Avg.	Representative Window
1(Full)	335.60000 GHz	335.60000 GHz	Single Continuum	1875.000 MHz(1675 km/s), 7.813 MHz(6.979 km/s) (2-bit)+VLBI	8	<input type="radio"/>

Show image spectral windows

Baseband-2

1(Full)	337.60000 GHz	337.60000 GHz	Single Continuum	1875.000 MHz(1665 km/s), 7.813 MHz(6.938 km/s) (2-bit)+VLBI	8	<input type="radio"/>
---------	---------------	---------------	------------------	---	---	-----------------------

Feedback

Validation Validation History Log

Q ↑ ↓ ?

Overview

Contextual Help

- Please ensure you and your co-Is are registered with the [ALMA Science Portal](#)
- Create a new proposal by either:

Phase I: Science Proposal



Sensitivity



Requested RMS over is

For a peak flux density of , the S/N is

Achieved RMS over the total bandwidth is

For a continuum flux density of , the achieved S/N is

For a peak line flux of , the achieved S/N over 1/3 of the source line width (/ 3 =) is

Line width / bandwidth used for sensitivity (/) =

Justify your requested RMS and resulting S/N for the spectral line and/or continuum observations.

For line observations also justify the bandwidth used for the sensitivity calculation.

Spectral dynamic range

Spectral dynamic range is related with the bandpass accuracy.
It is 1000 (B3-B6), 400 (B7), 250 (B8), 170 (B9), and 150 (B10).